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**Response of Humidity and Clouds to
Tropical Deep Convection**

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29 **Abstract**

30

31 Currently available satellite data can be used to track the response of clouds and humidity
32 to intense precipitation events. A compositing technique centered in space and time on locations
33 experiencing high rain rates is used to detail the characteristic evolution of several quantities
34 measured from a suite of satellite instruments. Intense precipitation events in the convective
35 Tropics are preceded by an increase in low-level humidity. Optically thick cold clouds
36 accompany the precipitation burst, which is followed by the development of spreading upper
37 level anvil clouds and an increase in upper tropospheric humidity over a broader region than that
38 occupied by the precipitation anomalies. The temporal separation between the convective event
39 and the development of anvil clouds is about 3 hours. The humidity increase and associated
40 decrease in clear-sky longwave emission persist for many hours after the convective event, even
41 though the domain-averaged column integrated water vapor remains essentially constant over the
42 composite period. Large-scale vertical motions from reanalysis show a coherent evolution
43 associated with precipitation events identified in an independent dataset. Upward vertical
44 motion anomalies associated with precipitation events begin with stronger upward motion
45 anomalies in the lower troposphere, which then evolve toward stronger upward motion
46 anomalies in the upper troposphere, in conjunction with the development of anvil clouds.
47 Greater upper tropospheric moistening and cloudiness are associated with larger scale and better
48 organized convective systems, but even weaker more isolated systems produce sustained upper
49 level humidity and clear-sky OLR anomalies.

50 **1. Introduction**

51

52 Deep convective systems are prominent features of the tropical atmosphere that have
53 important roles at a spectrum of space and time scales from local diurnal cycles to the planetary-
54 scale Hadley and Walker circulations. Upward motion occurring in the ascending branches of
55 the Hadley and Walker circulations is realized not in the form of large-scale continuous ascent
56 but in the form of a large number of relatively small-scale discrete convective plumes occurring
57 in an otherwise subsiding environment (Yanai et al. 1973). Because it is the ensemble of
58 transient deep convective processes that results in the measured mean ascent in the Tropics,
59 understanding the large spatial and long temporal scale tropical circulation requires
60 understanding deep convective processes.

61 In addition to its dynamical importance for the tropical circulation, deep convection is the
62 primary source of high clouds and free-tropospheric water vapor, which strongly impact the
63 radiation budget of the planet. Stratiform anvil clouds that tend to have a negative net radiative
64 forcing and thin cirrus clouds that tend to have a positive net radiative forcing (Hartmann, et al.
65 2001), spread outward from deep cumulonimbus clouds and are the most prominent cloud in the
66 convective regions of the Tropics. Moist boundary layer air is transported by deep convection to
67 the dry mid- and upper-troposphere, where even slight humidity increases can strongly reduce
68 the radiation emitted to space (Shine and Sinha 1991; Udelhofen and Hartmann 1995; Spencer
69 and Braswell 1996; Allan et al. 1999; Held and Soden 2000; Colman 2001). Because the ability
70 of the Tropics to retain heat determines how much energy is available for export to the rest of the
71 Earth (Pierrehumbert 1995), deep convection strongly affects not only the local radiation budget,
72 but also the energy budget of the entire planet. Indeed, Soden and Fu (1995) find that the

73 frequency of deep convection is strongly correlated with changes in upper tropospheric humidity,
74 and that these variations are responsible for about half of the regional greenhouse effect
75 variations. Several studies (Salathe and Hartmann 1997; Pierrehumbert and Roca 1998; Dessler
76 and Sherwood 2000) demonstrate using trajectory analyses that transport of moisture away from
77 deep convection is the primary mechanism by which the free troposphere is moistened, and that
78 the sublimation of detrained condensate is not a significant moisture source. Inamdar and
79 Ramanathan (1994) provide evidence for a super-greenhouse effect, in which the moistening
80 effect of deep convection reduces the ability of the surface-atmosphere system to radiate away
81 “excess” energy in regions of high SST.

82 The ability to properly simulate deep convection and the corresponding cloud and
83 humidity fields remains a challenge to global climate models. Understanding how clouds and
84 humidity will change under greenhouse warming is largely dependent on understanding the
85 convective processes which determine their distribution, and this requires detailed measurements
86 of tropical convective systems. In this study we use profiles of humidity retrieved by the
87 Atmospheric Infrared Sounder (AIRS) and cloud properties from the Moderate Resolution
88 Imaging Spectroradiometer (MODIS) onboard NASA’s Aqua satellite to investigate the
89 evolution of high clouds, humidity, and clear sky outgoing longwave radiation (OLR_{CS})
90 associated with deep convection. The spatial and temporal humidity and cloud distributions are
91 investigated by compositing about locations of deep convection, identified by intense rain rates
92 (RRs) from the Tropical Rainfall Measuring Mission (TRMM) Multisatellite Precipitation
93 Analysis (TMPA). In addition, we assess the sensitivity of the upper tropospheric moisture and
94 cloud evolution to the scale of the convection.

95 It is well known that while some tropical deep convection “pops up” in regions distant
96 from synoptic support, the vast majority of deep convection is organized into convective
97 complexes, which have been extensively documented (e.g., Houze and Betts 1981; Gamache and
98 Houze 1983; Houze 1982; Chen and Houze 1997; Nuret and Chong 1998; Sherwood and
99 Wahrlich 1999). These studies have largely relied on ground-based and geostationary satellite
100 observations of individual convective systems during field campaigns. The present study aims to
101 further detail the evolution of the moisture and cloud fields in the vicinity of tropical deep
102 convective events throughout the tropical Pacific. We take a statistical approach as a means of
103 showing climatological characteristics of convective systems rather than highly-detailed
104 observations of individual systems.

105 The AIRS instrument is uniquely suited for measuring the humidity distribution: It
106 provides radiosonde-quality humidity and temperature retrievals at high vertical resolution
107 throughout the troposphere, unlike geostationary water vapor channels which sense vapor in a
108 thick upper tropospheric layer between about 500 and 200 hPa; its global coverage allows
109 observations from a broader range of locations than those within the radiosonde network or
110 within field campaign domains; the duration of observations allow for a larger and more diverse
111 set of samples than those of field campaigns; and the cloud-clearing techniques employed in the
112 AIRS retrieval algorithm (Susskind et al. 2003) allow for sampling with greater confidence in
113 cloudy regions where infrared retrievals from other instruments become contaminated.

114

115 **2. Data and Quality Control**

116

117 The Aqua satellite was launched on May 4, 2002 as part of NASA's Earth Observing
118 System into a sun-synchronous near-polar orbit at 705.3 km above the earth's surface. The
119 satellite ascends (descends) across the equator at approximately 1:30 PM (AM) local time
120 providing global coverage approximately every two days, depending on the sensor. We use
121 retrievals over the tropical Pacific ITCZ region (5°N to 15°N, 120°E to 260°E) between January
122 2003 and December 2005 from three sensors onboard Aqua: AIRS, MODIS, and AMSR-E. We
123 also use a precipitation dataset (TMPA) that includes measurements from a suite of polar-
124 orbiting and geostationary satellites. Finally we make use of horizontal and vertical winds from
125 the NCEP/NCAR Reanalysis.

126

127 *a. AIRS*

128

129 The descriptions of the AIRS instruments are provided in detail in Aumann et al. (2003).
130 AIRS is actually a suite of instruments, a hyperspectral infrared instrument (AIRS) with 2378
131 channels between 3.7 and 15.4 μm and a 13.5 km footprint at nadir, the Advanced Microwave
132 Sounding Unit (AMSU-A) with 15 microwave channels between 23 and 90 GHz and a 40.5 km
133 footprint at nadir, and a visible / near IR sensor with four channels between 0.40 and 0.94 μm
134 and a 2.3 km footprint at nadir (Aumann et al. 2003). Within each AMSU-A field of view (FOV)
135 is an array of three-by-three AIRS FOVs, and within each AIRS FOV is an array of eight-by-
136 nine visible / near IR FOVs. The visible / near IR channels are primarily used to flag the
137 presence of low clouds (Gautier et al. 2003). Because information from all sensors is used
138 simultaneously, the geophysical (level 2) retrievals are reported at the nominal resolution of 40.5
139 km at nadir.

140 The AIRS retrieval algorithm makes use of a cloud-clearing technique described in detail
141 in Susskind et al. (2003). Briefly, the technique takes advantage of the fact that while IR
142 retrievals are strongly affected by the presence of clouds, microwave temperature retrievals are
143 largely insensitive to the presence of clouds. Thus, any horizontal inhomogeneity in the
144 radiances observed by the three-by-three array of IR footprints within the microwave footprint is
145 largely caused by varying amounts of clouds within each IR FOV. One important assumption in
146 this approach is that only the relative amount and not the radiative properties of a given cloud
147 type vary between the IR FOVs. A second assumption is that the geophysical properties that are
148 retrieved in the clear portions of the FOVs are identical in each IR FOV. Making use of this
149 technique requires no *a priori* assumptions about or modeling of the cloud properties (height,
150 emissivity, etc.), nor does it limit the IR retrievals to rare clear-sky scenes.

151 We use retrievals of clear sky outgoing longwave radiation (OLR_{CS}) and profiles of water
152 vapor mixing ratio (w), saturation water vapor mixing ratio (w_s), and temperature (T) from the
153 AIRS version 4, level 2 (swath) product. The profile of w is a layer quantity, representing the
154 mean mixing ratio between the standard pressure levels, while profiles of w_s and T are level
155 quantities, representing the value at the pressure level. For temperatures higher than 273.15 K,
156 w_s is calculated with respect to water, otherwise it is calculated with respect to ice using the
157 Buck (1981) formulation. We calculate the layer geometric mean w_s such that the calculated RH
158 is a layer-mean RH between pressure levels.

159 Atmospheric T retrievals have been compared with ECMWF analyses and dedicated
160 radiosondes and are found to be accurate to 1°C for every 1 km thick layer (Fetzer et al. 2005).
161 Similarly, w profiles have been compiled in 2 km layers and compared with dedicated
162 radiosondes. At the Tropical West Pacific (TWP) site, Fetzer et al. (2005) find the following

163 biases: -10.5% (1013-700 hPa), -0.7% (700-500 hPa), -2.3% (500-300 hPa), -16.1% (300-200
164 hPa), and 15.1% (200-150 hPa).

165 We apply all appropriate quality assurance flags to the AIRS dataset. AIRS products
166 have a hierarchy of quality flags (Susskind et al. 2006), based on whether all steps in the retrieval
167 algorithm are performed satisfactorily. Generally, precipitation and/or cloud fractions exceeding
168 80% in the AIRS fields of view cause the retrieval to fail. We perform additional quality control
169 by removing all retrievals in which the RH with respect to liquid water (calculated using the
170 Buck (1981) w_s formulation) exceeds 100% at any layer within the retrieval. This removes
171 spuriously high RHs while allowing for humidities that are supersaturated with respect to ice, as
172 is frequently observed in the upper troposphere (c.f, Gettelman et al. 2006). Approximately 7%
173 of the AIRS retrievals were removed due to supersaturation with respect to liquid water. Finally,
174 the AIRS data are re-gridded to 1° horizontal resolution, keeping a record of the number of
175 observations contained within each 1° box.

176

177 *b. MODIS*

178

179 The Moderate Resolution Imaging Spectroradiometer (MODIS) is a whiskbroom
180 scanning radiometer with 36 channels between 0.415 and 14.235 μm . The retrieval of cloud
181 properties for the level 2 products is described in detail in Platnick et al. (2003). As many as 20
182 channels are used in the cloud detection algorithm to create a cloud mask at 1 km resolution,
183 which is essentially a measure of the confidence that the FOV is clear. CO_2 slicing within the
184 broad 15 μm absorption band is then used to infer cloud top pressure and the effective cloud
185 amount on a 5x5 pixel (5 km at nadir) scene, assuming at least 4 of the 25 pixels are flagged as

186 probably cloudy or cloudy. Temperature profiles from the GDAS gridded meteorological
187 product (Derber et al. 1991) are then used to derive cloud top temperature (CTT). Optical
188 thickness (τ) retrievals, which make use of the 0.65, 0.86, and 1.2 μm bands in addition to
189 inferences about cloud phase, are only provided for the daytime observations (ascending orbits in
190 the Tropics). We remove cloud retrievals that are affected by sun glint, over land, and where the
191 cloud mask is undetermined.

192 Using the cloud fraction, CTT, and τ data at 5 km (nadir) resolution, we calculate
193 histograms of average high (CTT < 245 K) cloud fraction in three bins of τ at 1° horizontal
194 resolution as in Kubar et al. (2007). To ensure sufficient sampling for each histogram, we
195 require that at least 65 MODIS pixels are contained within each 1° grid space (consistent with
196 Kubar et al. (2007)). For each 1° grid space, we bin the cloud fractions based on τ ranges that
197 are chosen to distinguish high clouds with negative radiative forcing ($\tau \geq 4$) from high thin
198 clouds with positive radiative forcing ($\tau < 4$). The clouds with negative forcing are further
199 separated into anvil ($4 \geq \tau < 32$) and thick ($\tau \geq 32$) clouds. The assumption here, which will be
200 supported by the results, is that the intermediate optical depth cold cloud corresponds to
201 extended upper level anvil cloud associated with convection, while the thick cold cloud is more
202 closely associated with heavy precipitation.

203

204 *c. AMSR-E*

205

206 AMSR-E is a conically-scanning passive microwave radiometer sensing polarized
207 radiation at six frequencies between 6.9 and 89 GHz. The AMSR-E instrument and retrieval
208 algorithms are explained in detail in Japan Aerospace Exploration Agency (2005). We use

209 retrievals of column-integrated water vapor (WVP) over the ocean from the AMSR-E version 5
210 ocean product. WVP is retrieved using the 18.7, 23.8, and 36.5 GHz brightness temperatures.
211 First, cloud liquid water index (CWI) is derived from brightness temperature, atmospheric
212 transmittance, and vertical mean atmospheric temperature at 18.7 and 36.5 GHz. Then, WVP is
213 calculated from the CWI, atmospheric transmittance at 18.7 and 23.8 GHz, and a set of
214 regression coefficients which minimize differences between WVP and the precipitable water
215 from radiosondes. These data are mapped onto a 0.25° resolution grid by Remote Sensing
216 Systems and WVP retrievals are discarded in the presence of heavy rain. We re-grid these data
217 to 1° resolution, keeping a record of the number of observations contained within each 1° box..
218

219 *d. TMPA*

220

221 The Tropical Rainfall Measuring Mission (TRMM) Multisatellite Precipitation Analysis
222 (TMPA) provides precipitation rates at three hour resolution between 50°S and 50°N from a
223 suite of polar orbiting and geostationary satellites. Huffman et al. (2007) provide a detailed
224 description of the dataset. The data are collected from the Microwave Imager (TMI) on TRMM,
225 Special Sensor Microwave Imager (SSM/I) on Defense Meteorological Satellite Program
226 (DMSP) satellites, AMSR-E on Aqua, and the Advanced Microwave Sounding Unit-B (AMSU-
227 B) on the National Oceanic and Atmospheric Administration satellite series. These polar
228 orbiting passive microwave sensors cover about 80% of the earth's surface between 50°S and
229 50°N over three-hour periods. Precipitation data for the remaining 20% is inferred from window
230 channel IR data collected by geostationary satellites. Three additional data sources are the
231 TRMM Combined Instrument (TCI) estimate, the Global Precipitation Climatology Project

232 (GPCP) monthly rain gauge analysis and the Climate Assessment and Monitoring System
233 (CAMS) monthly rain gauge analyses.

234 Briefly, the TMPA RR estimates are made using the following procedure: the microwave
235 estimates are calibrated and combined, IR estimates are created using the calibrated microwave
236 precipitation, the microwave and IR estimates are combined, and finally the rain gauge data are
237 incorporated as a means of scaling the retrieved precipitation estimates. The data are reported at
238 the nominal 3-hourly observation times (0000, 0300, ..., 2100 UTC), averaging polar orbiting
239 data that are ± 90 minutes from these times. We make use of the version 6 3B42 product, which
240 we re-grid from 0.25° to 1° horizontal resolution.

241

242 *e. NCEP/NCAR Reanalysis*

243

244 Finally, 6-hourly horizontal and vertical winds from the NCEP/NCAR Reanalysis
245 (Kalnay et al. 1996) over the same period are used. These data are linearly interpolated from 2.5°
246 to 1° horizontal resolution and to 3-hourly resolution.

247

248 **3. Methodology**

249

250 We are interested in the moistening effects of tropical deep convection from a statistical
251 and climatological perspective rather than from a case-by-case perspective; thus we composite
252 over many thousands of deep convective events in the Pacific ITCZ over the three-year period
253 Jan 2003 – Dec 2005. We locate deep convection by seeking RRs that exceed the 90th percentile,
254 which – after gridding the data to 1° resolution – is 1.6 mm hr^{-1} for this time period in the

255 Tropical Pacific (-25°N to 25°N , 120°E to 260°E). The cumulative sum of precipitation
256 (expressed as a fraction of the total precipitation) is plotted as a function of instantaneous RR in
257 Figure 1. During the three-year period, 57% of total rainfall in the Tropical Pacific fell in RR
258 events that exceeded the 90th percentile (vertical line in Figure 1). Thus by choosing the 90th
259 percentile of RR as a threshold, we concentrate on very intense convection which makes up a
260 substantial portion of the total accumulated precipitation while still retaining a large sample size.

261 We composite the meteorological fields in 11×11 grids of 1° grid spaces surrounding
262 each RR grid space exceeding this threshold value and for 24 hours before and 24 hours after the
263 time of deep convection. Where multiple adjacent grid spaces exceed the RR threshold, the
264 composites centered at each rainy grid space are averaged into one distinct realization so as to
265 maintain the most conservative estimate of the number of independent samples.

266 Even though AQUA passes over a given location only once every 12 hours in the
267 Tropics, the precipitation data are provided every three hours everywhere, thereby resulting in
268 sampling of geophysical quantities at a wide range of time offsets from the deep convection
269 events. We choose composite temporal increments of 3 hours by averaging all retrievals that fall
270 within ± 90 minutes of the three hour increments. For example, any Aqua overpass that occurred
271 between 4.5 and 1.5 hours prior to a high RR observation is placed at hour -3.

272 The use of a reference frame centered on intense RRs that is fixed in time over the 48
273 hour period surrounding the convective events is chosen to study the effects of convection on the
274 environment in an Eulerian sense: generally, convective systems pass through the domain from
275 east to west within the tropical easterlies. Unlike the studies of Soden (1998, 2004) and
276 Sherwood and Wahrlich (1999), we do not attempt to track convective systems but rather
277 concentrate on the effects of convection on the environment through which the convection

278 passes. Thus we sample a spectrum of convection, from events that pop up stochastically to
279 convective systems that propagate in time.

280 Composite anomalies are computed by subtracting from each time lag the temporal mean
281 of the composite-mean pattern over the 48-hour period. We refer to the fractional coverage of
282 the anomaly as the fraction of the $11^\circ \times 11^\circ$ domain that is occupied by grid spaces that are greater
283 than or equal to e^{-1} of the maximum anomaly observed over the entire composite period. In the
284 case of OLR_{CS} and ω_{UT} , the region is defined relative to the maximum (absolute) negative
285 anomaly. We define the magnitude of the anomaly as the maximum positive anomaly at each
286 time lag, except for OLR_{CS} and ω_{UT} , in which case the magnitude is the maximum (absolute)
287 negative anomaly at each time lag. We choose these metrics rather than simply taking spatial
288 averages across the domain because they retain scale and magnitude information, allowing for
289 separation of small-scale but very anomalous features from large scale weakly anomalous
290 features which might appear identical in spatial averages. For anomalies that are three-
291 dimensional (e.g. RH anomalies), we first average the composite anomalies over the appropriate
292 levels (e.g., between 500 and 200 hPa for upper tropospheric anomalies), then calculate the
293 fractional coverage such that we are only reporting the horizontal extent of the anomalies. As
294 will be shown below, many anomalies propagate out of the domain, so it is not always possible
295 to say with certainty when they have spread to their largest extent or have reached their
296 maximum amplitudes. In these cases we simply put a lower bound on the time at which these
297 occur.

298

299 **4. Results**

300

301 Results are shown for convection observed within the tropical Pacific ITCZ, defined as
302 5°-15°N, 120°E-260°E. Aside from spatial pattern differences due to underlying SSTs, the
303 results shown here are robust throughout the Tropics, assuming a climatologically convective
304 region is chosen for analysis. Composites generated using only observations from ascending
305 AQUA passes (1:30 pm local time) show no differences from those generated using only
306 descending AQUA passes (1:30 am local time), nor do seasonal differences affect the results.

307

308 *a. Composites about all RRs exceeding the 90th percentile*

309

310 Plan views of eight composite-mean variables as a function of time relative to the high
311 precipitation event are shown in Figure 2. As required by the compositing technique, the central
312 grid space in the plan view at hour 0 contains the maximum RR and the composite-averaged RR
313 is significantly lower in the hours preceding and proceeding this time. High RRs are confined to
314 the central portion of the domain and are observed for only a few hours before and after peak
315 RR.

316 Composite-mean WVP is highest near the center of the domain, with larger meridional
317 than zonal gradients (i.e., the moist region is elongated in the east-west direction). The zonal
318 elongation of mean WVP reflects the underlying SSTs in the analysis region, which extends
319 from the northern edge of the western Pacific warm pool in the west to a narrow warm tongue of
320 relatively high SSTs in the east (c.f., Figure 1 of Kubar et al. 2007).

321 OLR_{CS} is correspondingly lower where the composite mean WVP is higher (i.e., near the
322 center of the domain), as the bulk of the radiation escaping to space originates from a higher
323 altitude and therefore lower temperature where it is more moist. A subtle but important

324 difference between the OLR_{CS} and WVP will become apparent in the anomaly patterns discussed
325 below.

326 Composite-mean high thick cloud fraction looks nearly identical to the precipitation,
327 maximizing at hour 0 and falling off rapidly away from hour 0, in addition to remaining confined
328 to the center of the domain. This strengthens our confidence that the optical depth threshold
329 chosen to define thick high clouds ($\tau \geq 32$) is indeed capturing the deep convective cores.

330 Whereas the high thick clouds only cover the heavily precipitating region and peak in
331 phase with the precipitation, the anvil cloud fractions cover a larger area of the domain and peak
332 at hour +3. The coverage of anvil clouds tends to be elongated in the zonal direction and
333 confined in the meridional direction. This spatial pattern is mainly due to the inclusion of
334 eastern Pacific convection, which occupies a broad range of longitudes but is confined to a
335 relatively narrow band of latitudes over the highest SSTs (not shown). The spatial patterns and
336 evolution of anvil cloud fraction lends credence to our choice of optical depth ranges defining
337 anvil clouds: We expect greater spatial coverage of anvils relative to deep convective cores, as
338 well as a time lag between peak convection and their maximum extent as they spread outward
339 from deep convection. This result is also consistent with previous studies (Soden 2000, 2004;
340 Tian et al. 2004) that show a time lag between the peak convective cloud fraction and the peak
341 anvil cloud fraction.

342 High thin cloud is well distributed in space and time outside of where the high thick and
343 anvil clouds are predominant. Largest high thin cloud fractions tend to be near the edge of the
344 anvil clouds. Note that the local minimum in high thin cloud fraction near the center of the
345 domain between hour -6 and hour +6 is simply an artifact of there being only a finite area for all

346 three cloud types to occupy: The high thin cloud fraction must be lower to accommodate the
347 large thick and anvil cloud fractions present near the deepest convection.

348 Plan-views of the anomalies calculated by subtracting from each time lag the temporal
349 mean of the composite-mean pattern over the 48-hour period are shown in Figure 3. In addition,
350 fractional coverages and anomaly magnitudes of these fields are plotted in Figure 4. Fractional
351 coverage of the anomaly region refers to the fraction of grid spaces in the $11^\circ \times 11^\circ$ domain at
352 each time lag that are greater than or equal to e^{-1} of the maximum anomaly observed over the
353 entire composite period. Anomaly magnitude refers to the maximum anomaly at each time lag.
354 In the cases of ω_{UT} and OLR_{CS} , the quantities are calculated with respect to the maximum
355 (absolute) negative anomaly. The errorbars represent the 95% confidence limits averaged in
356 space over the domain.

357 RR anomalies look identical to composite-mean RRs because we have isolated very
358 extreme RR events. Anomalously high RRs are confined to the central portion of the domain
359 and peak at approximately 2 mm hr^{-1} at hour 0.

360 Peak WVP anomalies (Figure 4d) coincide with the peak RR (Figure 4b), which is
361 consistent with the instantaneous correlations between RR and WVP shown in Bretherton et al.
362 (2004). Anomalies of WVP clearly show a westward propagation of a moist signature. Thus it
363 is likely that the majority of the convection observed in these composites does not occur
364 spontaneously but rather is organized into convective systems that propagate from east to west in
365 time, as documented by Reed and Recker (1971). The WVP anomaly size remains essentially
366 constant as the system crosses the domain (Figure 4c).

367 As alluded to above, OLR_{CS} anomalies behave quite differently than WVP anomalies.
368 Rather than peaking at hour 0, they reach maximum amplitude following the convection and are

369 anomalously lower in the 24 hours following convection than in the 24 hours prior to convection
370 (Figure 4d). As will be shown below, the region of reduced OLR_{CS} corresponds to a sustained
371 moist anomaly in the upper troposphere following deep convection.

372 Whereas anomalously high thick cloud fractions only cover the heavily precipitating
373 region and peak in phase with the precipitation, the anvil cloud fraction anomalies cover a larger
374 area of the domain, peak at hour +3, and remain extensive for several hours following convection
375 (Figures 4a-b).

376 The region occupied by anomalously large fractions of high thin cloud (not shown) is
377 larger at nearly all times than either high thick or anvil cloud, though the anomalies themselves
378 are quite small ($\sim 2\%$). This difference would be more dramatic if the fractional coverages were
379 calculated over a larger domain that captures the broad horizontal extent of high thin cirrus
380 clouds.

381 The asymmetry in OLR_{CS} anomalies in the presence of symmetric WVP anomalies can
382 be explained by the pattern of RH anomalies, shown in Figure 5. Whereas the low-level (below
383 700 hPa) RH anomalies peak just prior to the maximum RR (hour 0), the upper troposphere (600
384 hPa - 200 hPa) is anomalously moist between 3 and 24 hours following the convection in both
385 regions. The RH anomalies in the upper troposphere are much larger than those in the lower
386 troposphere, peaking around 10% at hour +9 (Figure 4b). Anomalies above 200 hPa are
387 negligible.

388 These results are consistent with the lag correlations between radar-derived rain rate and
389 lower and upper tropospheric RH from radiosondes shown in Sobel et al. (2004): Lower
390 tropospheric RH (RH_{LT}) is highly correlated with rain rate at lags of -6 hours and 0 hours, while

391 peak correlation between upper tropospheric RH (RH_{UT}) and rain rate occurs at a lag of +6
392 hours. The peak correlation is stronger for RH_{UT} than for RH_{LT} .

393 The fractional coverage of the RH_{UT} anomaly region increases until at least hour +15
394 (Figure 4a). Beyond hour +18, the RR anomaly has moved out of the domain, but the residual
395 RH_{UT} anomaly near the western edge of the domain is still greater than that at the center of the
396 domain at hour 0.

397 It is important to note that fractional anomalies from the composite mean mixing ratio
398 (not shown) exhibit the same patterns and magnitudes as the RH anomalies, indicating that the
399 positive RH anomalies at upper levels are not due to temperature changes. The temperature
400 perturbations (not shown) are negligible (on the order of 0.1 K throughout the troposphere) and
401 changes in RH are almost entirely due to mixing ratio changes. The absence of substantial
402 temperature perturbations over the course of several individual convective events and over the
403 course of a composite easterly wave was noted by both Sobel et al. (2004) and Reed and Recker
404 (1971), respectively.

405 To elucidate the co-evolution of humidity and vertical motion, spatial averages across the
406 $11^\circ \times 11^\circ$ domain are calculated of omega and of anomalies of omega and RH (Figure 6).
407 Although upward motion is observed throughout the troposphere at all times in the composite
408 period, it shifts from bottom-heavy prior to convection to top-heavy following the convection.
409 Anomalous upward motion reaches peak amplitude in the upper troposphere at hour 0 (Figure
410 4b), but occupies the deepest region of the troposphere at hour -3 (Figure 6). Surprisingly, the
411 lower troposphere is anomalously subsiding at hour 0 (Figure 6). Anomalously top-heavy
412 vertical motion exits the domain by hour +15. Both the fractional coverage and the magnitude of

413 the upper tropospheric vertical motion (ω_{UT}) anomaly reach a broad peak around hour 0 (Figures
414 4a-b).

415 Rough calculations show that the magnitude of spatial-mean composite-anomalous ascent
416 in the mid-troposphere is consistent with the spatial mean composite RR anomalies. Simple
417 scaling arguments show that a 1 mm day^{-1} precipitation anomaly – which is on the order of
418 spatially-averaged RR anomalies observed in this study (not shown) – should correspond to a
419 mid-tropospheric ω anomaly of -10 hPa day^{-1} . Indeed, we observe spatially-averaged ω
420 anomalies that are approximately this magnitude, which indicates that the reanalysis is (roughly)
421 capturing the vertical motion associated with deep convection. Poorer agreement between these
422 quantities is observed at smaller spatial scales, as would be expected given the coarse resolution
423 of the reanalysis compared to the scale of deep cumulonimbus updrafts.

424 Further consistency is evident between the evolution of the spatially-averaged vertical
425 motion anomalies and spatially-averaged humidity anomalies shown in Figure 6c. Spatially-
426 averaged RH anomalies clearly show a moist anomaly prior to peak convection at the low levels
427 (below 700 hPa) followed by a significantly larger moist anomaly at upper levels after the peak
428 convection. Roughly, moistening occurs where there is ascent (e.g., in the lower troposphere
429 between hours -21 and -3 and in the upper troposphere between hours -12 and +12) and drying
430 occurs where there is descent (e.g., in the upper troposphere between hours -24 and -18 and
431 between hours +15 and +24). This is to be expected, given that the humidity tendency is
432 proportional to the magnitude of upward motion in the presence of a strong downward gradient
433 of absolute humidity.

434 The vertical motions provided on the $2.5^\circ \times 2.5^\circ$ reanalysis grids are unlikely to capture all
435 small-scale motions that characterize deep convection or motions driven by local cloud radiative

436 effects, and are largely model output rather than observations in the remote areas of the Tropical
437 Pacific studied here. Nonetheless, the reanalysis does a remarkably good job – at least in the
438 composite sense – of producing a vertical motion field in an environment of tropical deep
439 convection that evolves in a similar manner to that shown in observational studies (e.g., Fig. 8 of
440 Reed and Recker 1971, Fig. 6 of Gamache and Houze 1983, Fig. 8f of Nuret and Chong 1998)
441 and that is consistent with humidity and precipitation anomalies observed from space.

442

443 *b. Composites as a function of spatially-averaged RR*

444

445 We now sort the composites by the spatially-averaged RR at hour 0. While still requiring
446 that the central RR grid space at hour 0 in each composite exceeds the 90th percentile, the
447 composites are now stratified by the average RR over the entire domain. This separates large
448 systems with heavy precipitation occurring over broad regions from small systems that are more
449 confined in space. We refer to the lowest (highest) quartile of spatially-averaged RR as the low
450 (high) RR regime. For consistency, the anomaly regions are calculated with respect to the
451 maximum anomalies observed using all observations (i.e., the maxima from the unsorted results
452 discussed above rather than the maxima within each RR regime). Figures 7-9 show fractional
453 coverage and magnitude of anomalies of RR, WVP, OLR_{CS}, high thick, anvil, and thin cloud
454 fraction, ω_{UT} , and RH_{UT} for each RR regime, along with the results using all observations (i.e.,
455 the unsorted results shown in Figure 4). In general, the composite anomaly plan views (not
456 shown) have similar patterns to those generated using all observations (Figures 3 and 5), but the
457 size and amplitude of the anomalies tends to increase with spatially-averaged RR.

458 Anomalously high RRs cover the greatest fraction of the domain at hour 0 in all regimes,
459 with fractional coverage increasing with RR regime (Figure 7a). Because the scale of the
460 convective system increases with RR regime, the anomalous RR region arrives in the domain
461 earlier and exits later with increasing RR regime. The anomaly amplitudes (Figure 7b) are
462 indistinguishable among RR regimes at hour 0 because the composites are all centered on RRs
463 exceeding the 90th percentile. Thus sorting by spatially-averaged RR at hour 0 tends to separate
464 large systems (4th quartile) from small systems (1st quartile).

465 The size of the region occupied by WVP anomalies remains essentially constant at about
466 10% across the composite period in all RR regimes (Figure 7c). Though the WVP anomaly
467 peaks at hour 0 in all regimes, the anomaly amplitude is larger in the lower RR regimes (Figure
468 7d). This is consistent with the relationship between instantaneous RR and absolute WVP
469 derived in Bretherton et al. (2004). An implication of the relationship is that a larger WVP
470 anomaly is required in a relatively dry environment (lower RR regime) compared to a relatively
471 moist environment (higher RR regime) to produce the same RR anomaly, as is the case here.
472 Again, the symmetry of WVP anomaly amplitude about hour 0 can be contrasted with the OLR_{CS}
473 anomaly (Figure 7f).

474 The fractional coverage of the OLR_{CS} anomaly increases from just prior to convection
475 until reaching peak coverage following convection (Figure 7e). The anomalies are generally
476 greater and more expansive for the high RR regimes and reach maximum size at the same time
477 as the corresponding RH_{UT} anomalies (discussed below). Maximum (absolute) OLR_{CS}
478 anomalies between -5 and -9 W m⁻² are observed at hour +3 in all RR regimes (Figure 7f), but
479 the fractional coverage tends to remain expansive until at least hour +12 in all RR regimes.

480 The fractional coverage and anomaly magnitude of high thick cloud fraction evolves in a
481 nearly identical manner to RR anomalies, though the fractional coverage is much smaller (Figure
482 8a). Similar to the RR anomalies, high thick cloud fraction anomalies grow rapidly to a peak at
483 hour 0 and then rapidly decrease (Figure 8b). The fractional coverage is likewise stratified by
484 RR regime, with thick cloud occupying a larger portion of the domain throughout the composite
485 period in the higher RR regimes. Plan views of thick cloud fraction anomalies show that the
486 anomaly region expands zonally and becomes more persistent with RR regime (not shown).

487 The evolution of anvil cloud fraction anomalies is quite different from those of RR and
488 deep convective cloud. Anomalously high anvil cloud fractions reach their maximum extent and
489 maximum amplitude at hour +3 in all RR regimes (Figures 8c-d). After hour +3, the anomaly
490 magnitude tends to be slightly greater in the higher RR regimes. The fractional size of the
491 domain in which anvil cloud fractions are anomalously high is greater for higher RRs at all times
492 in the composite. Thus deep convection that occupies a larger area of the domain results in
493 broader anvil cloud coverage, though peak coverage consistently occurs at hour +3 across the
494 regimes.

495 Because of the spatial and temporal ubiquity of high thin clouds throughout the
496 composite period and throughout the composite domain, high thin cloud fraction anomalies are
497 small (Figure 8f) and changes in their fractional coverage (Figure 8e) largely reflect changes in
498 the coverage of high thick and anvil cloud fractions.

499 Between hour 0 and hour 6, the fractional coverages of the RH_{UT} anomalies in the four
500 regimes are virtually indistinguishable (Figure 9a). Between hour +6 and hour +15, the moist
501 anomaly in the low RR regime tends to maintain its size, covering approximately 10% of the
502 domain, while the anomalies in the other regimes continue to spread until at least hour +15.

503 Anomaly size increases with spatial-averaged RR, with the second, third, and fourth RR quartiles
504 having RH_{UT} anomalies that occupy approximately 25%, 50%, and 70% of the domain,
505 respectively, at their peak extents. The region of anomalously high RH_{UT} remains extensive for
506 the remainder of the composite period in all regimes. Clearly, deep convection is pumping
507 moisture into the upper troposphere over a larger scale in the higher RR regimes throughout the
508 composite period. In the lower RR regimes, the convection is occurring on a smaller scale and
509 so the upper tropospheric moistening is confined to the center of the domain (not shown).

510 The RH_{UT} anomaly amplitudes peak at hour +9 in the lowest three RR quartiles and at
511 hour +12 in the 4th RR quartile (Figure 9b). After hour +9, the RH_{UT} anomalies are clearly
512 stratified by RR regime, with larger anomalies in higher RR regimes. Thus both the size and
513 amplitude of moist anomalies increase with the size and amplitude of convection. It is important
514 to recall that the anomalies are calculated individually for each regime with respect to the mean
515 over their respective composite periods. Both the composite-mean RH_{UT} (not shown) and the
516 maximum RH_{UT} anomaly are greater in the higher RR regimes than the lower RR regimes. It is
517 interesting to note that RH_{UT} anomalies always peak several hours after the peak anvil extent.

518 The ω_{UT} anomalies (Figure 9c-d) are likewise stratified by the RR regime, with high RR
519 regimes characterized by largest ω_{UT} anomalies (in both extent and magnitude). ω_{UT} anomalies
520 in the low RR regime are essentially zero, indicating that the reanalysis is not capturing the
521 updrafts that characterize deep convection on scales smaller than about 250 km (the resolution of
522 the reanalysis). The spatial extent of ω_{UT} anomalies exhibits a broad peak between hour -9 and
523 hour +9 (except in the lowest quartile, where no ω_{UT} anomalies exist) and is larger for larger RR
524 regimes. In all but the lowest RR regime, the anomalous ascent covers a large portion of the
525 domain between hour -12 and hour +12. Upper tropospheric ascent peaks at hour 0 in all RR

526 regimes (except RR quartile 1). It should also be noted that all RR regimes except the 1st quartile
527 show evidence of an upward propagating ω anomaly over the composite period (not shown).

528

529 **5. Conclusions**

530

531 In addition to its importance for the tropical mean circulation, deep convection is the
532 main mechanism for vertically redistributing clouds and humidity throughout the Tropics. This
533 latter virtue is of critical importance for the energy balance of the entire planet, and serves as a
534 fundamental link between discrete mesoscale atmospheric processes and the Earth's climate.
535 Making observations of these small scale processes in great detail in a range of locations is
536 absolutely necessary for better understanding the Earth's climate and using this information to
537 model climate change. The analysis herein has illustrated a means of studying the complex
538 interactions of precipitation, clouds, humidity, radiation, and dynamics at relatively fine spatial
539 and temporal scales, and can be extended over the entire globe.

540 In this study, a compositing technique centered in space and time on locations in the
541 Pacific ITCZ experiencing intense precipitation was used to investigate the response of clouds,
542 humidity, clear-sky OLR, and vertical motion to tropical deep convection. Nearly 60% of the
543 total precipitation in the Tropical Pacific falls in events upon which these composites are based.
544 Moistening occurs at low levels prior to convection and in the upper troposphere for several
545 hours following convection. The magnitude of the RH anomaly at upper levels is much greater
546 than that at low levels, and spreads outward for at least 15 hours to cover a much larger region
547 than that occupied by the precipitation anomaly. Whereas thick convective cloud fractions peak
548 at the time of the intense precipitation event, anvil clouds spread outward from the convective

549 region in time, reaching maximum spatial extent three hours later. Although the domain-
550 averaged column integrated water vapor remains essentially constant over the composite period,
551 OLR_{CS} is significantly lower following convection, highlighting its sensitivity to the coincident
552 upper tropospheric moist anomaly. Large-scale vertical motion is upward throughout the
553 troposphere during the composite period, but anomalous ascent shifts from the lower troposphere
554 to the upper troposphere over the course of the convection. This transition from bottom-heavy to
555 top-heavy vertical motion is consistent with previous observational studies of tropical convection
556 and nicely corresponds to the pattern of moistening observed here.

557 Broader regions of anomalously high RRs tend to be associated with larger anvil cloud
558 fraction, OLR_{CS} , RH_{UT} , and ω_{UT} anomalies, as is shown by sorting the composites by the
559 spatially-averaged RR at the time of peak convection. The temporal evolution of the anomalies
560 is consistent across all spatial scales of intense convection, however. Additionally, the results
561 shown here are consistent throughout the convective Tropics for all seasons, regardless of
562 whether the observations come from the ascending or descending AQUA orbit.

563 Tropical convection is organized on a broad range of spatio-temporal scales. An
564 interesting extension of the analysis presented in this work would be to investigate the similarity
565 of the results shown here to those at larger spatial and longer temporal scales. For example, does
566 the persistence in upper tropospheric humidity anomalies following precipitation anomalies also
567 exist at larger spatial and temporal scales? If so, how does this shape our understanding of
568 Tropical deep convection and its role in climate, and what can this tell us about interactions at
569 various scales?

570

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572

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585

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678 **List of Figures**

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683

684 Figure 2. Plan views of composite-mean RR, WVP, OLR_{CS}, high thick, anvil, and thin cloud
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689

690 Figure 3. Same as in Figure 2, but for the anomalies from the composite mean. Note that the
691 anomalies are calculated with respect to the temporal mean of the composite mean pattern over
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693

694 Figure 4. (Left column) Fractional coverage and (right column) magnitude of (top row) RH_{UT},
695 thick and anvil cloud fraction, ω_{UT} , (bottom row) WVP, OLR_{CS}, and RR anomalies. Note that
696 the amplitudes of ω_{UT} and OLR_{CS} anomalies are plotted with reversed sign for the purposes of
697 comparison and that the y-axes are different on the top and bottom rows. Error bars represent the
698 95% confidence intervals.

699

700 Figure 5. Plan views of anomalies of RH (%), plotted as a function of pressure between 1000
701 hPa and 150 hPa. Each domain is $11^\circ \times 11^\circ$, centered on a RR grid space exceeding the 90th
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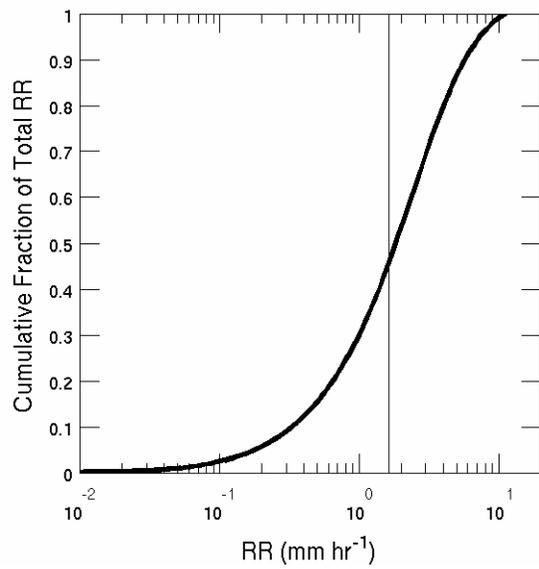
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706 Figure 6. Spatial averages over the $11^\circ \times 11^\circ$ domain of (a) composite-mean ω , (b) ω anomalies,
707 and (c) RH anomalies. The thick solid line is the zero contour.

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710 (middle row) WVP, and (bottom row) OLR_{CS} anomalies for composites in each quartile of
711 spatially-averaged RR. The thick solid line shows the fractional coverage and amplitude of these
712 quantities for all composites centered on RRs exceeding the 90th percentile. Error bars represent
713 the 95% confidence intervals. Note that the sign of the OLR_{CS} anomaly magnitude is reversed
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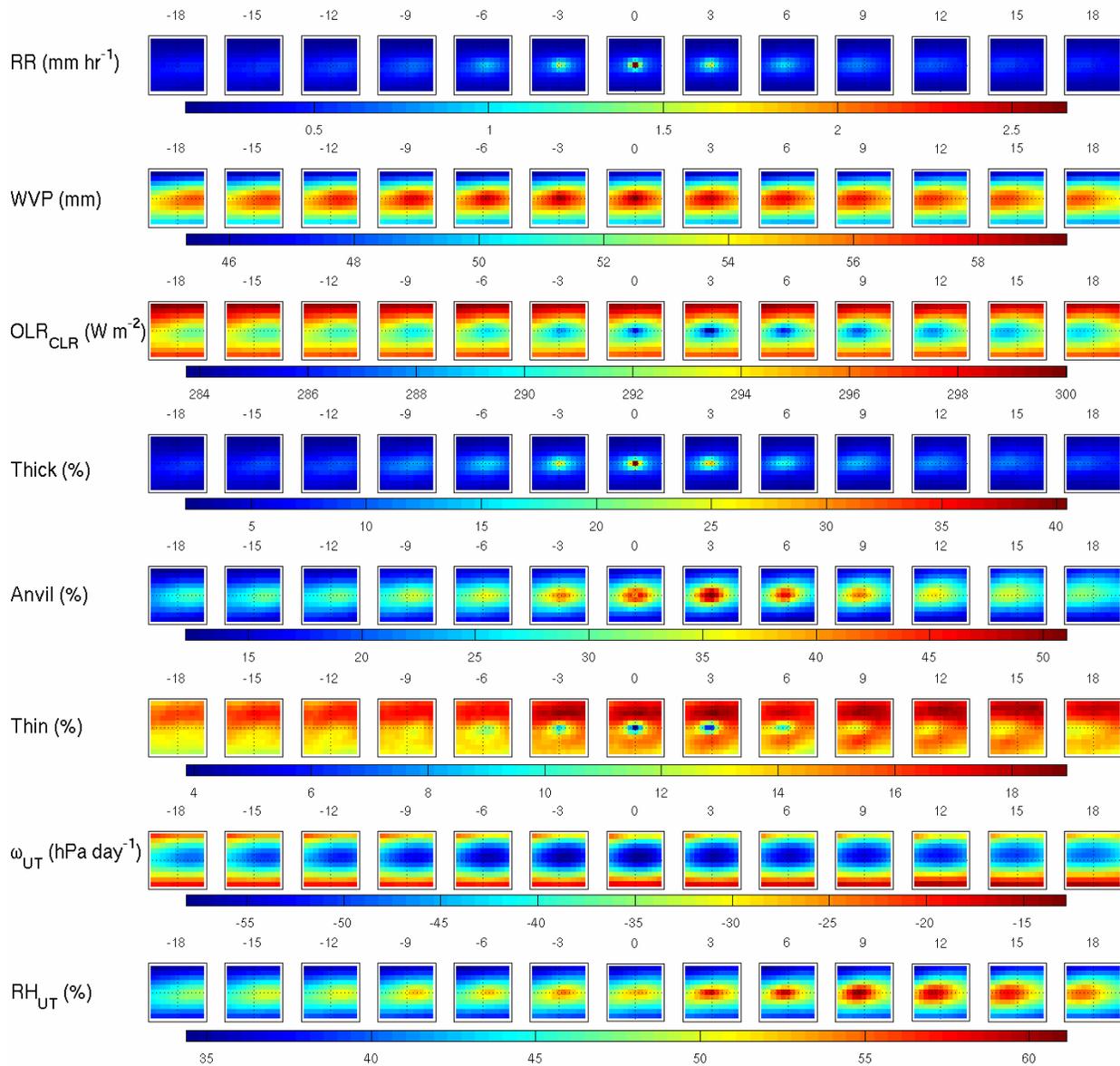
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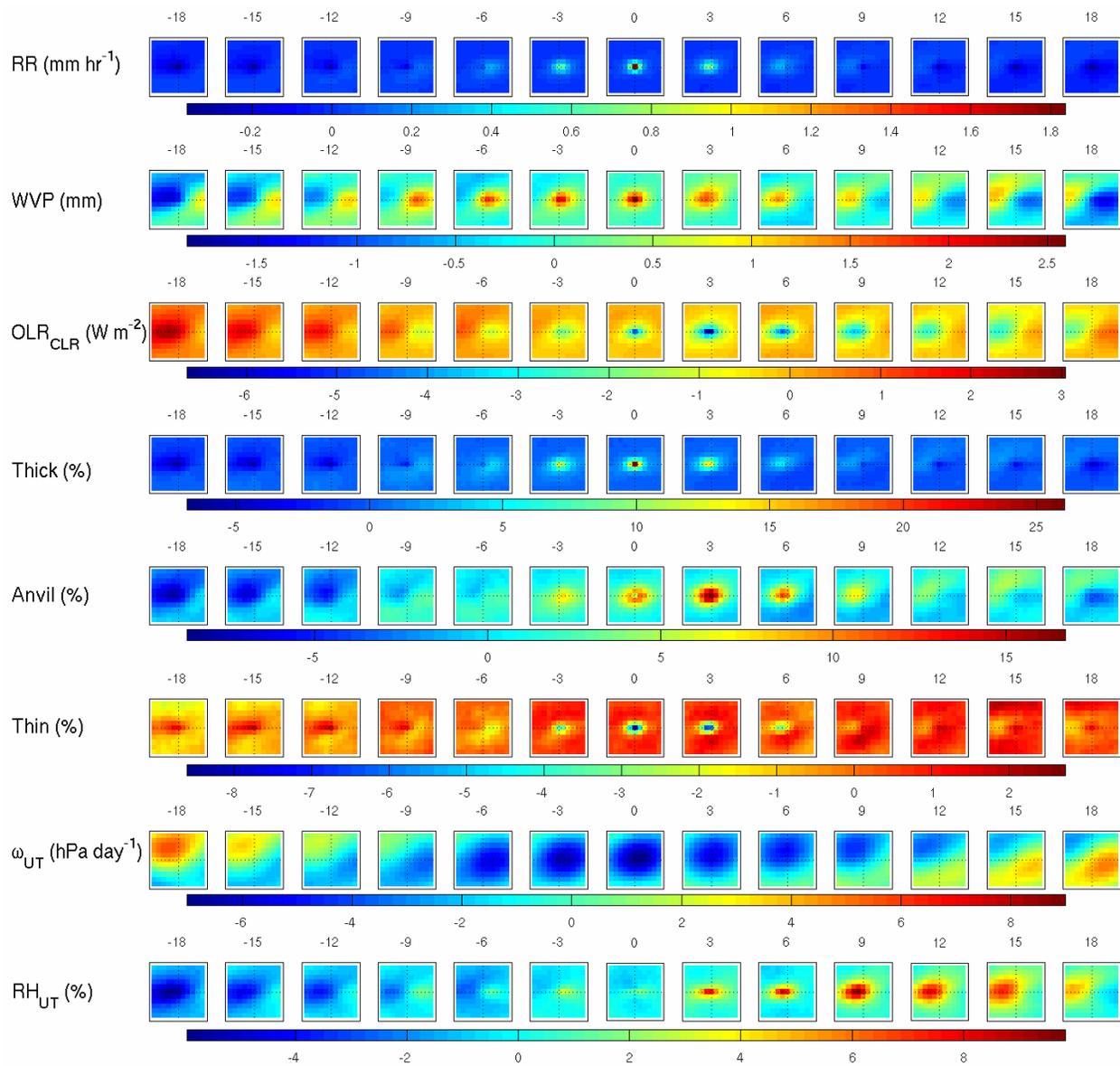
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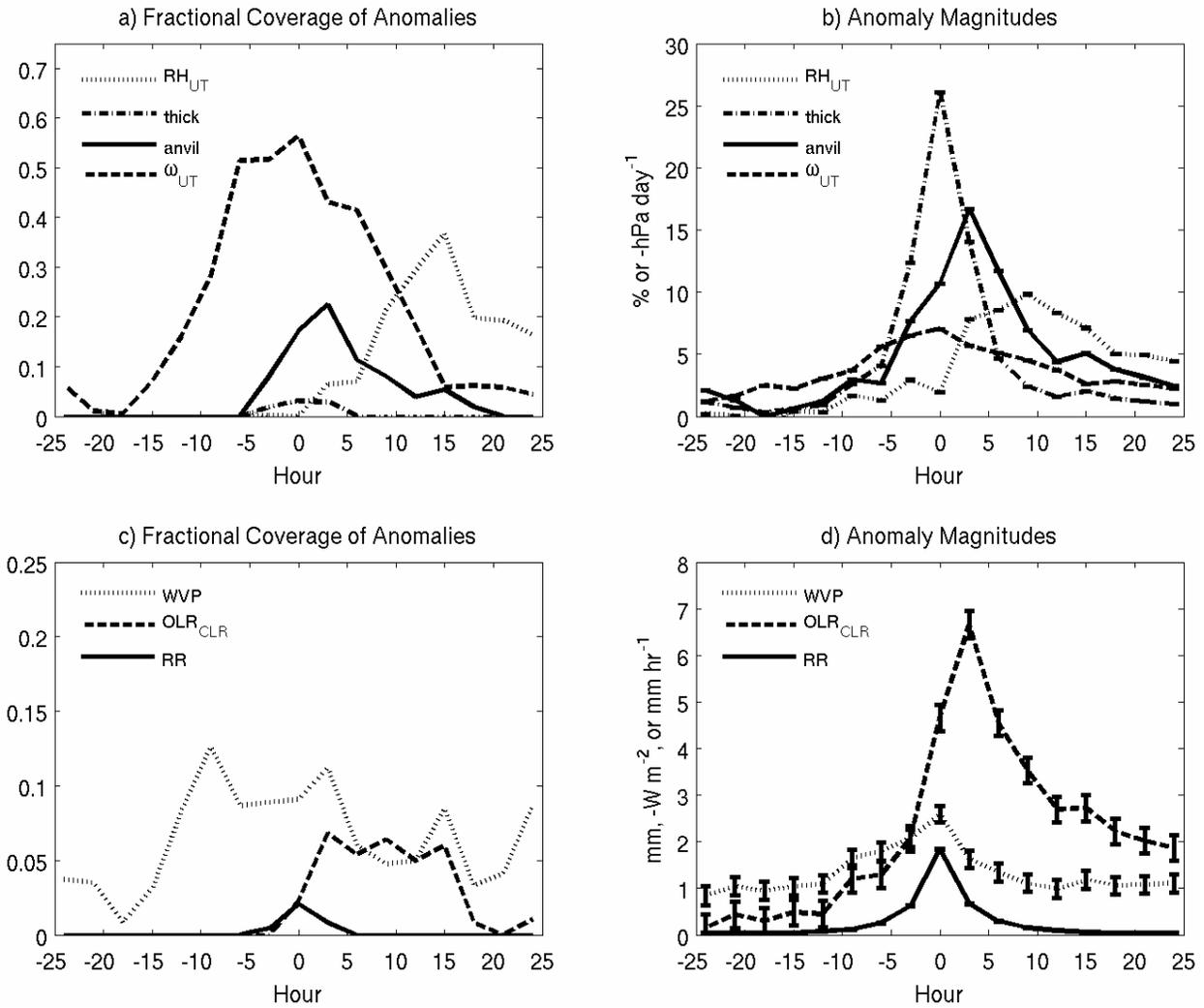
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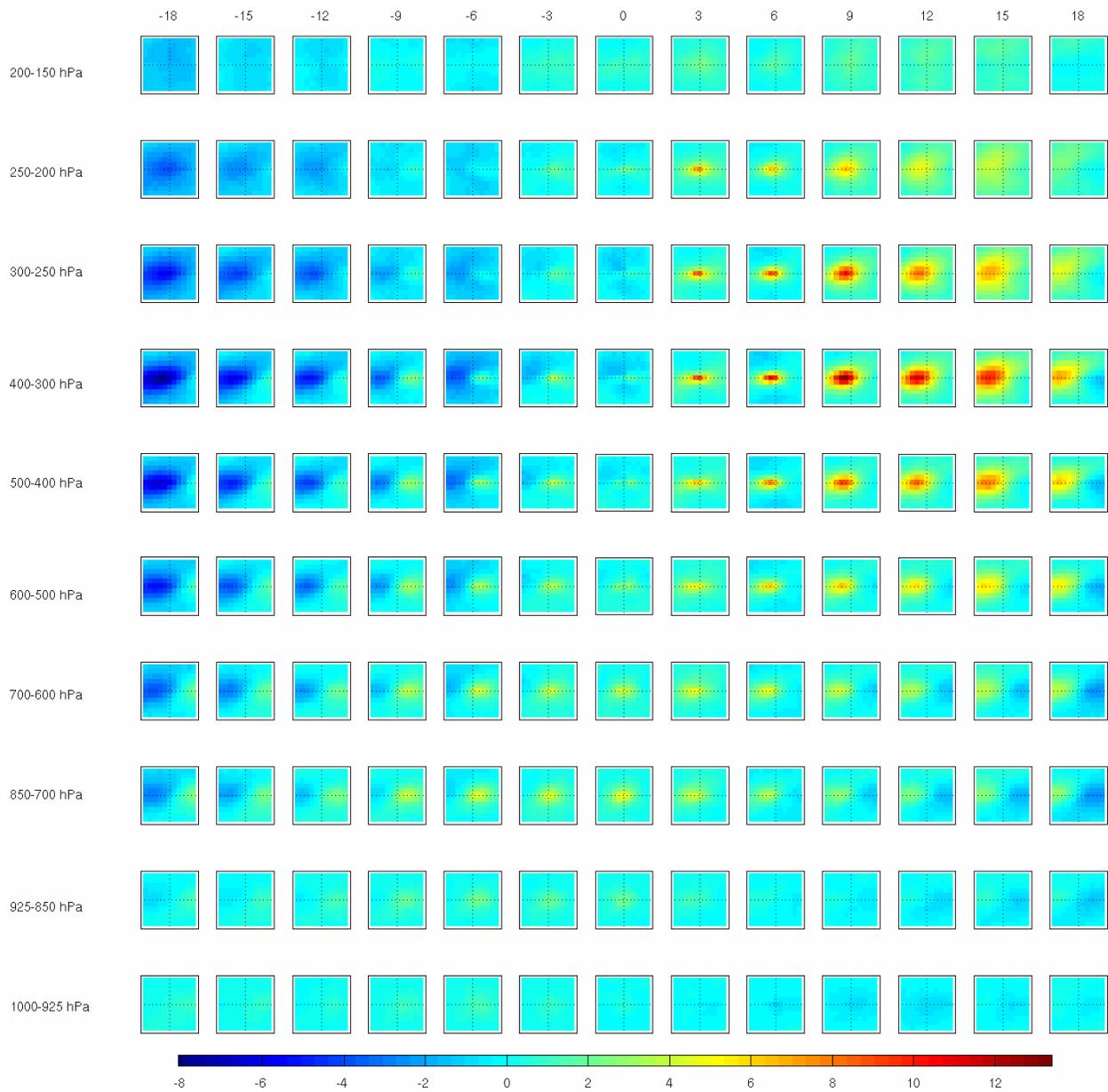
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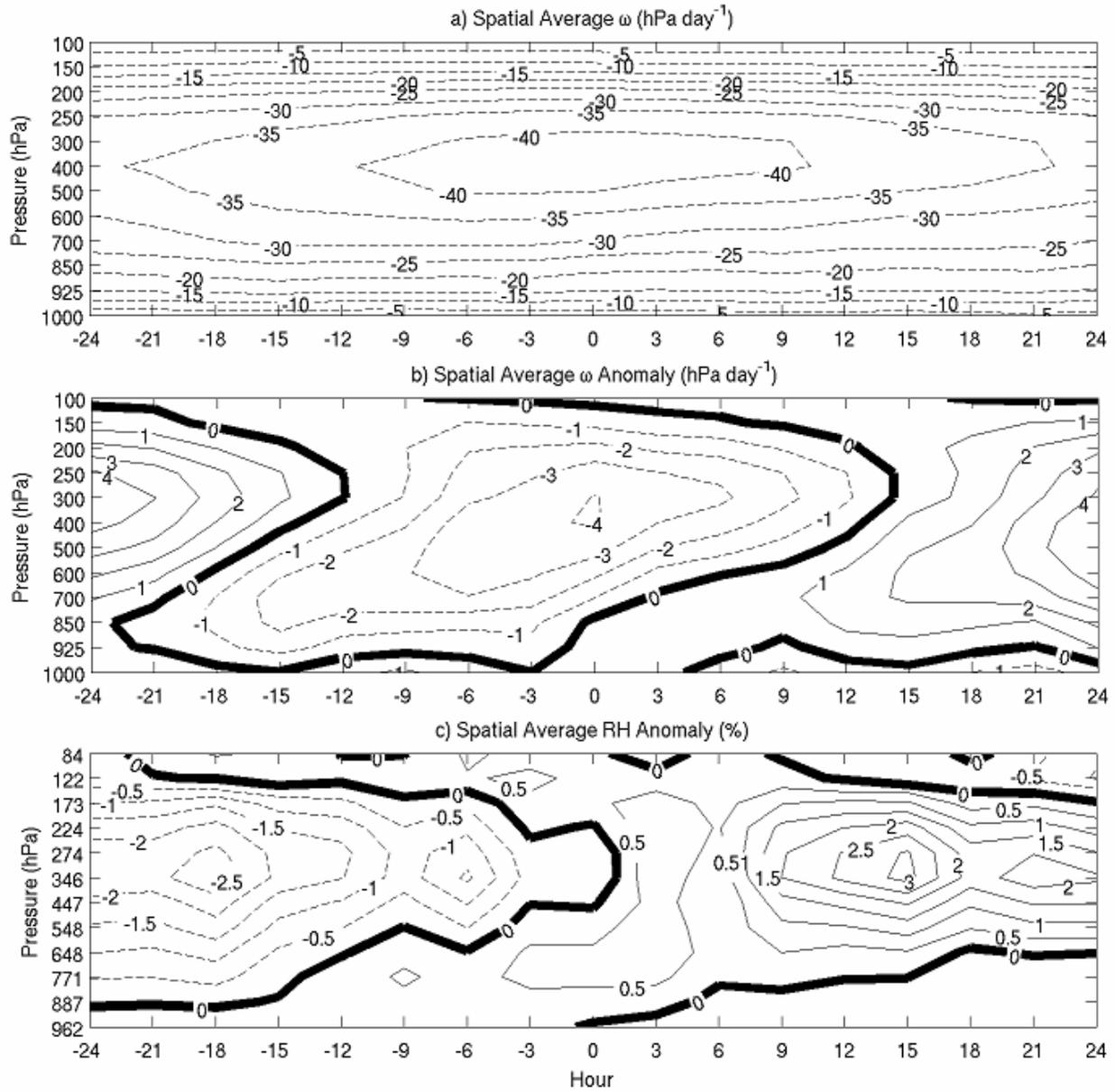
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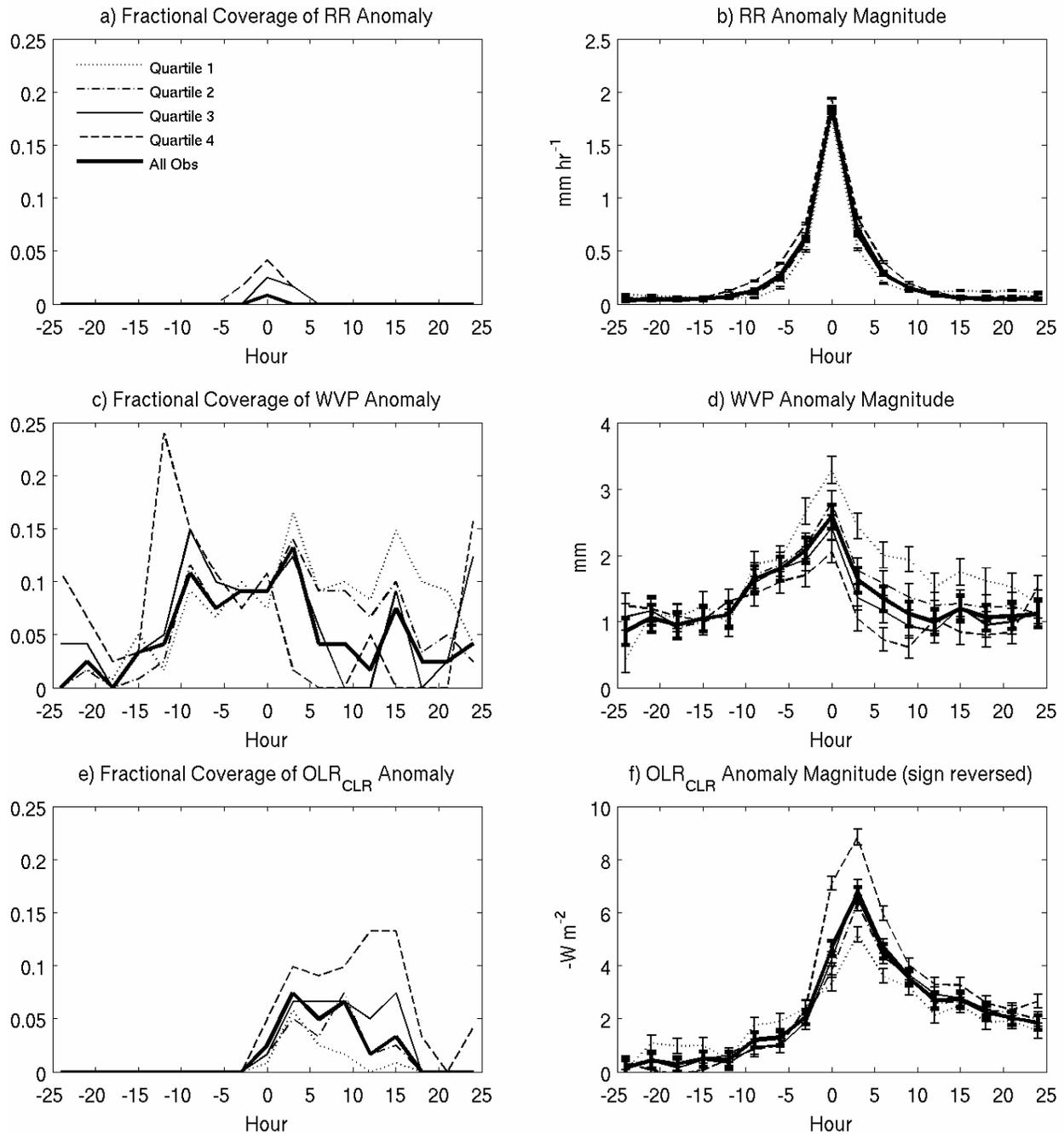
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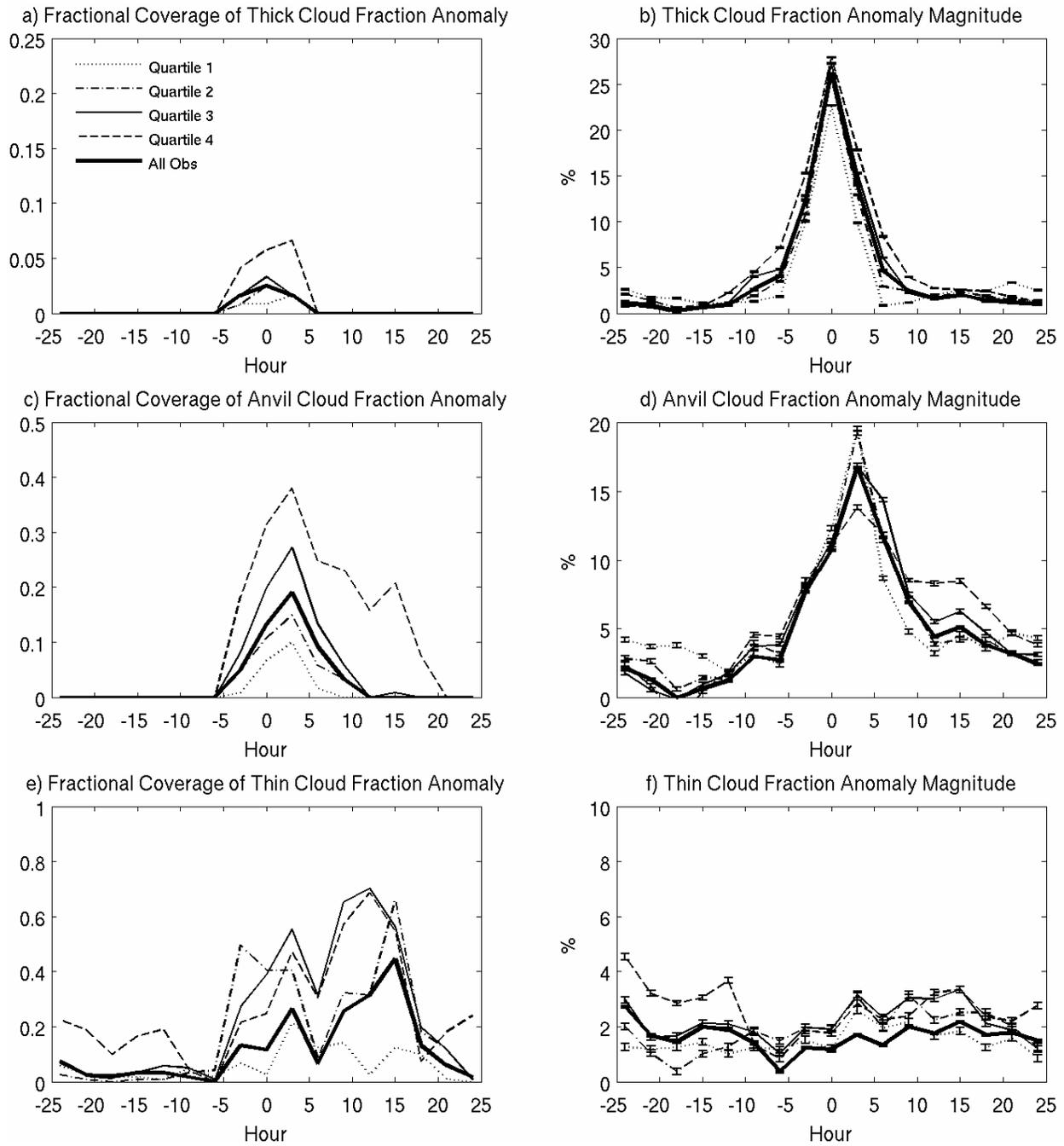
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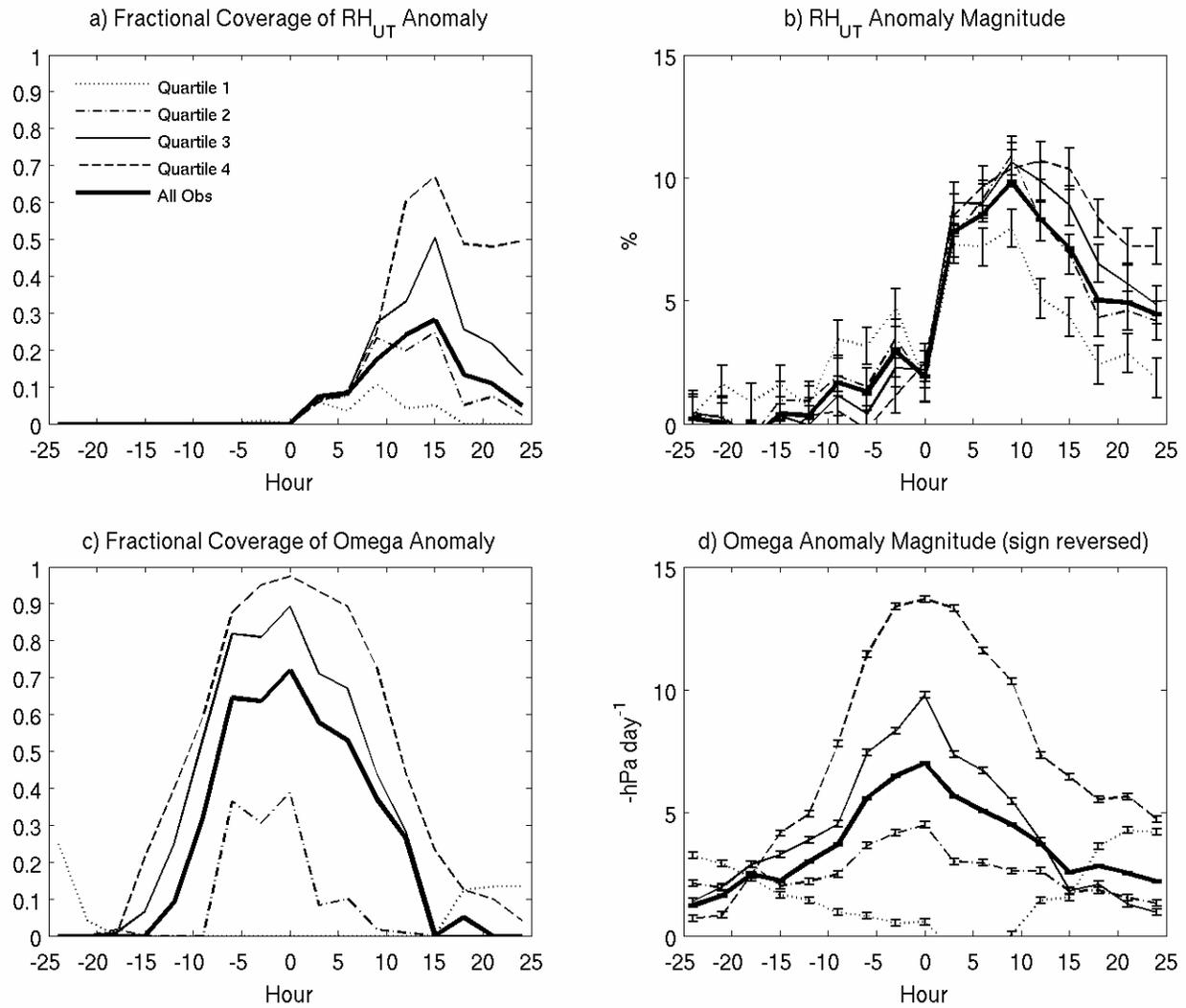
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Figure 7. (Left column) Fractional coverage and (right column) magnitude of (top row) RR, (middle row) WVP, and (bottom row) OLR_{CLR} anomalies for composites in each quartile of spatially-averaged RR. The thick solid line shows the fractional coverage and amplitude of these quantities for all composites centered on RRs exceeding the 90th percentile. Error bars represent the 95% confidence intervals. Note that the sign of the OLR_{CLR} anomaly magnitude is reversed for the purposes of comparison and that the y-axes are different among the anomaly magnitudes.



761

762 Figure 8. Same as in Figure 7, but for (top row) high thick, (middle row) anvil, and (bottom
 763 row) thin cloud fraction anomalies. Note that the y-axes are different for each row.



764

765 Figure 9. Same as in Figure 7, but for (top row) RH_{UT} and (bottom row) ω_{UT} anomalies. Note
 766 that the sign of the ω_{UT} anomaly magnitude is reversed for the purposes of comparison.

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