

# Constraining the Numbers, Luminosities, and Colors of Evolved Stars within Nearby Galaxies

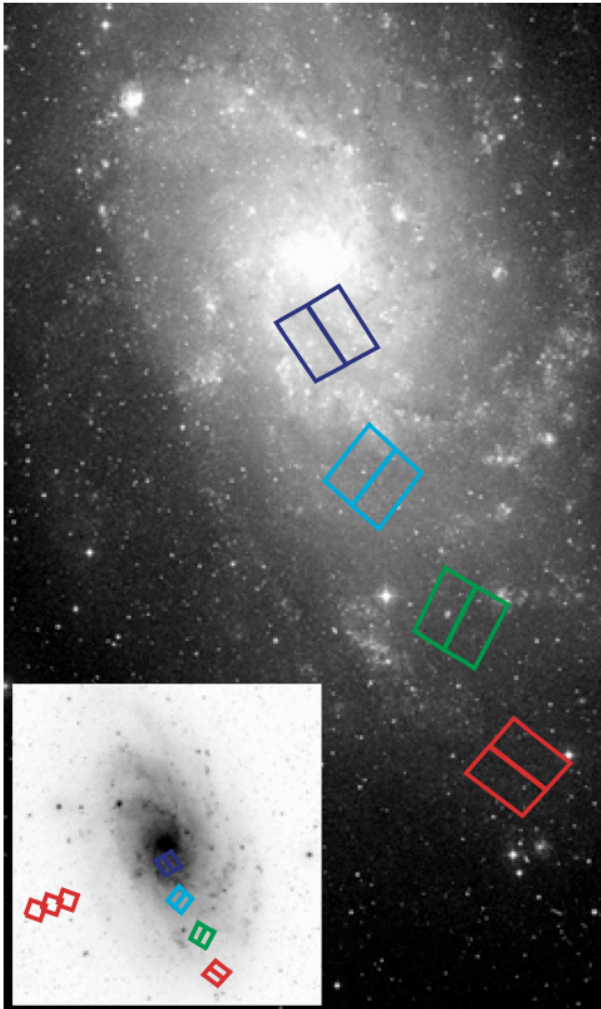
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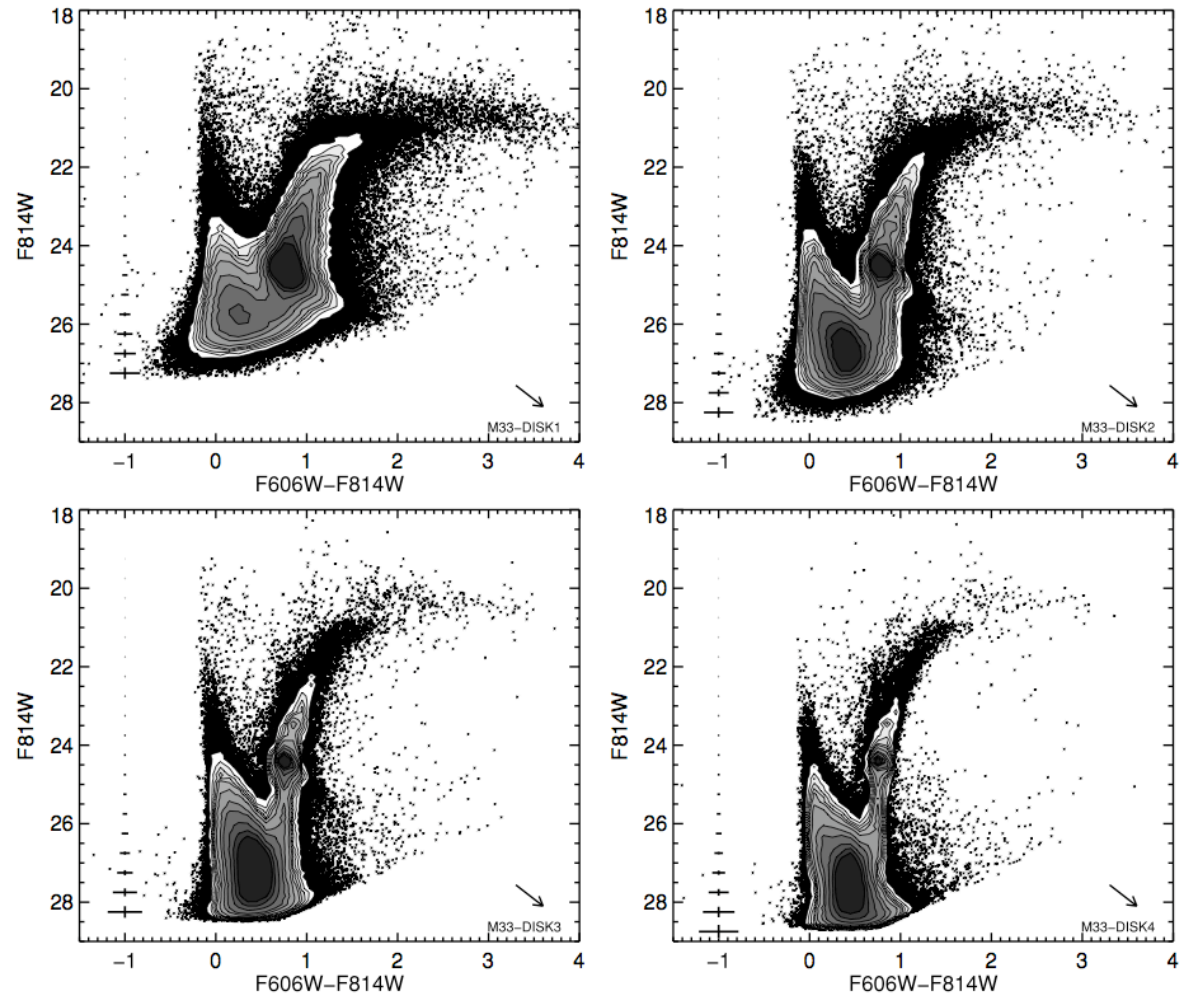


Nearby (2.5 Mpc) Dwarf  
Irregular Galaxy KKH 98  
HST F475W (Blue), HST  
F814W (Green), Keck AO K-  
band (red)

# The Colors and Magnitudes of Stars Can be Used to Reconstruct the Star Formation Histories of Galaxies

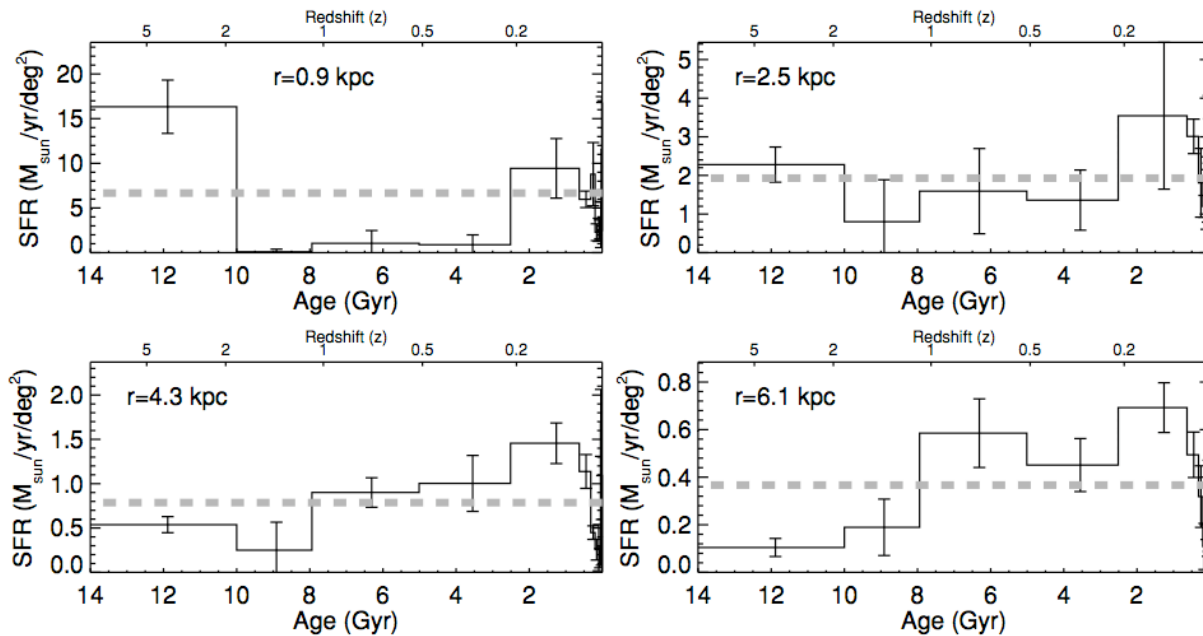
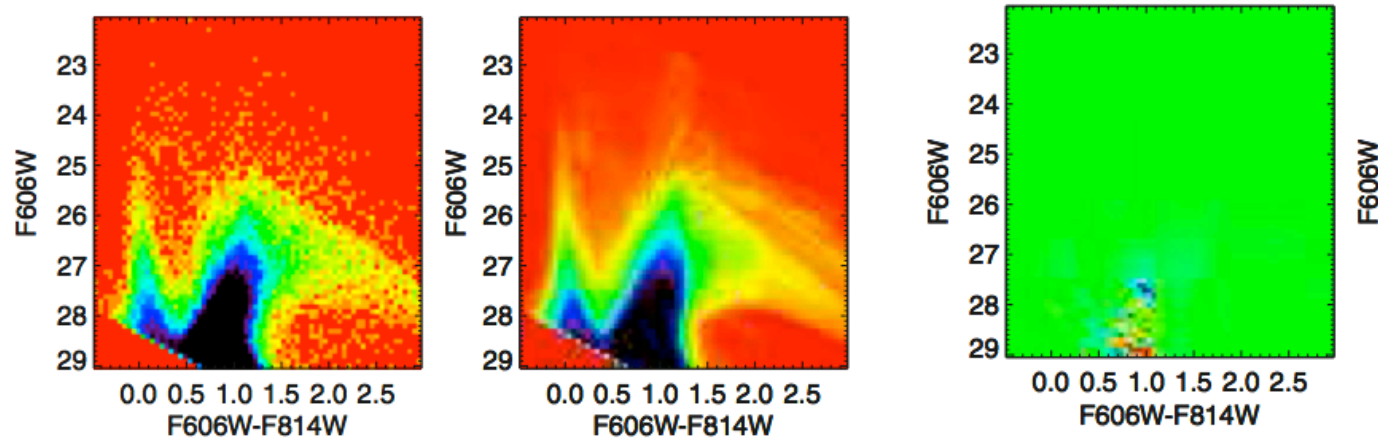


M33 HST ACS fields

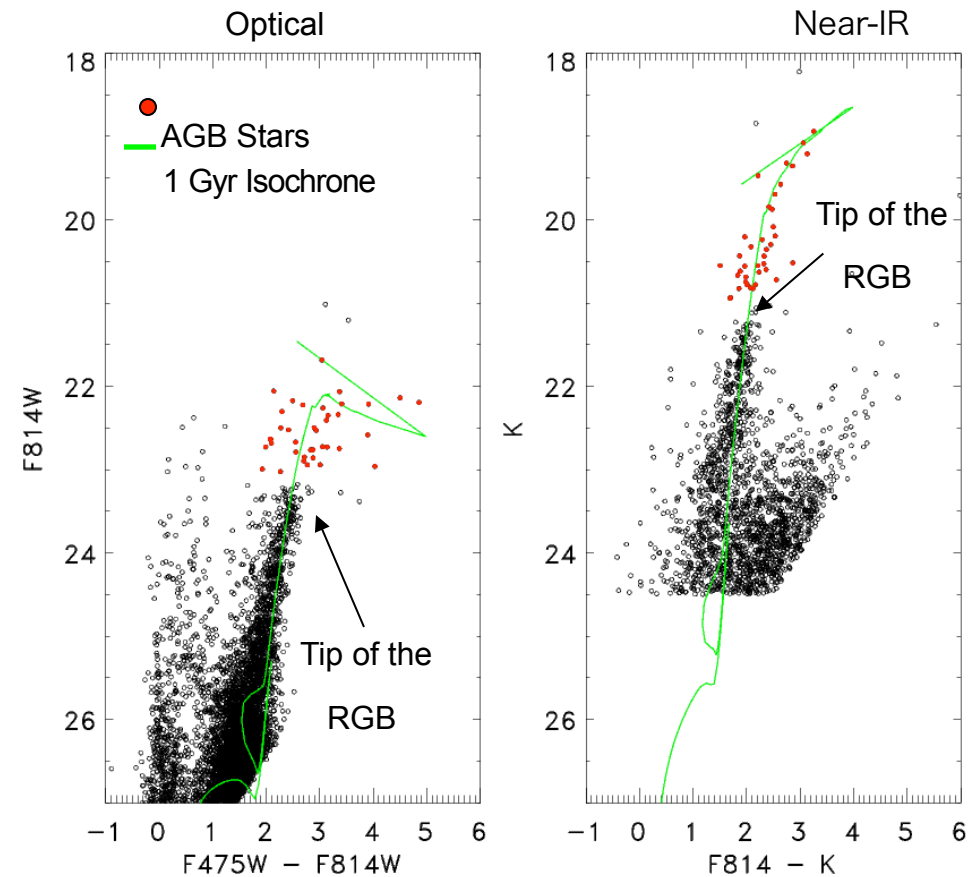
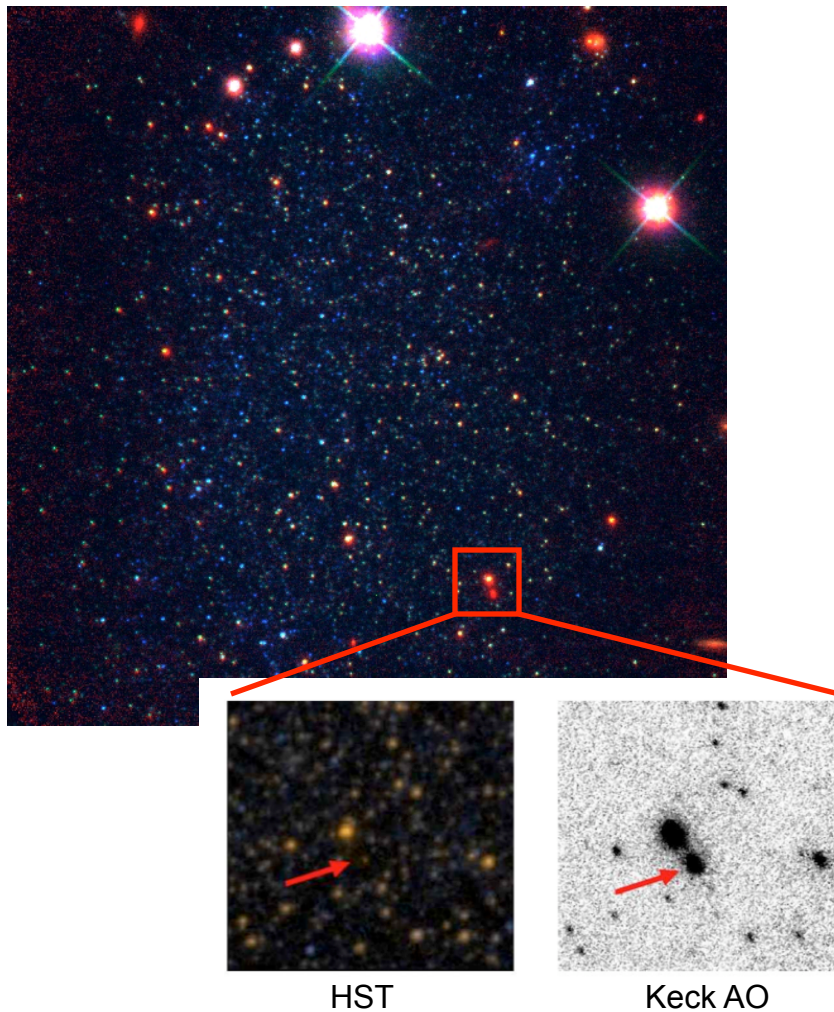


Williams et al. 2009

# The Colors and Magnitudes of Stars Can be Used to Reconstruct the Star Formation Histories of Galaxies

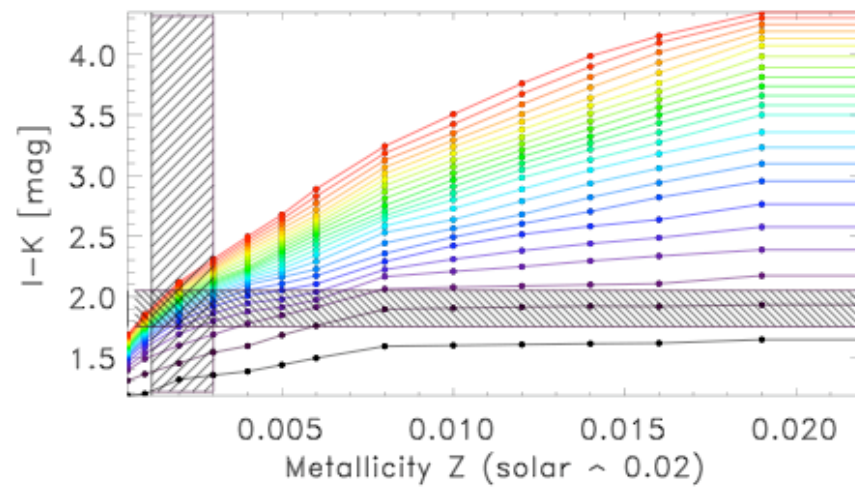
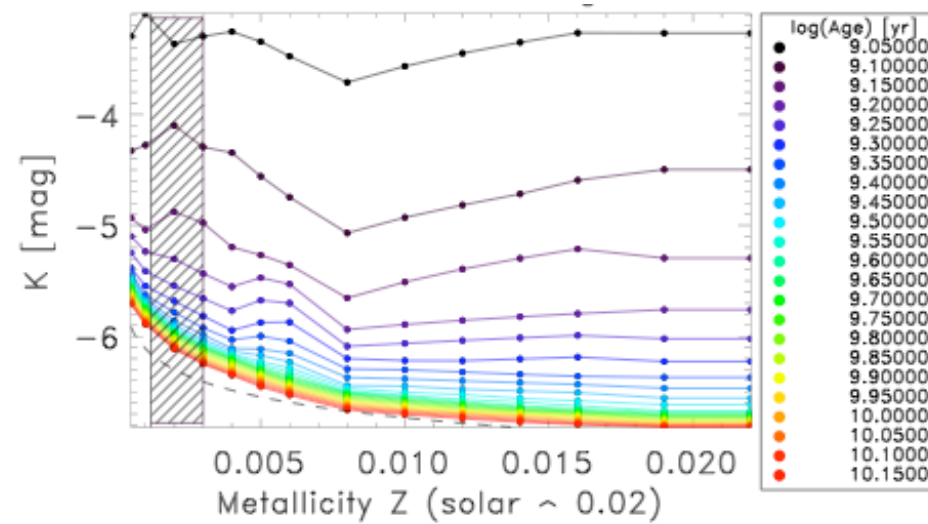


# In the Near-IR, Constraints on Star Formation History Come From the Asymptotic Giant Branch (AGB) and the Tip of the Red Giant Branch (TRGB)



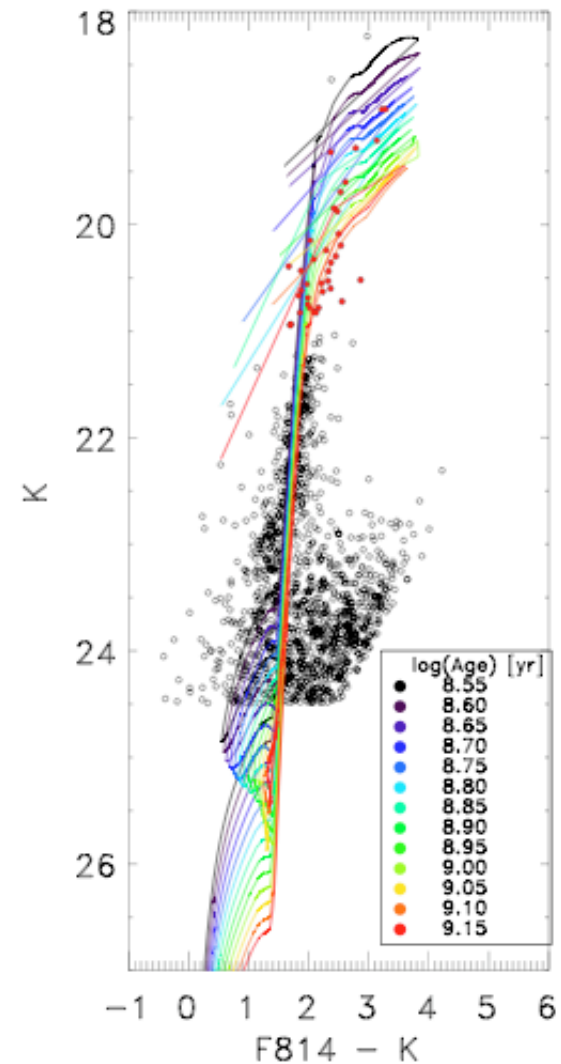
Melbourne et al. in Preparation

# The Color and Magnitude of the TRGB Indicates the Age of the Onset of Star Formation

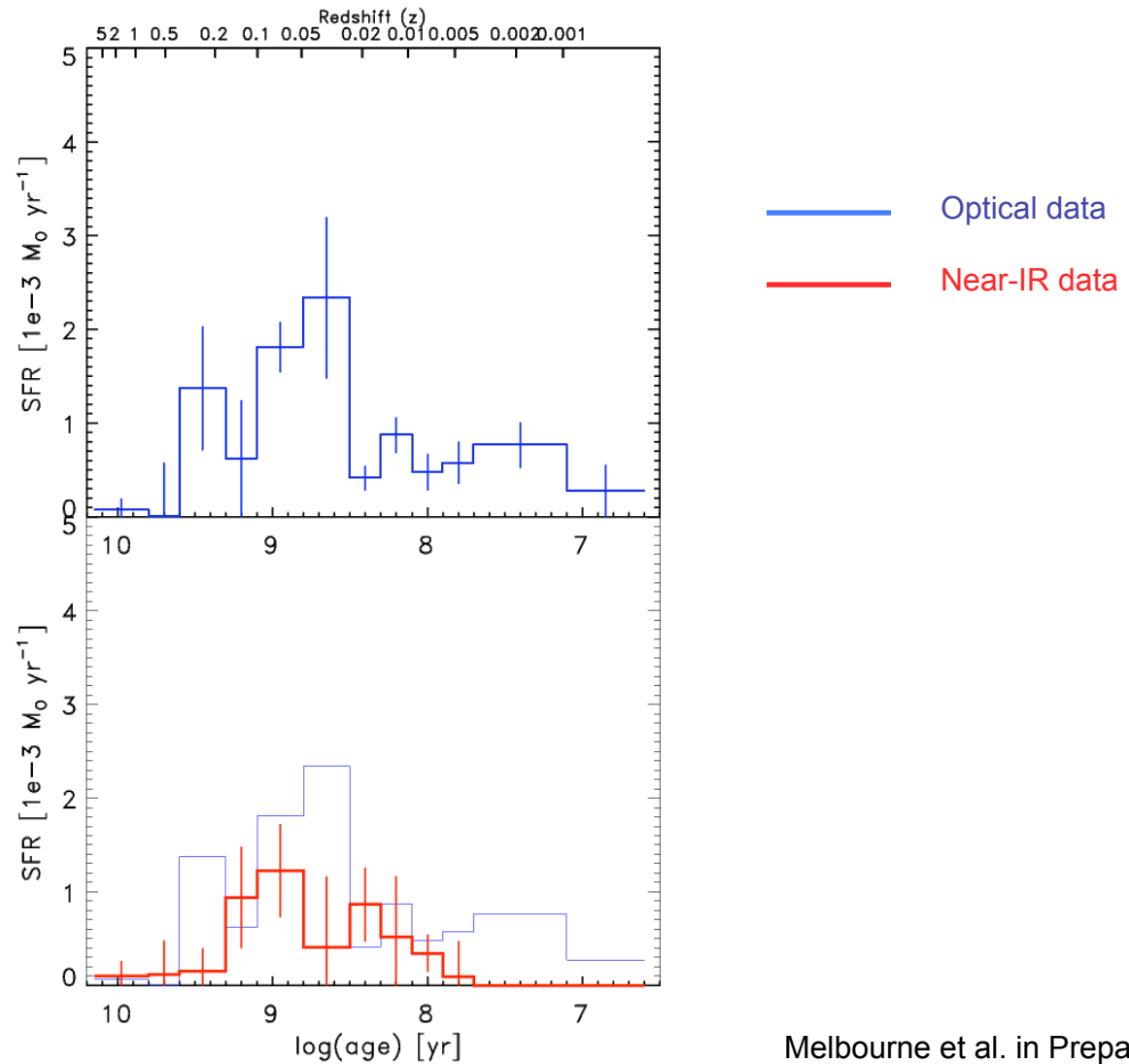


## The AGB is a Great Indicator of the Star Formation History from 0.5 - 5 Gyrs

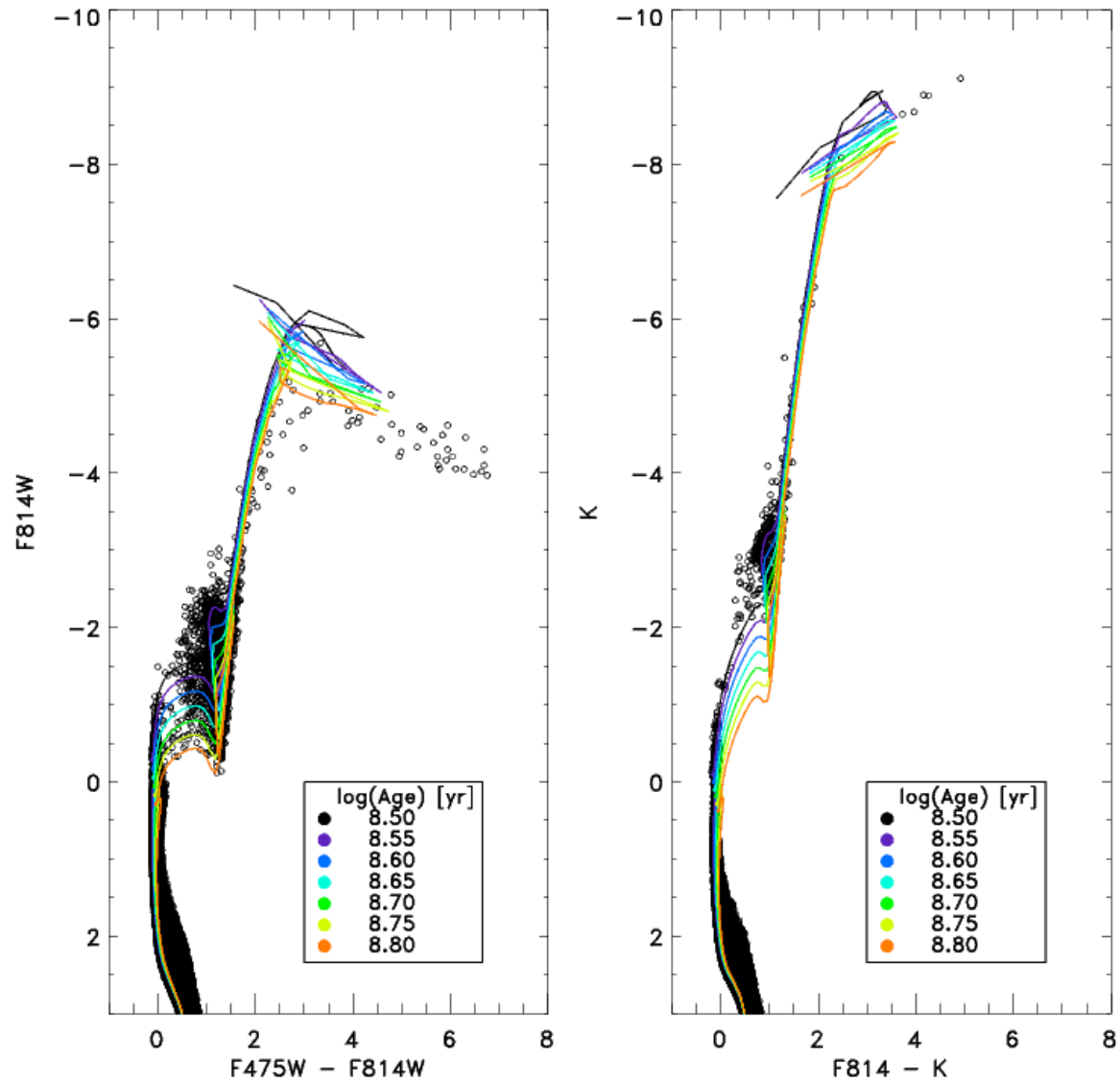
- Very Luminous in the Near-IR.
- Spatially sparse.
- Lie in Unique Sequences in Near-IR CMDs



In KKH 98, the SFH Deduced from the Optical Data are a Good Match the SFH from the IR



# Differences in SFH from Optical to IR Show Problem in Models of AGB





## AGB Stars Have Huge Implications for High-z Galaxies

- Stellar masses are based on SED fitting especially at redder wavelengths (less affected by dust).
- AGB stars contribute negligible mass but significant luminosity.
- At the redder wavelengths, AGB in  $\sim 1$  Gyr populations contribute 50% or more of the flux.
- If AGB not accounted for properly can get masses wrong by factors of two or more.
- Currently, best AGB models based on LMC, not a good fit to other galaxies.

## Our Plans for Studying the AGB in Nearby Galaxies

- Keck AO of ANGST Dwarf Galaxies (J, K) - data in hand
- Keck AO of M 33 Clusters (J, K) - 2 nights in Fall 09
- WFC3 Snapshot ANGST galaxies (H) - proposal approved
- ADP Archival Spitzer Imaging of ANGST galaxies - Pending.

# Summary

- Resolved photometry of stars can reveal star formation histories of nearby galaxies.
- This work can be done in the near-IR, where the IR bright stars are sufficiently sparse to resolve with Keck AO beyond 4 Mpc.
- The TRGB and AGB populations are important guide posts for stellar pops in the IR.
- To understand the high- $z$  universe, and to take full advantage of JWST and AO on TMT, the AGB needs to be better constrained observationally across a wide range of star formation histories and metallicities.