

Workshop on an Optical Clock Mission in ESA's Cosmic Vision Program
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Gravitational Physics with Optical Clocks in Space

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- Introduction
- Overview over some tests of General Relativity
- Scientific goals of proposed missions
- Scenarios of missions with optical clocks
- Clock developments
- Conclusions

Gravity and its foundations

General Relativity

Metric theory of gravity

- Gravitational redshift
- Lense-Thirring effect
- Shapiro delay
- Perihelion shift
- Schiff effect
- Earth & moon free fall
- Binaries dynamics
- ...

Einstein Equivalence Principle

Universality of Free Fall
(Weak equivalence princip.)

Local Position Invariance
(Universality of grav. Redshift
constancy of constants)

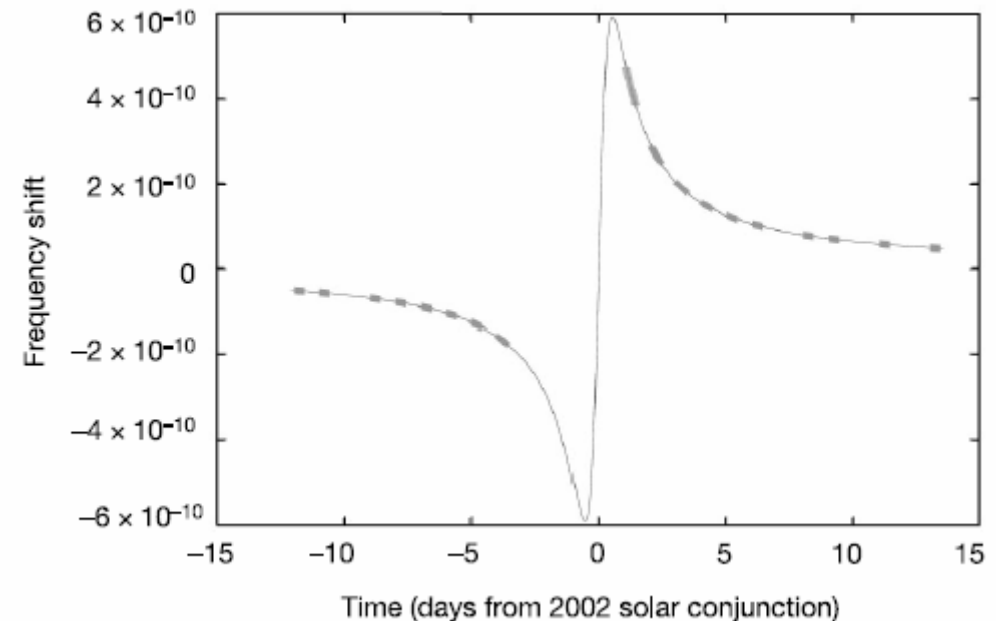
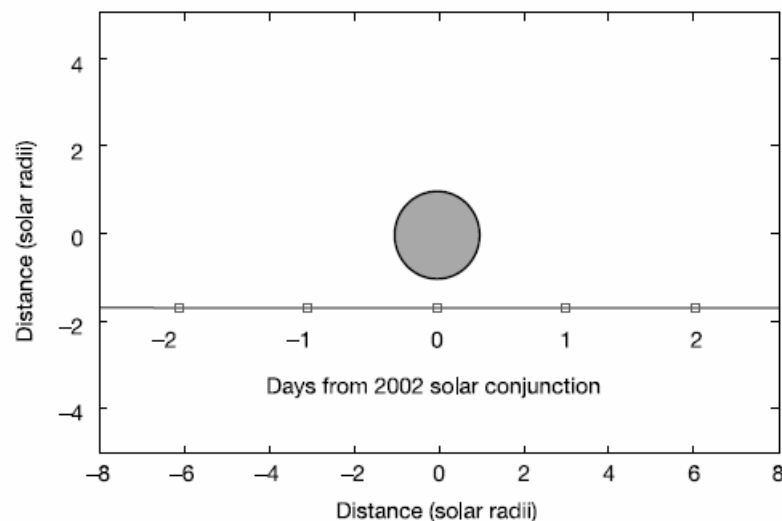
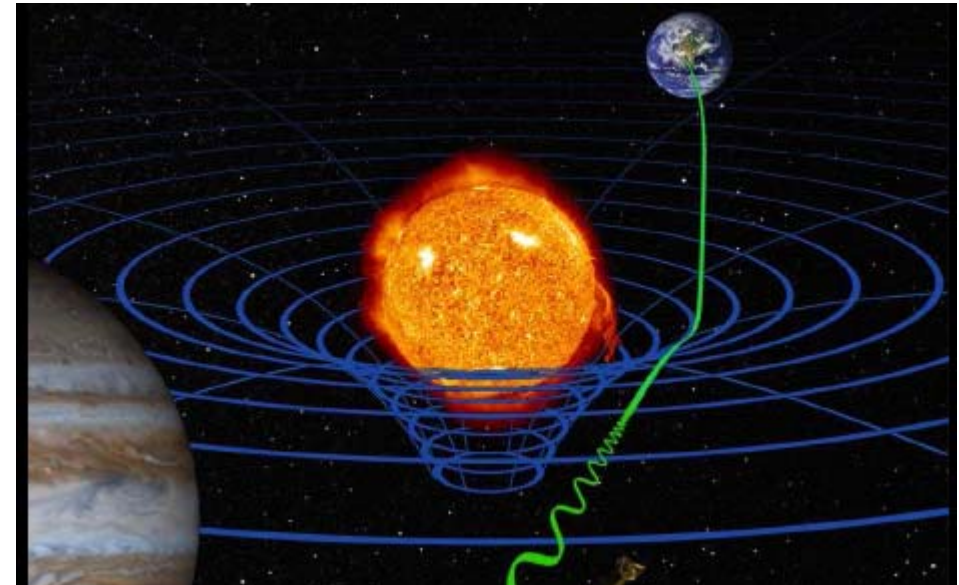
Local Lorentz Invariance
(Special Relativity)

Shapiro time delay

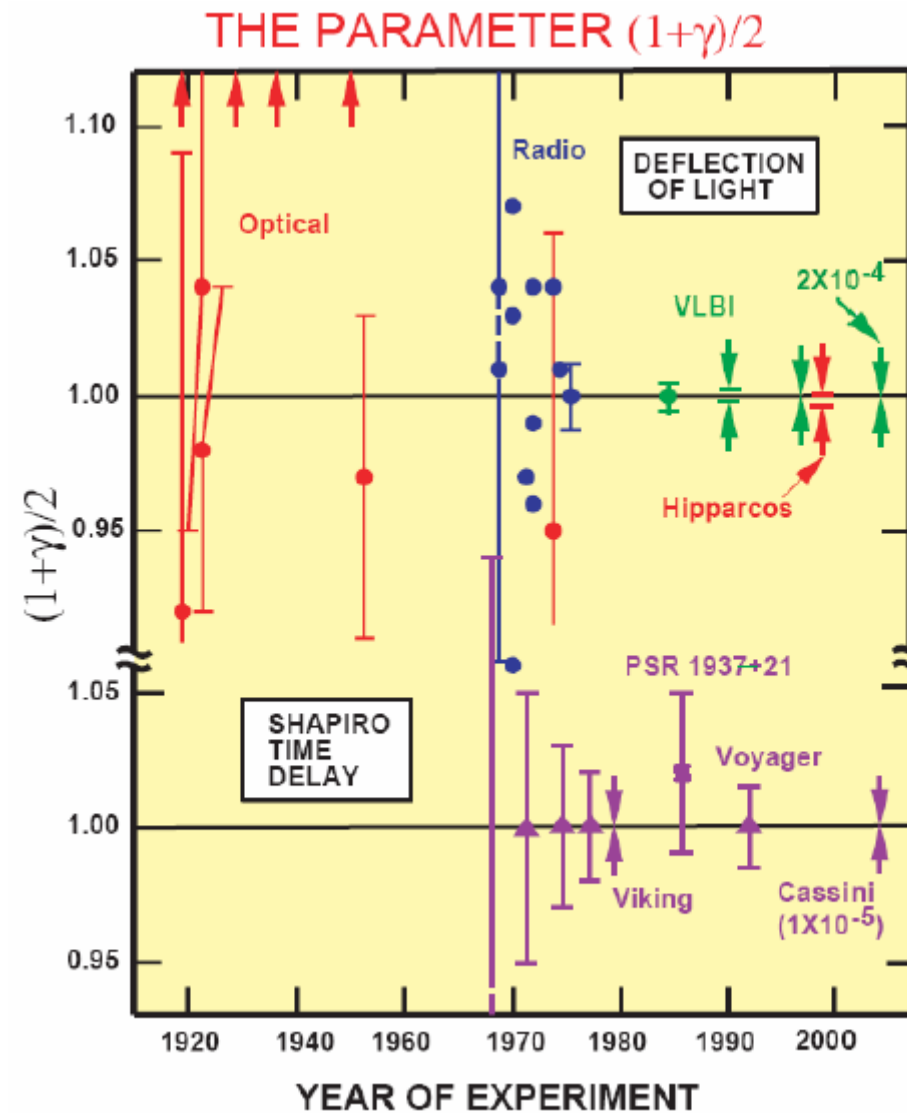
- Cassini Mission
(Tortora et al 2003)

$$\begin{aligned} \gamma_{\text{gr}} &= \frac{d\Delta t}{dt} = -2(1 + \gamma) \frac{GM_{\text{S}} db}{c^3 b dt} \\ &= -(1 \times 10^{-5} \text{s})(1 + \gamma) \frac{1 db}{b dt} \end{aligned}$$

- Achieved accuracy:
 $|1 - \gamma| < 2 \cdot 10^{-5}$



Time delay and deflection of light



From: C. Will (2006)

Nonlinearity of gravity

- Nonlinearity of metric

$$g_{00} = 1 + 2\frac{U}{c^2} + 2\beta\frac{U^2}{c^4} + \dots$$

- From Lunar Laser Ranging results and assuming only β and γ nonzero: $|1-\beta| < 2 \cdot 10^{-4}$

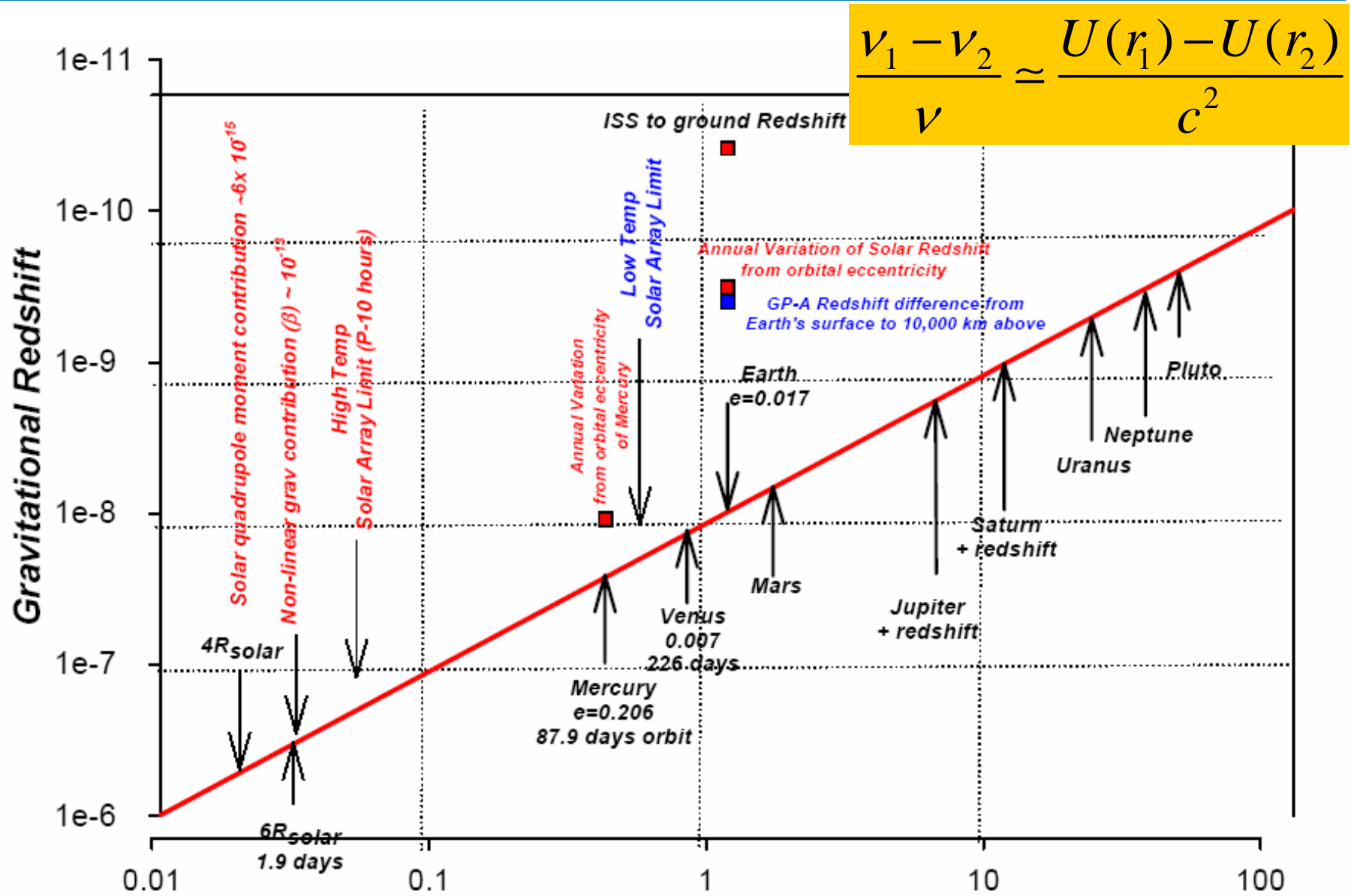
Testing the Gravitational Redshift of Clocks

The diagram illustrates a test of gravitational redshift. A grid represents the gravitational potential U , which is deeper near the Earth. A "Reference Clock" V_0 is located on the Earth's surface. A "Clock ensemble" consisting of two clocks, V_1 and V_2 , is shown in orbit. A yellow line represents the path of the clock ensemble, and a red dashed line represents the communication link between the reference clock and the ensemble. The potential U is shown as a grid that is more compressed near the Earth, indicating a stronger gravitational field.

$$v_i / v_0 = \sqrt{\frac{1 + 2U(r_i)/c^2}{1 + 2U(r_0)/c^2}}$$
$$\frac{\Delta v_i}{v_0} = \zeta_i \frac{\Delta U}{c^2} + \dots$$

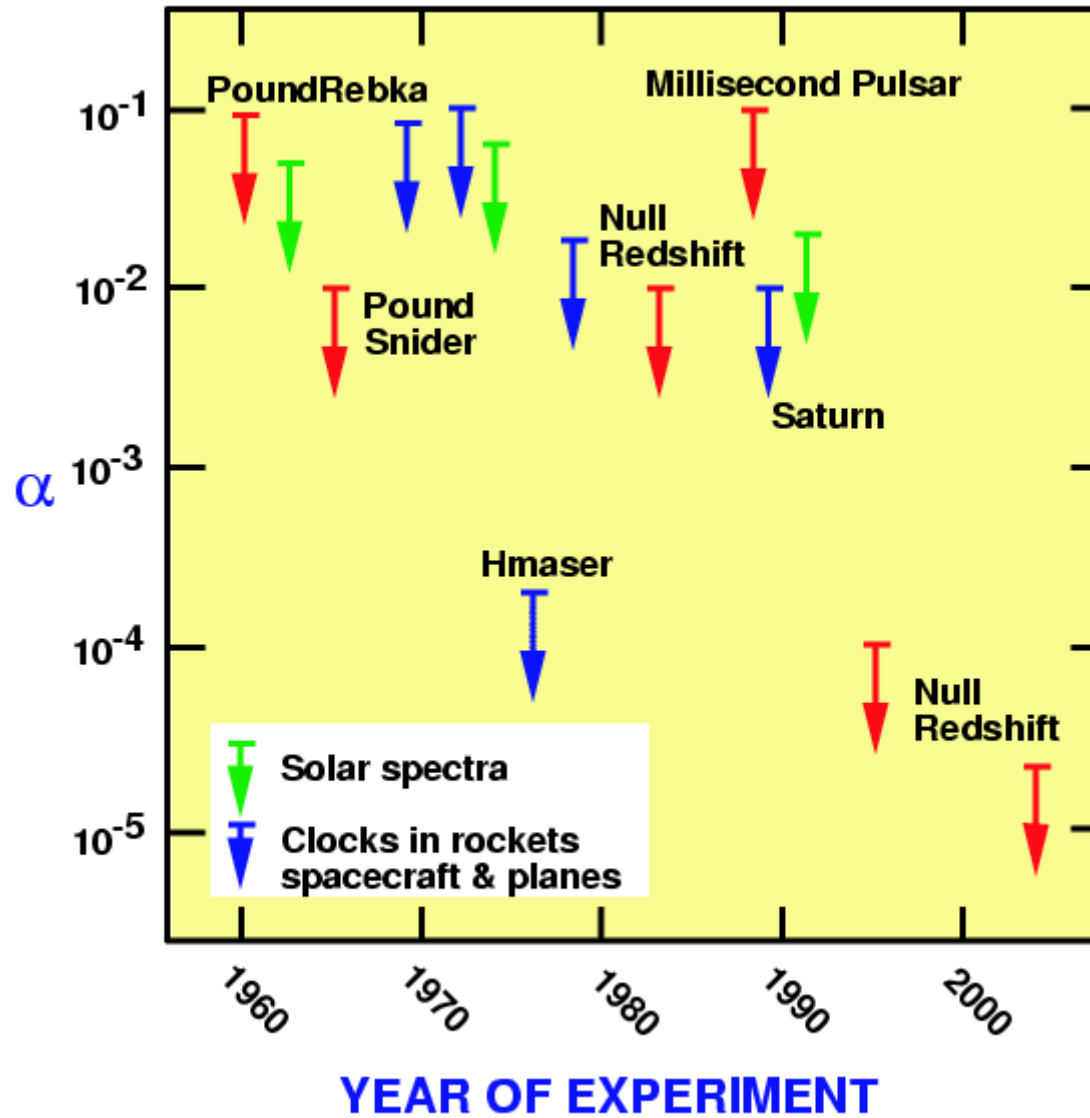
- Comparison with ground clock (*via microwave/optical link*)
 - **Absolute** gravitational redshift measurement
 - Test of higher-order relativistic corrections (*Linet & Teyssandier, 2002, Blanchet et al 2001, Ashby 1998*)
- Intercomparison of **dissimilar** on-board clocks:
 - Gravitational redshift **universality** test (Local Position Invariance): $\zeta_1 = \zeta_2$?

Gravitational Redshift



From: J. Prestage and L. Maleki, JPL *Distance from Sun (AU)*

TESTS OF LOCAL POSITION INVARIANCE



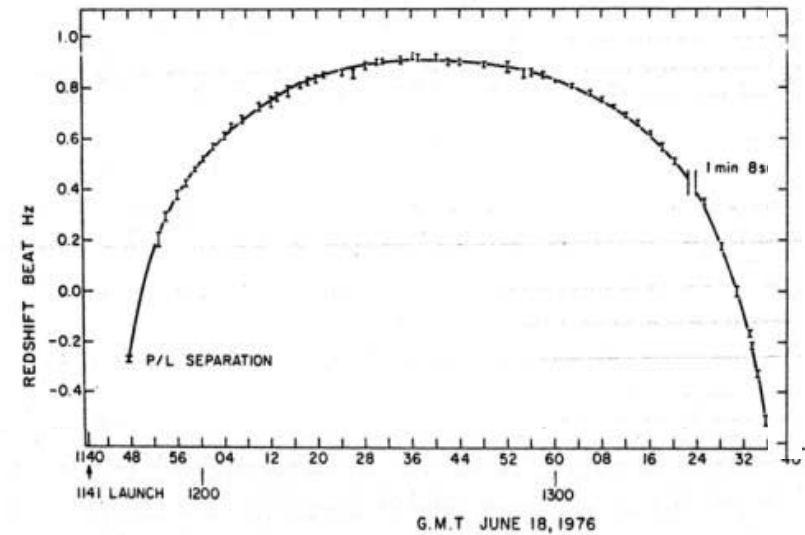
From: C. Will (2006)

$$\Delta v/v = (1+\alpha)\Delta U/c^2$$

Gravitational redshift: Past & upcoming missions

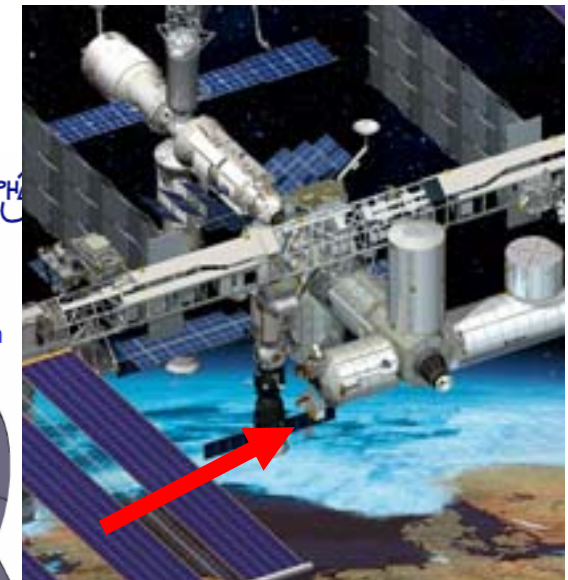
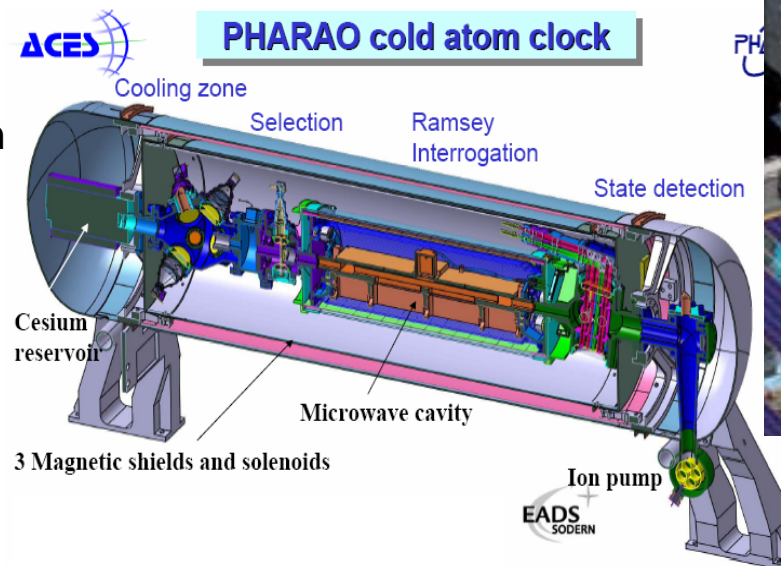
Gravity Probe A: hydrogen maser (1976)

- rocket flight to 10 000 km altitude
- tested relativistic Doppler effect and gravitational redshift to 70 ppm



ACES: Atomic clock ensemble (2012)

- PHARAO: cold atom microwave clock
- instability $1 \cdot 10^{-13}$ at 1 s, $4 \cdot 10^{-16}$ at 50 000 s
- accuracy $\sim 1 \cdot 10^{-16}$
- redshift test at 2 ppm
- technology demonstrator
- world-wide time dissemination and comparisons
- test of special and general relativity



Scientific Goals

- How scientifically powerful?
- „The most powerful test of gravitational theory“
 - Gravity Probe A: $7 \cdot 10^{-5}$ (redshift)
 - Cassini: $2 \cdot 10^{-5}$ (γ)
 - Gravity Probe B: goal $1 \cdot 10^{-5}$ (γ)
 - ACES: goal $2 \cdot 10^{-6}$ (redshift)
- **Proposals:**
 - Mercury Radioscience Orbiter Experiment: $\Delta\gamma \sim 2 \cdot 10^{-6}$, $\Delta\beta \sim 5 \cdot 10^{-6}$
 - GAIA $5 \cdot 10^{-7}$ (spacecraft at Lagrange-point L1)
 - ASTROD I: γ at $1 \cdot 10^{-7}$, β at $1 \cdot 10^{-7}$ (1 spacecraft, drag-free)
 - Gravitational Time Delay Mission: γ at $2 \cdot 10^{-8}$ (2 spacecraft, drag-free)
 - LATOR: γ at $2 \cdot 10^{-9}$ (3 spacecraft, incl. ISS, not drag-free)
 - ASTROD: γ at $1 \cdot 10^{-9}$ (3 spacecraft, drag-free)
- **Earth-based tests:** Local Position Invariance
(U/c^2 daily amplitude: $\sim 4 \cdot 10^{-13}$, yearly amplitude $\sim 2 \cdot 10^{-10}$)
Bauch and Weyers (2002), upcoming results with Cs & optical clocks

- *Damour and Nordtvedt (1993), Damour (1999), Damour, Piazza, Veneziano (2002):*
existence of scalar fields (dilaton) that violate EP, strength: $\gamma - 1$
 - model takes into account inflation and WMAP measurements;
 - γ is time-dependent, $=0$ in early universe, nearly 1 now; $1-\gamma \sim 5 \cdot 10^{-5} - 5 \cdot 10^{-8}$
 - *Within the dilaton model, the earth-based Equivalence Principle tests have already shown $|1-\gamma| < 2 \cdot 10^{-7}$ (to be improved by MICROSCOPE), and predict $d \ln \alpha / dt < 10^{-20} / \text{yr}$*
 - *But EP tests and γ measurement only probe hadronic matter and Coulomb energy; hyperfine and molecular clocks also probe leptonic matter (electron mass)*
- Alternative explanation to Dark Energy: extension of GR in the low-energy regime (*Carroll et al. 2004*), $1-\gamma \sim 10^{-9} - 5 \cdot 10^{-7}$
- *Sandvik et al (2002):* Local Position Invariance for α may be violated at level $\sim 10^{-4}$ (ruled out now)
- *See also Lämmerzahl (2006), Turyshev et al., in Dittus et al. (2007)*

Second-order effects

- Achievable values of U/c^2 in the solar system are of order
 - 1. 10^{-8} for a spacecraft going to Mercury or outer planets
 - 3. 10^{-7} for a spacecraft approaching sun
 - 1. 10^{-6} for a wave grazing the sun
- Clocks of $1 \cdot 10^{-18}$ accuracy, would allow a test of GR at 10^{-10} level
- Effects of second order in U/c^2 (still in „weak-field“ regime)
- Achievable values of U/c^2 in our solar system imply that resolution of measurement must be $1 \cdot 10^{-12}$ or better
 - ASTROD, LATOR, Gravitational Time Delay would be sensitive to second-order effects (probe sun field and aim at relative accuracies of measured PN parameter beyond $1 \cdot 10^{-6}$)
- Clocks could also allow a sensitive test of second-order effects

„The most precise test of general relativity“

$$v_i / v_0 = \sqrt{\frac{1 + 2U(r_i)/c^2}{1 + 2U(r_0)/c^2}}$$

Gravitational Redshift and PN formalism

- Contribution to redshift from the two PN parameters β, γ in a fully conservative metric theory without preferred location effects (*see Teyssandier et al (2007)*)

$$\frac{\nu_1 - \nu_2}{\nu} = \dots - (1 + \gamma) \left(\frac{U(r_1)}{c^2} \frac{\nu_1^2}{c^2} - \frac{U(r_2)}{c^2} \frac{\nu_2^2}{c^2} \right) + (\beta - 1) \left(\frac{U(r_1)^2}{c^4} - \frac{U(r_2)^2}{c^4} \right)$$

- Present accuracy of β ($2 \cdot 10^{-4}$) and γ ($2 \cdot 10^{-5}$) rules out any effects observable with clocks for solar-system level U
- Clocks test a different sector of the theory:
LPI violation, theories beyond PN theory

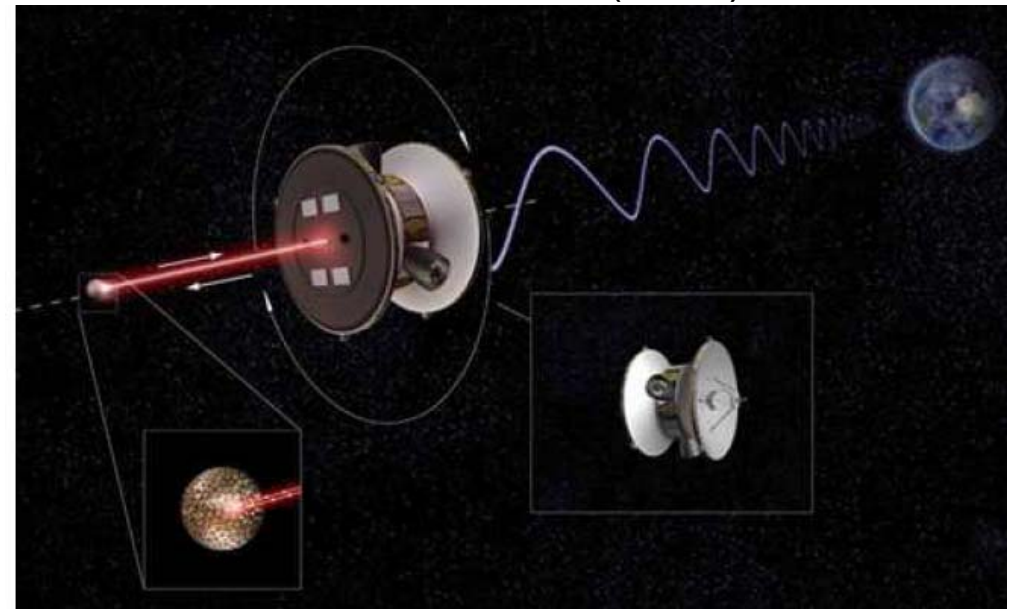
Complexity and Cost

- Drivers:
 - Number of spacecrafts: 1, 2 or 3
 - Distance of travel from earth
 - Type and number of dissimilar clocks
 - Frequency link to earth or between spacecrafts
 - no link: only Local Position Invariance test
 - link: also absolute gravitational redshift
 - link to earth: limited by inaccuracy of gravitational potential on earth ($\sim 10^{-18}$, similar to expected clock accuracy)
 - Drag-free
 - Additional measurements (e.g. Pioneer anomaly, Lorentz Invariance, geophysics, orbit dynamics)

Mission to outer solar system - Pioneer anomaly

Dittus et al., Firenze (2006)

- Main measurement: ranging of spacecraft while at large distance from earth (main s/c + free-flyer)
- Clock on board to sense anomalous acceleration: to achieve 1% accuracy in the anomalous acceleration, need a clock of 10^{-15} long-term (~ 10 years) accuracy



- Additional payload: optical clock would enable accurate measurement of gravitational redshift over planetary distance (first section of voyage) ($\Delta U/c^2 \sim 1 \cdot 10^{-8}$ relative to earth; earth gravitational potential limits accuracy at $1 \cdot 10^{-18}$, allowing 1% of second-order contribution)
- Link at $1 \cdot 10^{-18}$ over inter-planetary distances possible?
- Need to know distances to sun with $\Delta r_{\text{earth-sun}} \sim 15$ m, $\Delta r_{\text{s/c-sun}} \sim 140$ m, achievable
- LPI test could test second order-effect at 1%

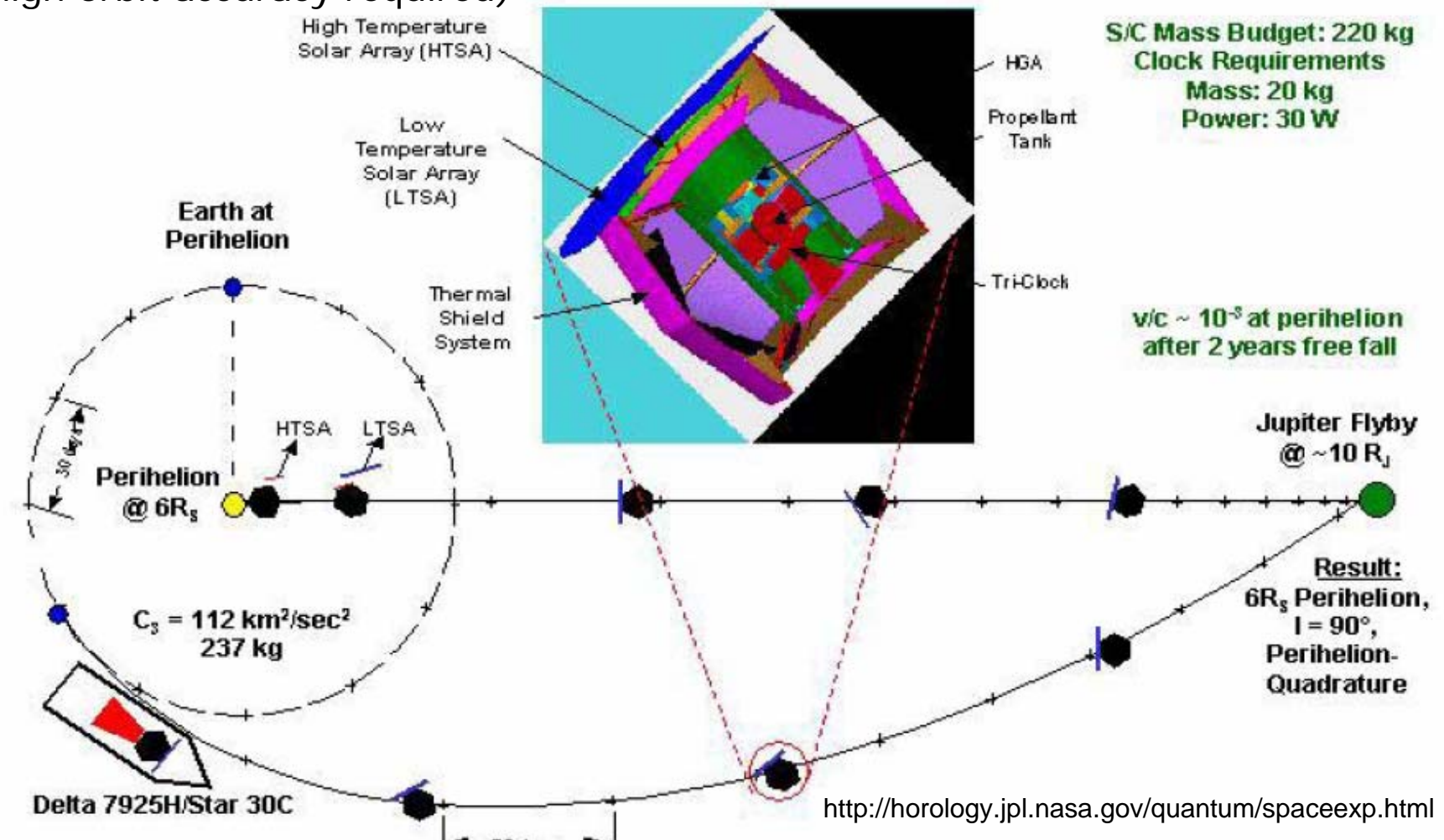
Mission to inner solar system

- Flight to mercury provides $\Delta U/c^2 \sim 1.5 \cdot 10^{-8}$, $\frac{1}{2} \Delta(U/c^2)^2 \sim 2 \cdot 10^{-16}$.
 - Comparison with earth clock can test second-order effects at 1% level
 - LPI test at 1 % of second-order
 - Interplanetary link at $1 \cdot 10^{-18}$ possible?
- Need to know distances to sun with $\Delta r_{\text{earth-sun}} \sim 15 \text{ m}$, $\Delta r_{\text{s/c-sun}} \sim 2 \text{ m}$, achievable
- Additional science goal: combine with time delay measurement when s/c is in conjunction
 - From ground: ASTROD I-type mission, $\Delta\gamma$ at $\sim 1 \cdot 10^{-7}$, $\Delta\beta$ at $\sim 1 \cdot 10^{-7}$
 - add second spacecraft; GTD-type mission, $\gamma \sim 1 \cdot 10^{-8}$

Solar Fly-by (SpaceTime Mission)

Maleki et al. - JPL

- Flyby at 6 solar radii gives a potential variation $\sim 3 \cdot 10^{-7}$ along orbit; LPI test using microwave ion clocks (room-temperature)
- Optical clocks would be an alternative, allowing LPI test at 10^{-11} level
- Gravitational redshift measurement making use of the full $\Delta U/c^2$ seems too difficult (very high orbit accuracy required)



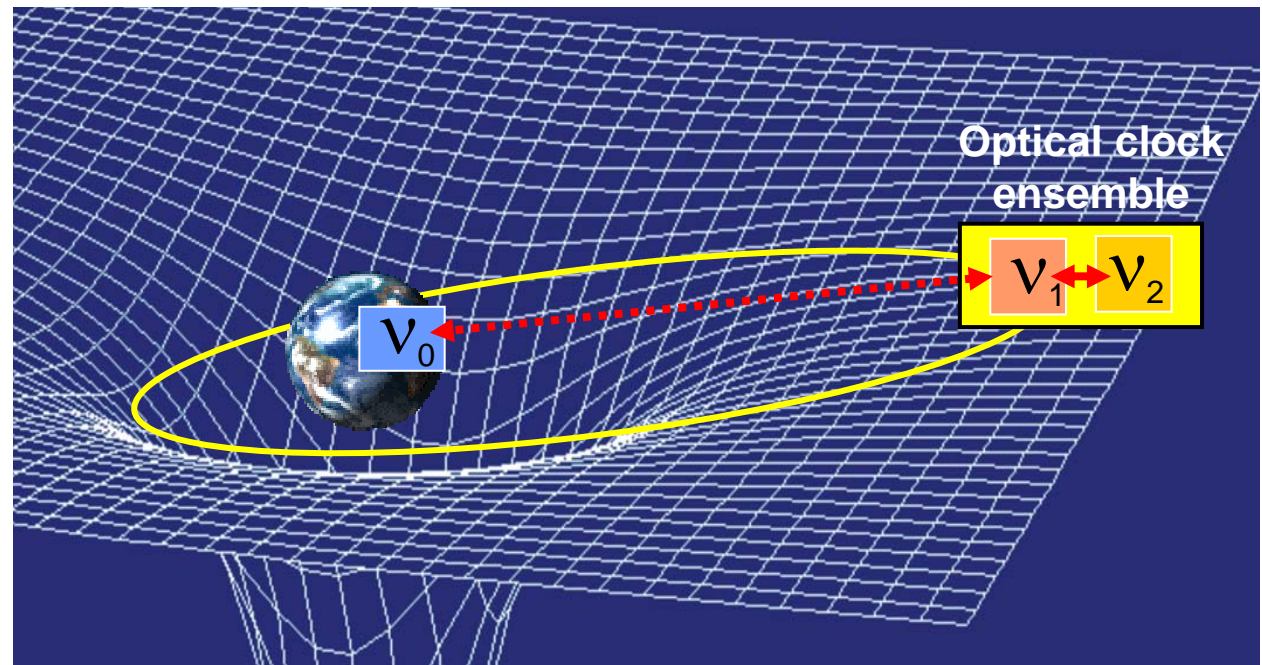
Earth orbit mission

- A constant distance, high-altitude earth orbit, e.g. geostationary: $\Delta U/c^2 \sim 6 \cdot 10^{-10}$, $\frac{1}{2} \Delta(U/c^2)^2 \sim 2 \cdot 10^{-19}$
 - But: current uncertainty in earth gravitational potential (~ 1 cm) implies a $\sim 1 \cdot 10^{-18}$ uncertainty
 - Future clocks and improved earth models could measure 2nd-order effect
 - Highly elliptic orbit: avoids earth gravitational potential uncertainty, as long as earth potential is constant to fraction of % over orbital period (~ 0.5 d);
 - Such an orbit also allows LPI test
 - variation in U is few 10^{-10} , so test barely at second-order level (averaging over many orbits)

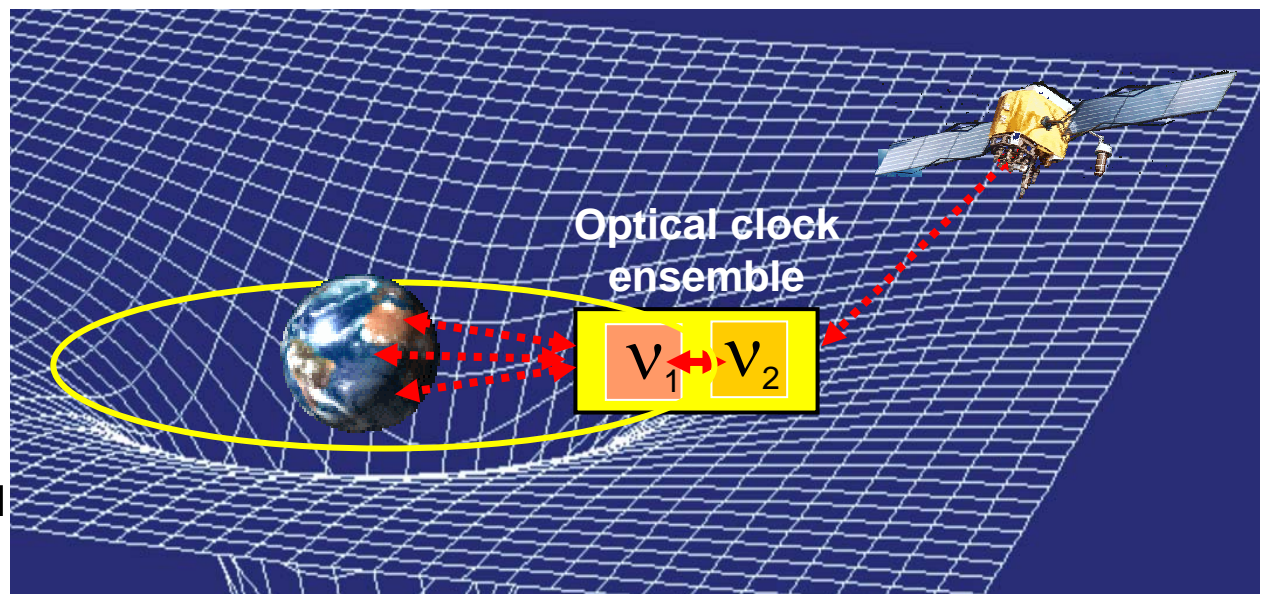
Gravity Explorer

Schiller et al, (2005, 2007)

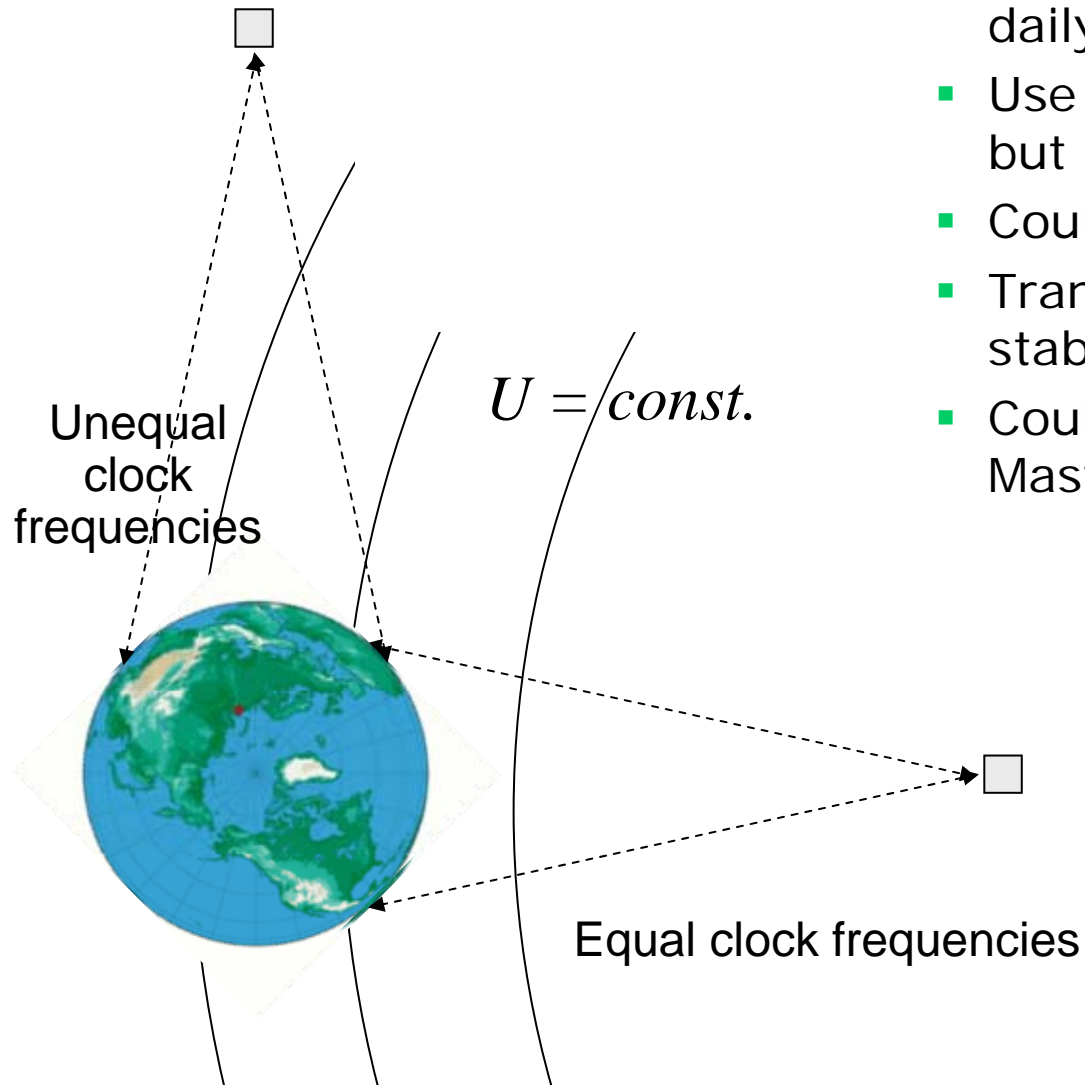
- Orbital phase I
(~ 1 year duration,
highly elliptic orbit)
 - Test of Local Position
Invariance and
of grav. Redshift
($2 \cdot 10^{-10}$ amplitude)



- Orbital phase II
(geostationary,
several years duration)
 - Master clock for earth
and space users
 - Geophysics
 - Ground clock
comparison (sun redshift
ampl. $4 \cdot 10^{-13}$)
 - LPI & Redshift in sun field
(amplitude $2 \cdot 10^{-12}$)

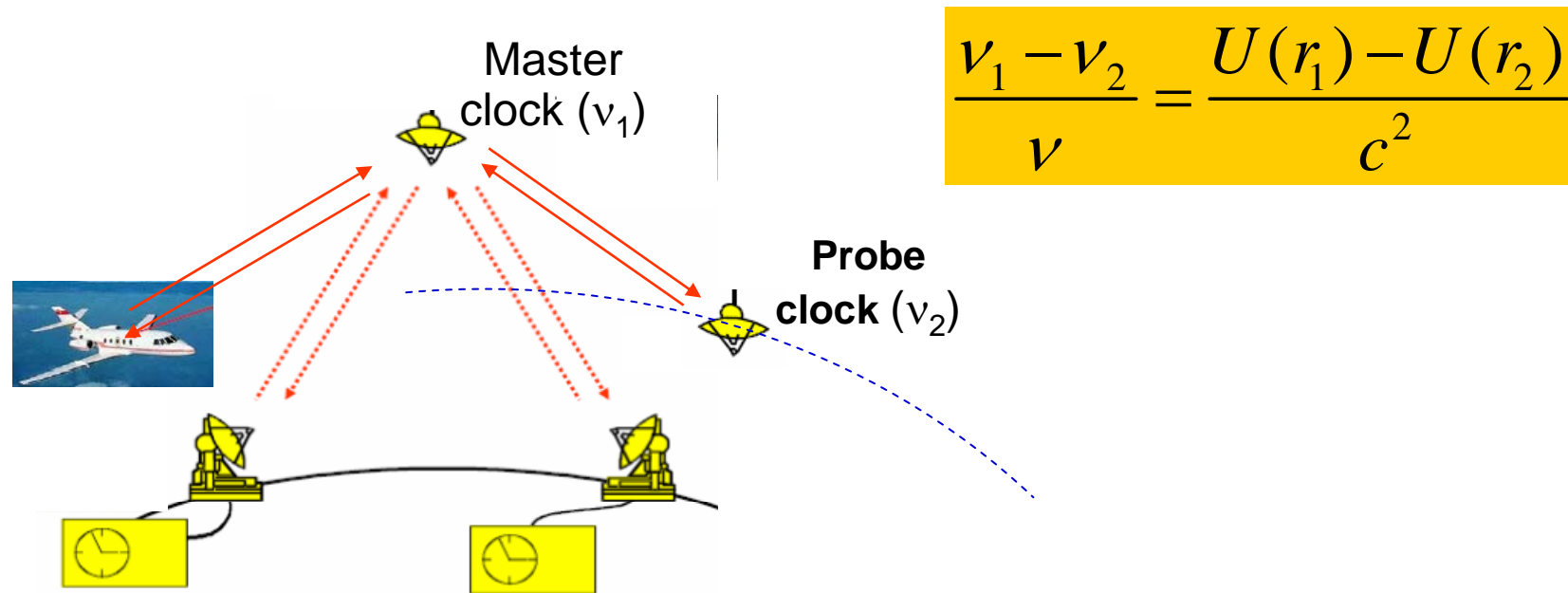


Ground clock comparisons



- Transponder satellite for terrestrial clock comparison
- Solar clock redshift has daily amplitude of $4 \cdot 10^{-13}$
- Use for testing gravity: small effect but long duration and lower cost
- Could also be used for geophysics
- Transponder could require a stable laser (frequency comb?)
- Could be combined with Master Clock concept

Clock comparisons



$$\frac{v_1 - v_2}{v} = \frac{U(r_1) - U(r_2)}{c^2}$$

- Comparison between master and probe clock (*via microwave/optical link*)
- *yields information about gravitational potential U*
- Features:
 - non-local measurement (cf. with two-satellite measurements such as GRACE or SSI)
 - measurement time to reach 10⁻¹⁸ ~ 10 h (? - limited by link)
 - independent of satellite acceleration

Gravity (geopotential) measurements

A clock accuracy $\frac{\Delta \nu}{\nu} = 10^{-18}$

yields $\Delta \left(\frac{U(r)}{c^2} \right) = 10^{-18} \Rightarrow \frac{\Delta U}{U} \sim 10^{-9}$ near earth's surface
(equivalent to 1 cm height change)

This requires a position accuracy:

$$\Delta r = \begin{cases} 30 \text{ cm} & \text{for a geostationary orbit} \\ 1 \text{ cm} & \text{for a LEO orbit} \end{cases}$$

Orbitography at a level of 3 cm via GPS is available,
- potential for improvement using laser ranging?

Compare:

GOCE: Measurement of g with $1 \text{ mGal} = 10^{-5} \text{ m/s}^2$ resolution, i.e. $\Delta g/g \sim 10^{-6}$

Absolute (corner-cube) gravimeters: resolution

$$\Delta g/g \sim 10^{-9}$$

Superconducting gravimeters (stationary): resolution

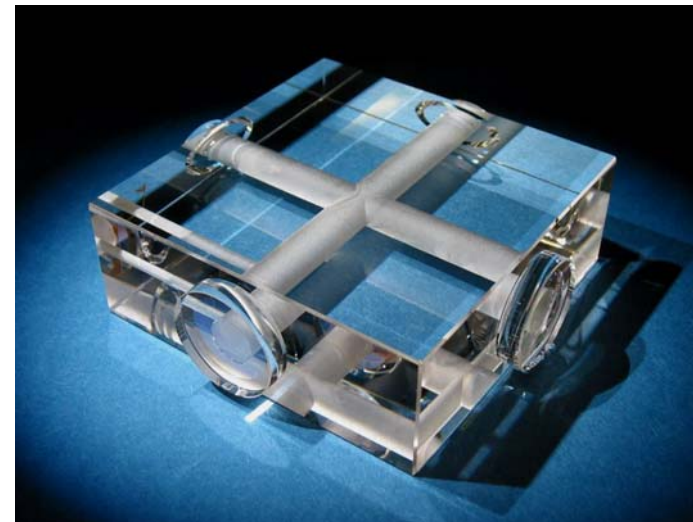
$$\Delta g/g \sim 10^{-11}$$

Geophysics perspectives

- Desirable is a geoid measurement with better than 1 mm accuracy (Sumatra earthquake produced geoid variations of -6 to +12 mm across the fault).
- This requires clocks with accuracy $\frac{\Delta\nu}{\nu} \sim 10^{-19}$

Further possible goals of a clock mission

- Test of isotropy of space (Michelson-Morley expt.):
requires an additional optical cavity & optoelec. components
- Test of constancy of speed of light (Kennedy-Thorndike expt.):
no additional components
- Test of Lorentz Invariance (large s/c velocity helpful)
- For particular earth orbits: test of Lense-Thirring effect (laser ranging retroreflectors; requires drag-free satellite)
- *See OPTIS proposal (Lämmerzahl et al 2001, 2004, Iorio et al 2004)*



C. Eisele, A. Nevsky, M. Okhapkin, S.S.

Clocks for Local Position Invariance Tests

- Frequencies depend on fundamental constants

$$\nu_i = \nu_i(\alpha, \alpha_S, G_F, m_e, m_N, g_N, \dots)$$

- Gravitational redshift experiments test whether some of these constants depend on the gravitational potential

$$\beta_j = \beta_j(U)? \quad \Rightarrow \quad \zeta_i = 1 + \sum_j \left(\nu_i^{-1} \frac{d\nu_i}{d\beta_j} \right) \left(\frac{d\beta_j}{d(U/c^2)} \right)$$

- Some constants can be related to more fundamental constants:

α

Electromagnetic interaction

$$m_p \propto \Lambda_{QCD} + \text{corrections } (m_q, m_s)$$

Strong interaction

$$g_N = g_N(m_q / \Lambda_{QCD}, m_s / \Lambda_{QCD})$$

$$m_e \propto \langle \phi \rangle = \text{Higgs vacuum field}$$

Weak interaction

$$\frac{\Delta(m_N/m_p)}{m_N/m_p} = c_\alpha \frac{\Delta\alpha}{\alpha} + c_\phi \frac{\Delta\phi}{\phi} + c_\Lambda \frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}} \quad c_\alpha, c_\phi, c_\Lambda \approx O(10^{-3})$$

Optical clock types

Completeness:

Scaling of transition energies (in units of Rydberg energy)

- Hyperfine energies [1,2] $g \frac{m_e}{m_p} \alpha^2 F(Z\alpha)$
 - Electronic energies (incl. relativistic effects) $G(\alpha)$ →
 - Vibrational energies in molecules [3,4] $\sqrt{m_e/m_N}$
 - Rotational energies in molecules m_e/m_N
 - Cavity frequency α^{-1}
 - Nuclear energies [5] $H(\alpha, m_q / \Lambda_{QCD}, m_s / \Lambda_{QCD})$
- | | |
|-------------------|--------------|
| Yb: | 0.31 |
| Sr: | 0.06 |
| Yb ⁺ : | (0.9, - 5.3) |

**Electronic transitions do not furnish a complete test!
Complement with hyperfine, nuclear, or molecular clock**

[1] Microwave cold atom clocks (PHARAO)

[2] (near-optical) range: highly charged atomic ions (S. Schiller, 2007)

[3] L. Hilico et al. (2000)

[4] S. Schiller and V. Korobov (2005)

[5] V. Flambaum (2006)

Clock choice

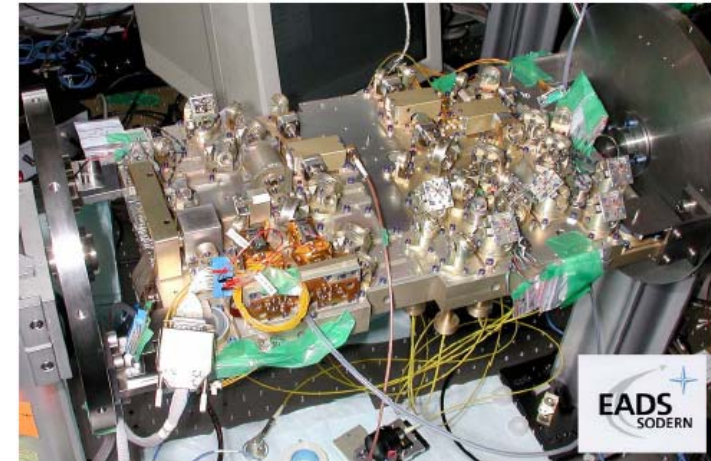
- A comparison of an atomic optical clock to a molecular optical clock is (within the Standard Model) sensitive to largest number of fundamental constants
- In gauge unification theories the dependencies of α and Λ_{QCD} on U are correlated (*Damour 1999, Langacker et al, Calmet & Fritzsche. 2002*)

$$\frac{\Delta\Lambda_{QCD}}{\Lambda_{QCD}} \simeq 40 \frac{\Delta\alpha}{\alpha}$$

- **Enhancement effects are desirable**
 - Sensitivity of nuclear transition frequencies to α , m_s , m_q are predicted to have $\sim 10^5$ -fold enhancement (Flambaum, 2006)
 - Other systems?

Space suitability

- Important optical clock components are already space-qualified
 - Single-frequency diode lasers (PHARAO)
 - Ultracold atom sources (PHARAO)
 - Opto-electronic components
 - Solid-state lasers and amplifiers (TESAT Spacecom)
 - Optical resonators (TESAT Spacecom)
 - Phase-locking (TESAT Spacecom)



Mass 22 kg, power 65 W

- Further optical technology on LISA Pathfinder
- Studies toward space qualification and space uses of frequency combs are under way (DLR, ESA)
- Ultracold atoms in free fall studies (Bose-Einstein Condensate) at ZARM Bremen (DLR)
- High-precision time transfer between satellites and earth to be tested in upcoming missions (ACES on ISS, T2L2 on JASON 2)
- Optical link experiments (LCT TerraSAR, LOLA,...)
- Quantum information research is likely to produce important technology also for compact optical clocks

Summary

- Different scenarios – different science output & costs
- LPI test only: „simplest“
- A complete test should include an absolute redshift measurement
new component: link
- A powerful test requires measuring the 2nd-order contributions
spacecraft needs to fly far away from earth
higher cost
- Combination with additional science goals
spacecraft bigger and more expensive
larger community

Cost, complexity, science output



Gravity Explorer
(earth orbit)

Gravity Explorer
(interplanetary orbit)

Optical Clocks +
ASTROD I

Optical Clocks +
Deep Space
Gravity Probe

Summary – Gravity Explorer proposal

S.S. et al. arxiv:gr-qc/0608081



- **Fundamental physics goals:**
(using clocks/links with 10^{-18} instability/accuracy)
 - Measure gravitational redshift with $\sim 10^4$ higher accuracy
 - Test higher-order relativistic effects in frequency comparison
 - Measure 2nd order Doppler effect with $\sim 10^2$ higher accuracy
 - Test independence of fine structure constant α on U with 10^2 higher accuracy*
 - Test independence of m_e/m_p on U with 10^2 higher accuracy*
 - **Additional possibilities**
 - With drag-free satellite, measure Lense-Thirring effect and perigee advance, ~ 10 times more accurately
 - Contribution to tests of time-independence of fundamental constants
 - Test of isotropy of speed of light (requires rotating satellite)
 - Other Local Lorentz Invariance tests
- **Gravity mapping**
 - Enable gravitational potential measurements at $2 \cdot 10^{-10}$ resol. (1 mm equiv.), requires clocks at 10^{-19} accuracy
- **Master clock for earth and space applications**
- **Enable distant ground clock comparisons**
- **Technology demonstration and validation**

*compared to future terrestrial experiments