

Global Radio η Aquarids 2013

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Dust Trail of Eta Aquariids in 2013

Date: Sat, 04 May 2013 14:15:42 +0900

From: Mikiya Sato <miya@kaiyohi.net>

To: Meteor science and meteor observing <meteorobs@meteorobs.org>

Subject: (meteorobs) **Dust Trail of Eta Aquariids in 2013**

I found out that the old dust trails of Eta Aquariids (ETA) will approach the earth in 2013.

Dear all,

Date(UT) May.06 05:45 - May.06 21:19

I am Mikiya Sato.

I found out that the old dust trails of Eta Aquariids (ETA) will approach the earth in 2013.

The peak may be continuous or broader because distribution of the old dust has spread.

The outline is as follows.

I expect that they are about 2 times of the usual activity at the maximum because this case is similar to Orionids from 2006 to 2010.

Time	Date	UT	Mag	Rate	Mag	Rate
811	May.06	05:45	45.682	+0.0018	-2.12	0.095
-910	May.06	06:27	45.710	-0.0017	-2.11	0.017
-1197	May.06	12:37	45.959	+0.0021	+3.44	0.013

Detection of the increase is not easy since observation condition of ETA is not so good in the Northern Hemisphere.

In addition, -1403, -1333, -1265, -1128, -985, and -835 trail also tend to approach the earth.

The **peak may be continuous or broader** because distribution of the old dust has spread.

But, the increase in the number of meteors is unknown about ETA.

I expect that they are about 2 times of the usual activity at the maximum because this case is similar to Orionids from 2006 to 2010.

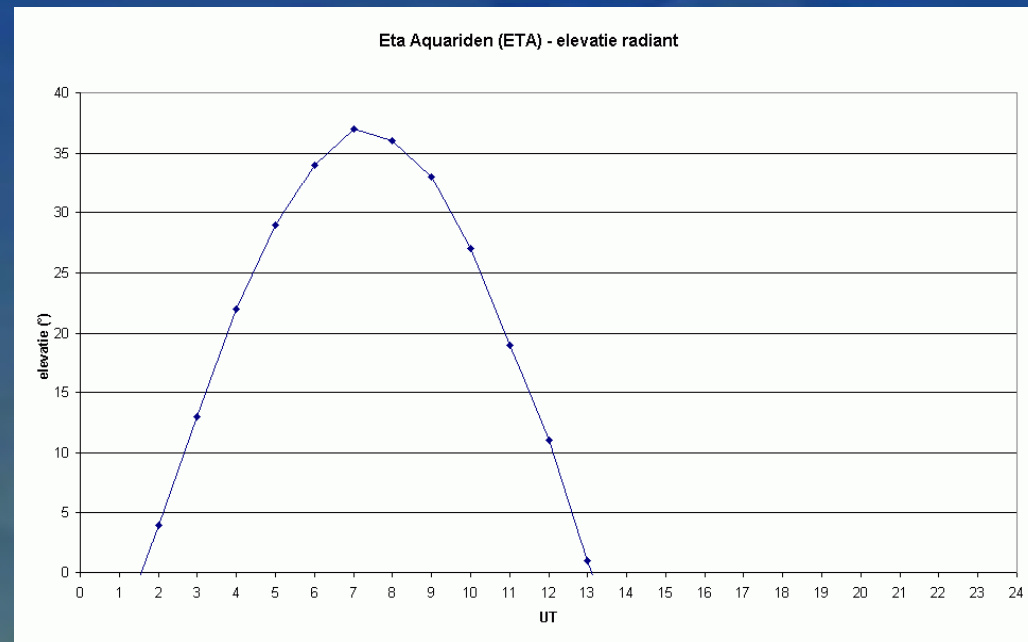
Detection of the increase is not easy since observation condition of ETA is not so good in the Northern Hemisphere.

However, please take notice of the appearance this year.

η Aquarids stream

- Radiant $\alpha = 336^\circ$ $\delta = -2^\circ$
- Daily path for $\lambda = 5^\circ$, $\beta = 51^\circ$

- Fast (66 km/s) \rightarrow less interesting for radio work
- Radio observations of previous years



'Global' hourly counts

Radio Meteor Observatory's On Line - Google Maps - Windows Internet Explorer

http://www.rmob.org/livedata/main.php

File Edit View Favorites Tools Help


★ Favorites Radio Meteor Observatory's On Line - Google Maps

Radio Meteor Observatory's On Line

[About Radio Meteor Observatory's DATA](#)

Click on red points onto this World Map to see observer data and use left zoom function for navigate into the map

[Google Maps observer coordinates HELP](#)



POWERED BY Google

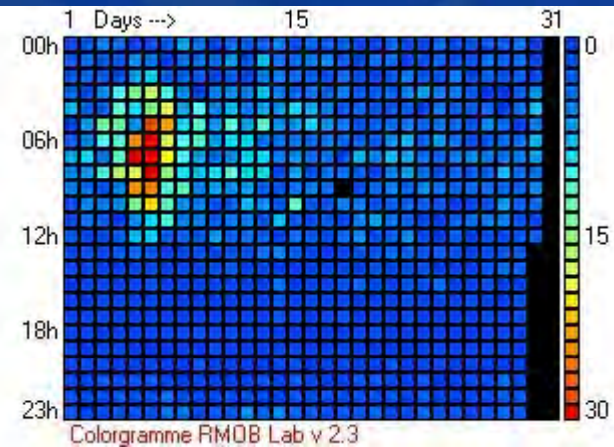
Server Powered by **INFOSAT** under **Linux**

Please see new version available for Colorgramme Lab on this site here : [Download](#)

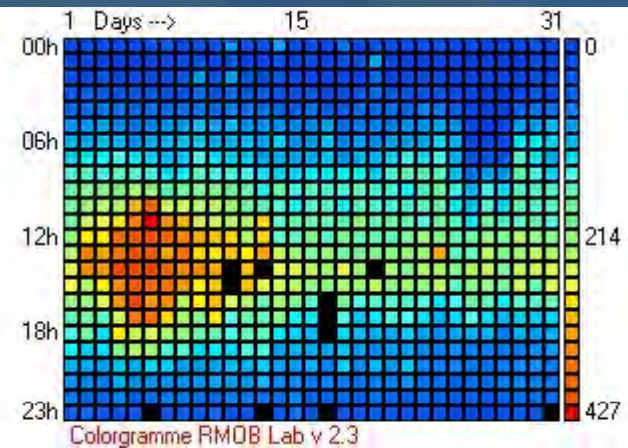
Internet 100%

May 2013

Observer : Felix Verbelen Location : 004°3539 East
Country : Belgium 050°5701 North
City : Kampenhout Frequency : 49.99 MHz
Antenna : 2-elements HB9CV Yagi Az. : 250° El. : 52°
RF Preamp : none
Receiver : ICOM-R75
Computer : PC-Pentium III

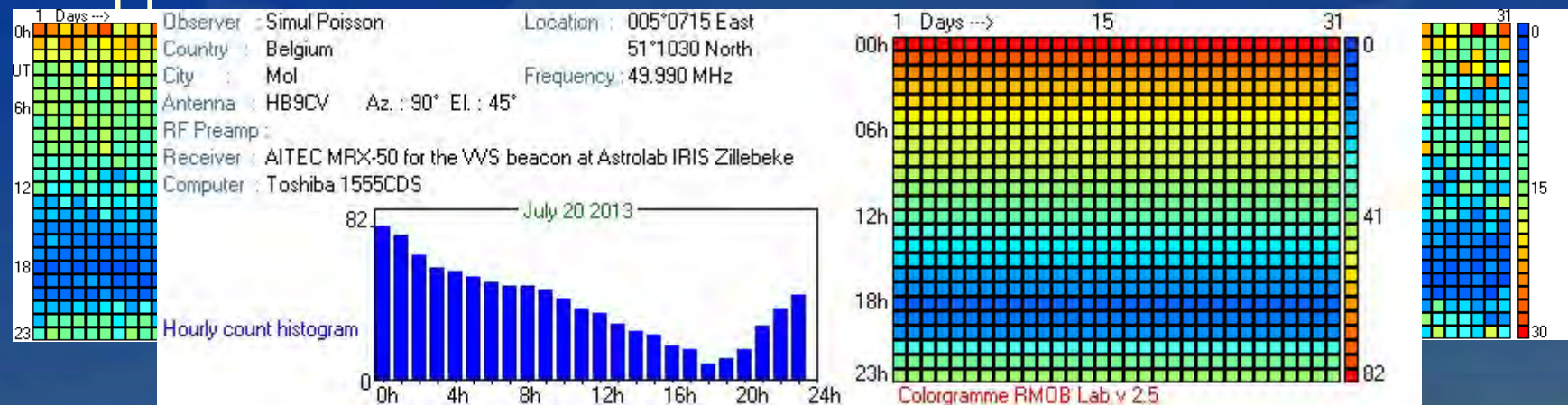


Observer : Mikhail Svoiski Location : 111°0000 West
Country : United States of America 033°0000 North
City : Tempe Frequency : 67.240
Antenna : Winegard GS-2200 active dipole Az. : 0° El. : 0°
RF Preamp :
Receiver : K89YIG Softrock VHF Ensemble
Computer : Compaq Armada M700

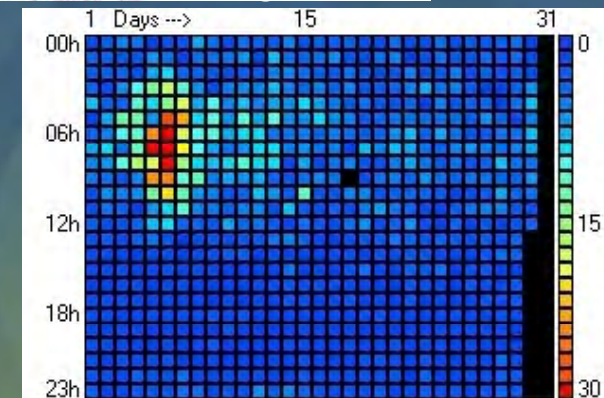


Patterns in function of counts

- number of 'random' events → Poisson distribution applies

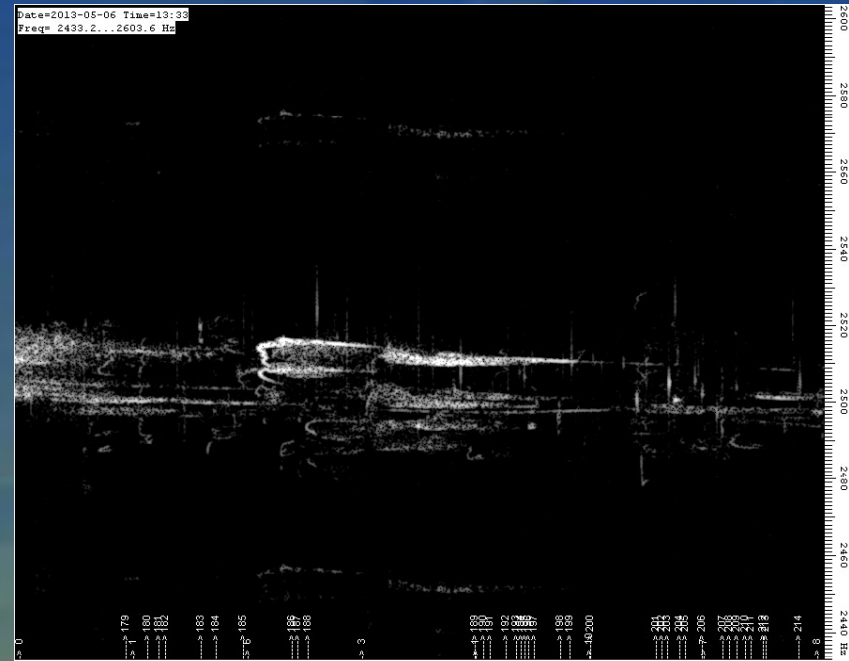
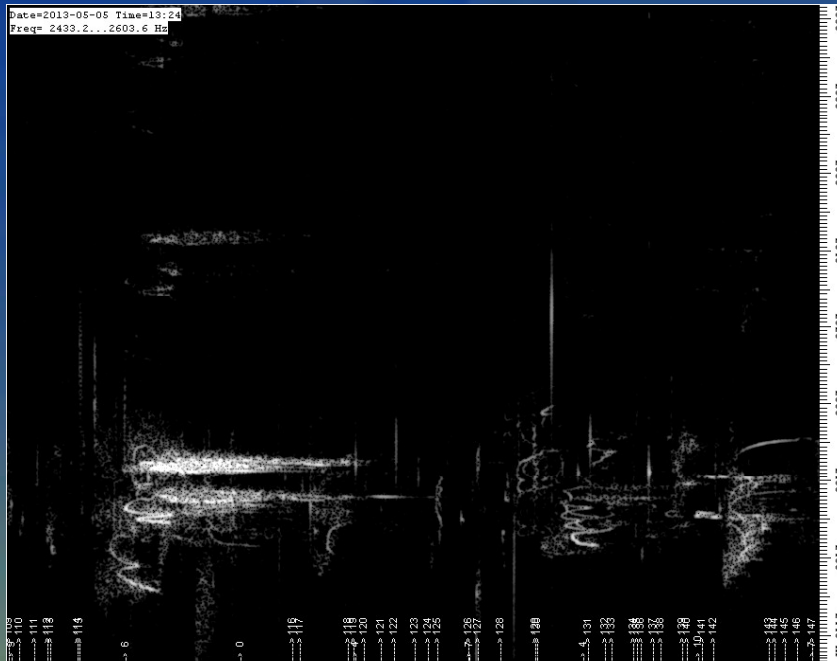


Verbelen May 2013



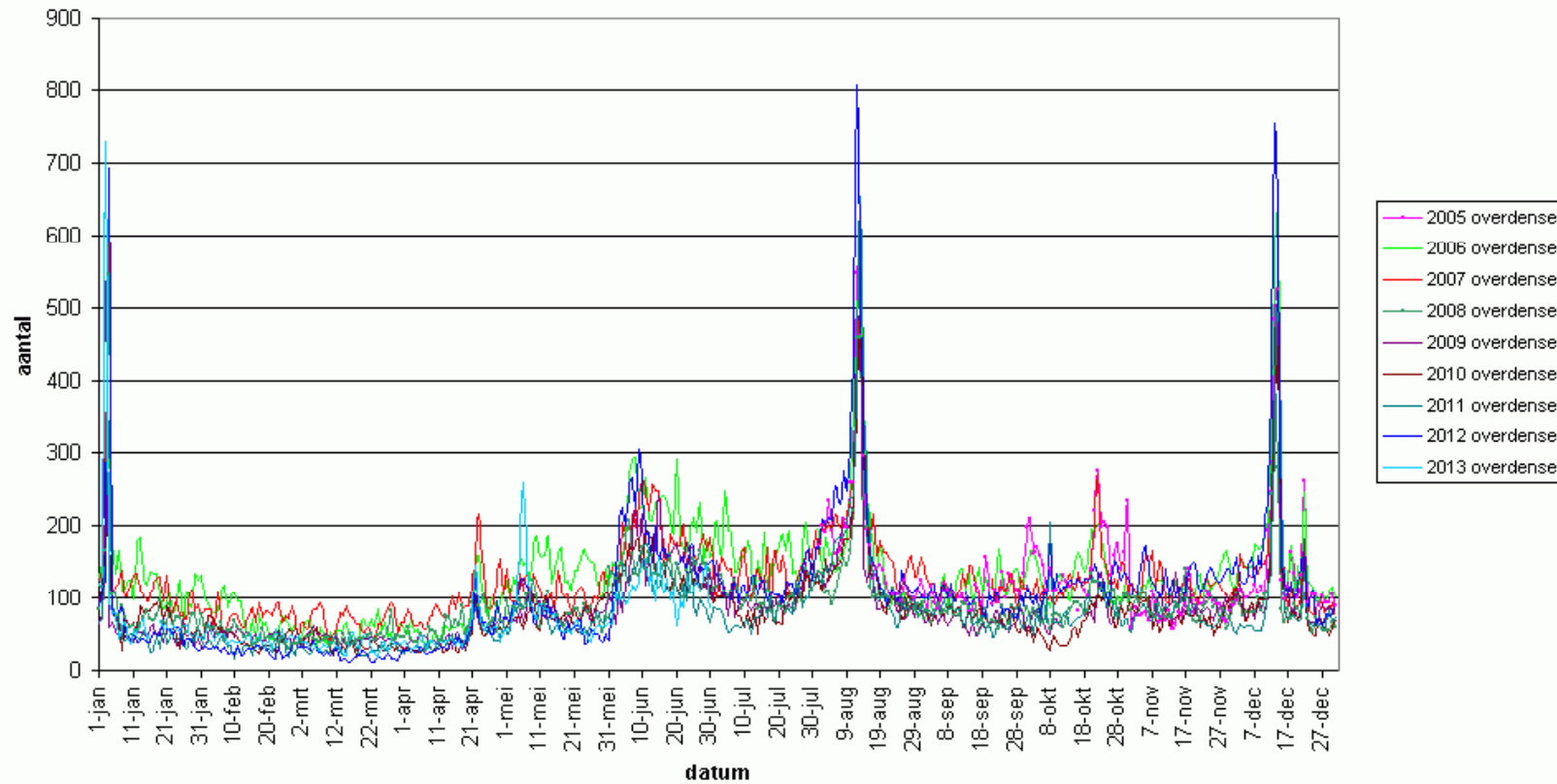
Details May 5 - 6

Svoiski
Automatic count?



Stability of radio counts

49.99 MHz - Radiometeoren - dagtotalen
overdense reflecties
Felix Verbelen (Kamphenhout)



Modelling stream activity

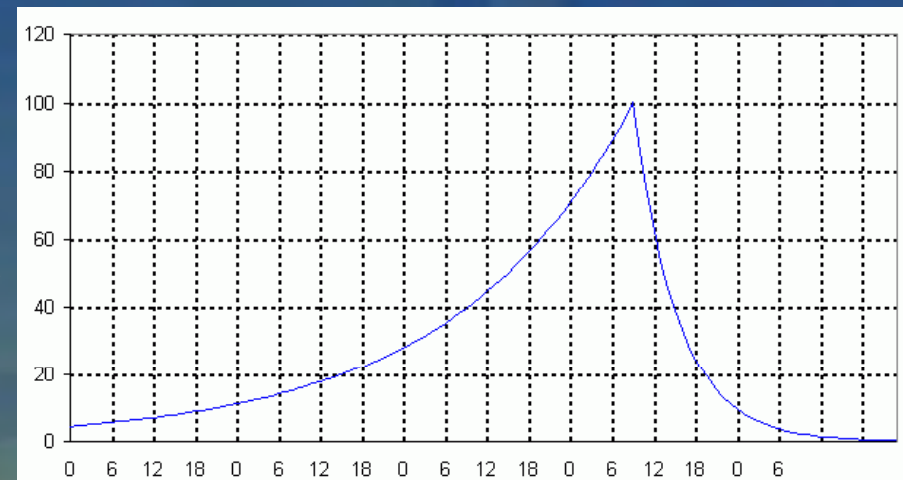
$$O(t) = S(T) + Z(t)OF(T)$$

$$T = \frac{t - t_0}{D}$$

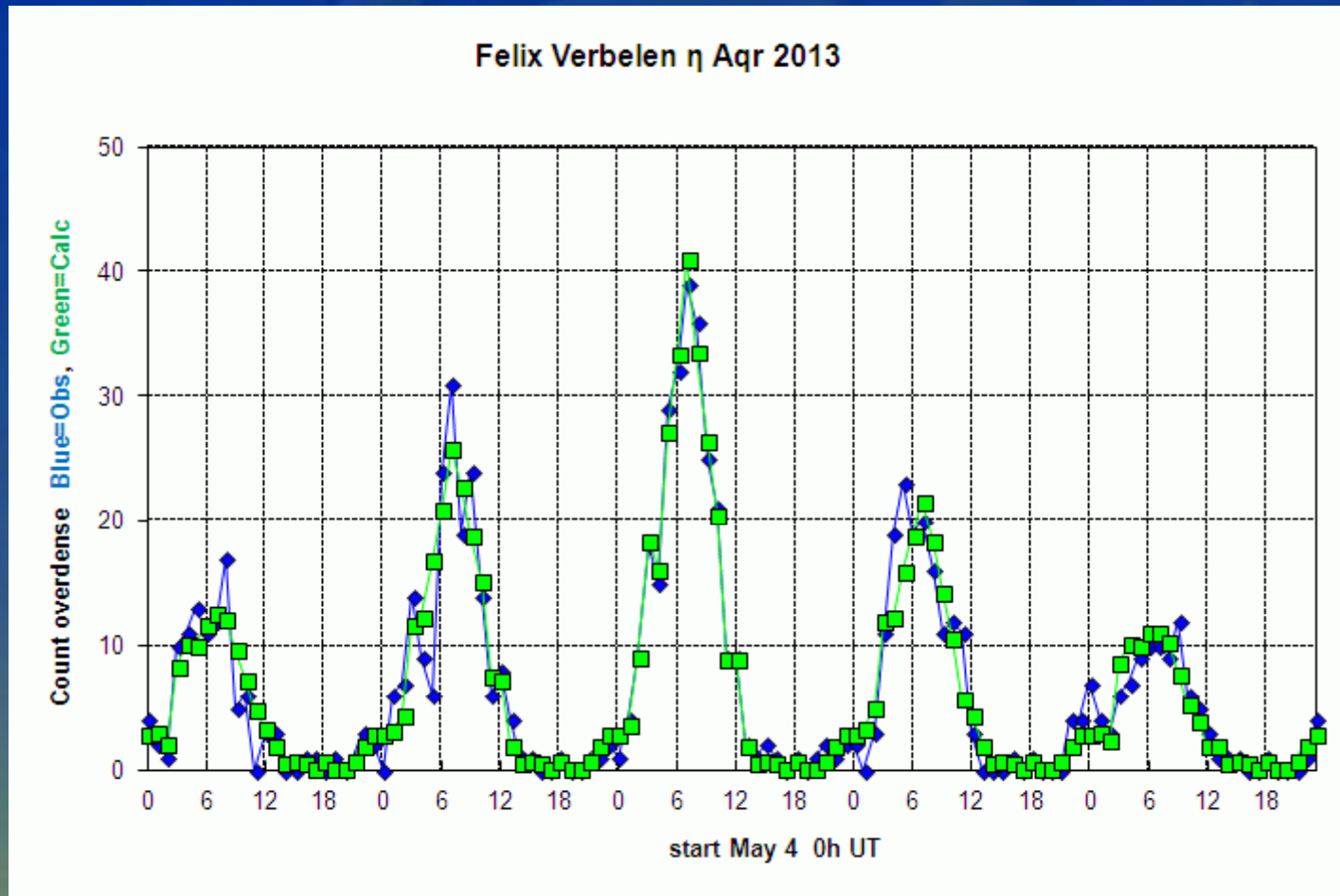
$$Z(t) = e^{-(t-t_M)/a}$$

$$Z(t) = e^{-(t_M-t)/b}$$

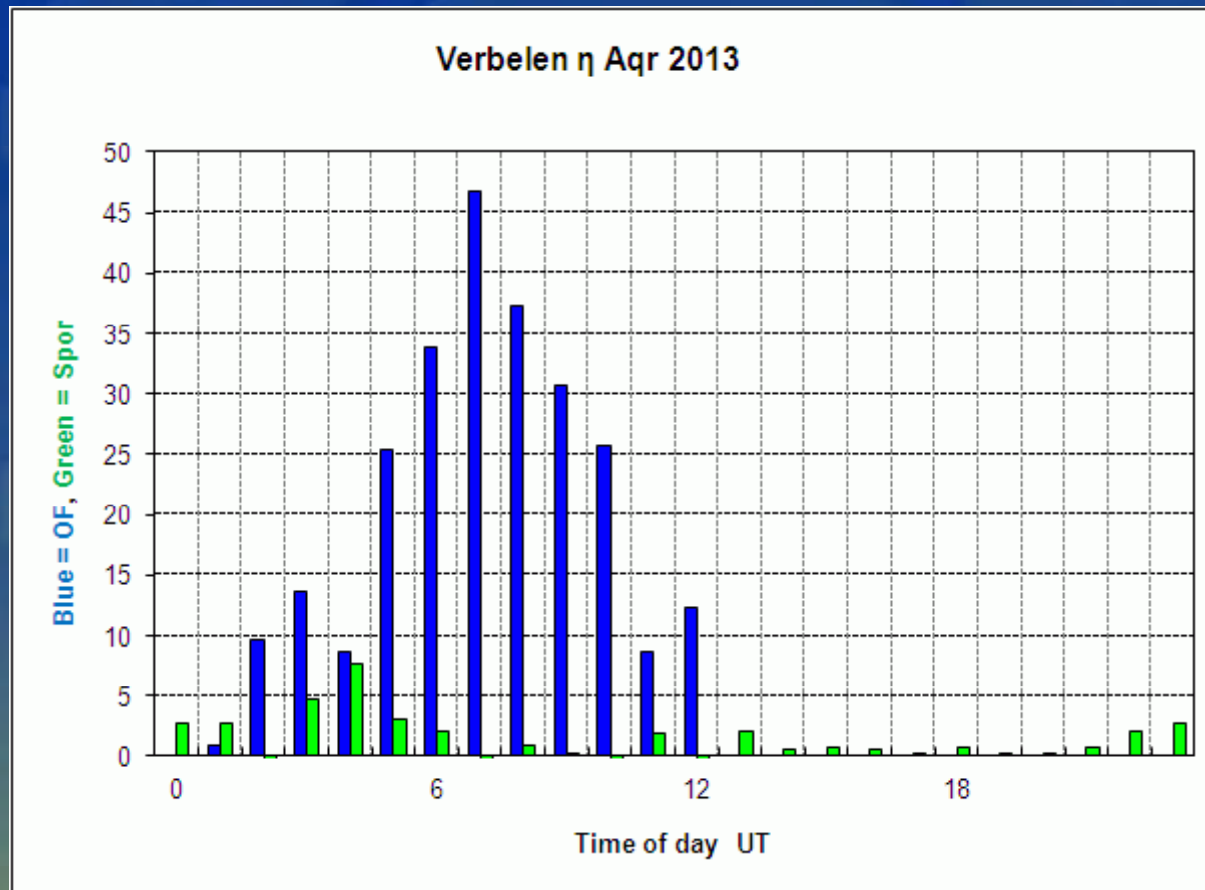
O observed 'activity'
 S sporadic background
 Z stream profile
 OF Observability Function
 t_M instance of maximum
 a rise time constant
 b decay



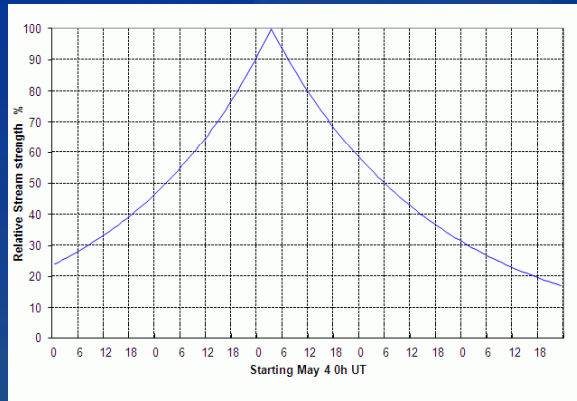
Applied to Verbelen η Aqr 2013



Applied to Verbelen η Aqr 2013



Maximum location

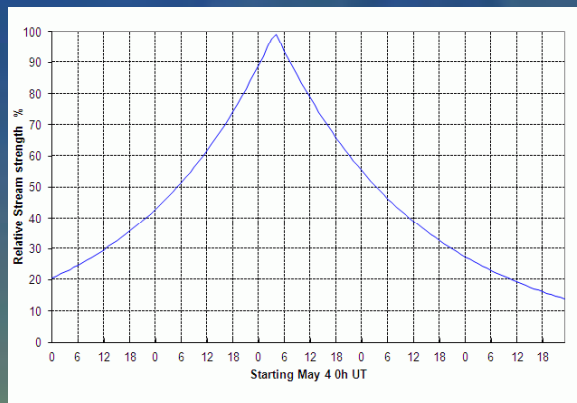


Verbelen:

$$t_M = \text{May 6, 3.4 h UT} \pm 4\text{h}$$

$$a = 36 \pm 7 \text{ h}, b = 39 \pm 9 \text{ h (asymmetric?)}$$

Observed maximum May 6, 7 h UT



Svoiski:

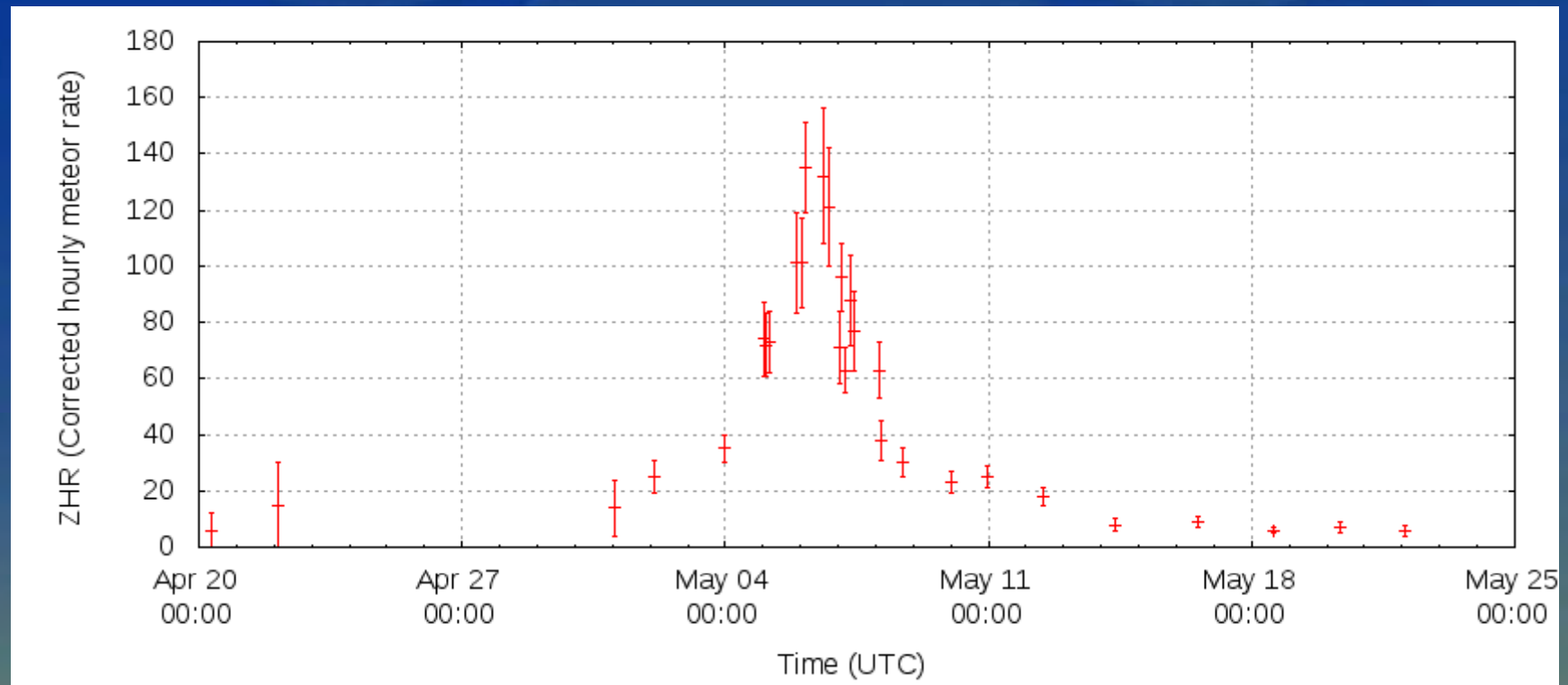
$$t_M = \text{May 6, 4.0} \pm 2.5 \text{ h UT}$$

$$a = 33 \pm 8 \text{ h}, b = 34 \pm 10 \text{ h}$$

Observed maximum May 6, 11 h UT

Visual comparison

<http://www.imo.net/live/eta-aquariids2013/>



Visual comparison

<http://www.imo.net/live/eta-aquariids2013/>

Time (UTC)	Solarlon	nINT	nETA	ZHR		Particle density
2013-04-20 07:52	30.215	2	0	6	±6	11 / 10 ⁹ ·km ³
2013-04-22 02:10	31.934	2	0	15	±15	27 / 10 ⁹ ·km ³
2013-05-01 01:58	40.680	1	1	14	±10	25 / 10 ⁹ ·km ³
2013-05-02 02:11	41.660	5	19	25	±6	46 / 10 ⁹ ·km ³
2013-05-03 23:27	43.489	10	40	35	±5	64 / 10 ⁹ ·km ³
2013-05-05 00:15	44.490	9	31	74	±13	135 / 10 ⁹ ·km ³
2013-05-05 01:36	44.545	6	43	72	±11	131 / 10 ⁹ ·km ³
2013-05-05 04:16	44.652	3	42	73	±11	133 / 10 ⁹ ·km ³
2013-05-05 21:56	45.366	6	31	101	±18	184 / 10 ⁹ ·km ³
2013-05-06 01:05	45.493	10	41	101	±16	184 / 10 ⁹ ·km ³
2013-05-06 02:45	45.560	6	67	135	±16	246 / 10 ⁹ ·km ³
2013-05-06 15:06	46.059	4	30	132	±24	240 / 10 ⁹ ·km ³
2013-05-06 18:11	46.183	3	32	121	±21	220 / 10 ⁹ ·km ³
2013-05-07 01:06	46.462	11	30	71	±13	129 / 10 ⁹ ·km ³
2013-05-07 02:20	46.512	8	68	96	±12	175 / 10 ⁹ ·km ³
2013-05-07 04:29	46.598	12	61	63	±8	115 / 10 ⁹ ·km ³
2013-05-07 08:28	46.759	4	30	88	±16	160 / 10 ⁹ ·km ³
2013-05-07 10:19	46.833	3	31	77	±14	140 / 10 ⁹ ·km ³
2013-05-08 01:50	47.459	5	36	63	±10	115 / 10 ⁹ ·km ³
2013-05-08 03:48	47.539	4	32	38	±7	69 / 10 ⁹ ·km ³
2013-05-08 16:51	48.065	5	31	30	±5	55 / 10 ⁹ ·km ³
2013-05-09 23:45	49.311	11	30	23	±4	42 / 10 ⁹ ·km ³

Eta Aquariids in Japan (May 6)

Date: Wed, 08 May 2013 19:14:43 +0900

From: Mikiya Sato <mail@kaicho.net>

To: Meteor science and meteor observing <meteorobs@meteorobs.org>

Subject: (meteorobs) Eta Aquariids in Japan (May 6)

From IMO web, it seems that two peaks of ETA were observed. The first peak time was about 3h on May 6. Probably, this was formed by -910 trail.

Dear all,

Similarly, the 2nd peak about 15h was caused by -1197 trail.

I am Mikiya Sato.

I could also watch enhanced Eta Aquariids, yesterday. Many bright meteors of ETA were observed even in Japan.

However, these two peaks might be continuous.

May 6, 2013

And, the outburst lasted about 2 days or more. This may mean that distribution of dust had spread because two or more kind of dust trails approached.

Time (UT)	3. Long	Perf	Lim	ETA	Alt	ZHR(1-20°)	Obs.	
1725-1755	46.17	0.50	1.0	5.90	10	20	120	M. Sato
1800-1830	46.20	0.50	1.0	5.90	15	30	129	T. Sato
1830-1850	46.21	0.33	1.0	5.70	7	21	90	M. Sato

*Location : 138d44.8'E 35d26.5'N 1200m Yamanashi, Japan.

>From IMO web, it seems that two peaks of ETA were observed. The first peak time was about 3h on May 6. Probably, this was formed by -910 trail. Similarly, the 2nd peak about 15h was caused by -1197 trail. However, these two peaks might be continuous.

And, the outburst lasted about 2 days or more. This may mean that distribution of dust had spread because two or more kind of dust trails approached.

I have to examine it in more detail in the future.

Thank you very much to all observers!

Thanks to / acknowledgments

- Felix Verbelen
- Micha Svoiski
- Pierre Terrier
- David Entwistle
- (meteorobs) mailing list
- Radio Meteor Observatories On Line
- IMO
- Astrolab IRIS, Zillebeke
- VVS