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1.

### LIGHTCURVES AND PERIOD DETERMINATION FOR **6146 ADAMKRAFFT**

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Minor planet 6146 Adamkrafft was observed over a period of 13 days (41 rotations) during August, 2001. Lightcurves obtained on 6 nights with an unfiltered CCD have yielded a rotational (synodic) period of 6.964 ± 0.009 hrs. The complete lightcurve is doubly periodic, with a total amplitude range of  $0.18 \pm 0.05$  magnitude. The rotation period is atypical of C and S class asteroids of this diameter.

### Introduction

Adamkrafft was discovered by C.J. van Houten, I. van Houten-Groeneveld and T. Gehrels at Mt Palomar on Sept 30, 1973 as a result of the Palomar-Leiden Trojan - 2 asteroid survey. Previous provisional designations were 1934 NG, 1979FA and 3262 T-2. This main-belt asteroid was subsequently named after the German sculptor Adam Krafft, 1460-1508. The name was cited in Minor Planet Circular No. 26765. The H magnitude is 13.0 and its assumed diameter is given as 16.6 km.

This asteroid was chosen from the suggested targets in the observing list of Pravec and Harris (2001). The most recent tabulation examined for previous lightcurve data was that of Harris (2001), but no lightcurve data are listed there.

### Observations and Results

At this opposition Adamkrafft was at Declination -23°, making it particularly suitable for southern hemisphere observers. Observations were made from Mt Tarana Observatory near Bathurst, NSW. The site is at 880m altitude and the latitude is S 33.4348, longitude E 149.7576. The equipment and methodology have been described by Bembrick (2001).

The observational circumstances for August, 2001 have been summarised in Table I, which also shows the percent of the rotational light curve covered on each night of observation. The data from each night were plotted as differential instrumental magnitude vs U.T. No light time corrections were applied. Data

from the better nights show two maxima and two minima, with the least noisy light curve on 25 August. Thus the epoch of zero phase was chosen as the prominent maximum of Aug. 25 at 13.85 UT (JDGeo 2452147.07708).

In all, 18 extrema were identified from the lightcurves and the time differences between these were used to estimate the rotational period. This led to an estimate of  $6.964 \pm 0.009$  hours. On this basis the observations cover 41 rotational cycles. Using the above epoch and period, the data were phase folded (Figure 1), with the magnitudes of the primary maxima from other nights being adjusted to the magnitude of the zero phase maximum on the night of epoch. Results were posted on the CALL website on Oct. 14, 2001.

As the composite lightcurve is somewhat noisy, a smoothed, phase-folded plot (Figure 1) was produced by the method of 'phase binning", using 32 overlapping bins. This smoothed curve is displaced 0.2 magnitudes from the composite curve in the figure. From this smoothed curve the average amplitudes of the extrema were estimated. The overall amplitude of this lightcurve is 0.18 magnitudes. The one-sigma error bars for these phase bins are generally 0.05 magnitudes or less, giving an indication of the overall accuracy of the quoted amplitudes.

### Discussion and Conclusion

Using AAVSO variable star software, a period search was made using data from all six nights. The results agree well with the graphical method outlined above. The best frequency was found to be 3.448 (cycles/day), which is a period of 6.96 hours. Tests with other software also found a frequency of 2.450 (period of 9.79 hours), however phase stacking using this period did not produce an acceptable composite lightcurve. Thus this period was rejected as an alias. To a first approximation, the maximum amplitude of 0.18 magnitudes implies a ratio a/b of 0.85, where a, b and c are the axes of a tri-axial ellipsoid and the rotation is about the shortest axis, c. The small difference of around 0.04 magnitudes between the two maxima implies little difference in albedo or reflecting surface area between the opposite hemispheres.

Minor planet 6146 Adamkrafft was observed over 41 rotational cycles and the synodic period determined as  $6.964 \pm 0.009$  hours. This is believed to be a secure result as all rotational phases of the lightcurve were observed. The lightcurve shows the doubly periodic features typical of an irregularly shaped, tri-axial ellipsoid.

With a revolution rate of 3.44/day, Adamkrafft is one of the faster rotators among the asteroids less than 50 km diameter. This group

2 of asteroids may have a bimodal distribution of rotation rates peaking at slightly less than one rev/day and at approximately 3 revs/day (Binzel et al. 1989). The deduced revolution rate is not typical of C or S class asteroids of this diameter (Burns and Tedesco, 1979).

#### References

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Table I. Observational Circumstances, 2001

Obs. Date	Hel. Lat. (B°)	Hel. Long. $(L^{\circ})$	R (AU)	Solar Phase Angle	Phase Coverage
13 Aug.	-8.14	316.32	1.731	5.4	28%
14 Aug.	-8.12	316.15	1.732	5.8	50
16 Aug.	-8.06	315.79	1.735	6.6	71
23 Aug.	-7.78	314.68	1.744	10.1	107
24 Aug.	-7.73	314.55	1.746	10.6	121
25 Aug.	-7.68	314.42	1.747	11.1	107

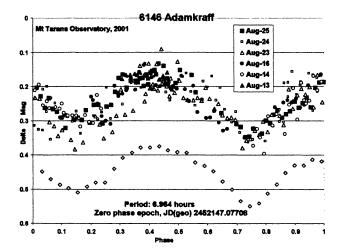


Figure 1. Composite Lightcurve for 6146 Adamkrafft.

### PHOTOMETRY OF 866 FATME, 894 ERDA, 1108 DEMETER, AND 3443 LETSUNGDAO

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Minor planets 866 Fatme, 894 Erda, 1108 Demeter, and 3443 Letsungdao were observed from Santana Observatory (MPC Code 646). The rotational period of 866 Fatme was determined to be  $20.03 \pm 0.01$  hours with an amplitude of  $0.28 \pm 0.05$  magnitude. 894 Erda was determined to have a rotational period of  $4.69 \pm 0.01$  hours with an amplitude of  $0.08 \pm 0.05$  magnitude. The rotational period of 1108 Demeter was determined to be  $9.70 \pm 0.01$  hours with an amplitude of  $0.17 \pm 0.01$  magnitude. Finally, 3443 Letsungdao was found to have a rotational period of  $3.44 \pm 0.02$  hours with an amplitude of  $0.30 \pm 0.04$  magnitude.

Santana Observatory is located in Rancho Cucamonga, California at an elevation of 400 meters and is operated by Robert D. Stephens. Details of the equipment used can be found in Stephens (2000). All of the asteroids whose results are presented here were selected from the CALL web site "List of Potential Lightcurve Targets" (Warner 2000).

Aperture photometry was performed using the software program Canopus developed by Brian Warner and including the Fourier analysis routine developed by Alan Harris (Harris et al., 1989). This program allows combining data from different observers and adjusting the zero points to compensate for different equipment and comparison stars. All observations were unfiltered. Dark frames and flat fields were used to calibrate the images.

Fatme is a main-belt asteroid discovered 25 February 1917 by M. Wolf at Heidelberg. It is named after a character from Abu Hassam by the German composer Carl Maria Friedrich Ernst von Weber. Three hundred five observations over 11 sessions between 14 May and 8 June 2001 used to derive the rotational period of  $20.03 \pm 0.01$  hours with an amplitude of  $0.28 \pm 0.05$  magnitude.

Erda is a main-belt asteroid discovered 4 June 1918 by M. Wolf at Heidelberg. It is named for the Norse goddess. Two hundred six observations over five sessions between 15 July and 19 July 2001 were used to derive the rotational period of  $4.69 \pm 0.01$  hours with an amplitude of  $0.08 \pm 0.05$  magnitude. Thirty observations by Laurent Bernasconi on 24/25 July 2001 as analyzed by Raoul Behred although which cannot independently determine the period, are consistent with this period.

Demeter is a main-belt asteroid discovered 31 May 1929 by K. Reinmuth at Heidelberg. It is named for the Greek goddess of the fruitful soil and of agriculture. One hundred fifty one observations over seven sessions 9 July and 29 July 2001 were used to derive the rotational period of 9.70  $\pm$  0.01 hours with an amplitude of 0.17  $\pm$  0.02 magnitude. Observations by René Roy and Stéphane Charonnel taken on 25 June 2001 and 3 July 2001 as analyzed by Raoul Behred, although which cannot independently determine the period, are consistent with this data.

Letsungdao is a main-belt asteroid discovered by Purple Mountain Observatory at Nanking on 26 September 1979. It is named in honor of the theoretical physicist Tsung-Dao Lee. One hundred fifty nine observations over three sessions spanning three days were used to derive the rotational period of  $3.44 \pm .02$  hours with an amplitude of  $0.30 \pm 0.04$  magnitude. Observations by Laurent Bernasconi and Stéphane Charonnel taken on 22 July 2001, 23 July 2001, and 31 July 2001 as analyzed by Raoul Behred derived a rotational period of  $3.438 \pm 0.0005$  hours.

### Acknowledgements

Many thanks to Brian Warner for his continuing work and enhancements to the software program "Canopus" which makes it possible for amateur astronomers to analyze and collaborate on asteroid rotational period projects and for maintaining the CALL Web site which helps coordinate collaborative projects between amateur astronomers.

### References

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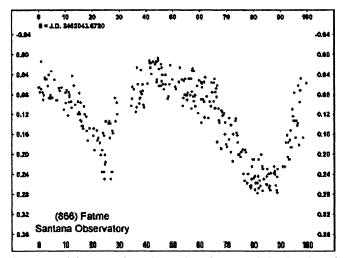


Figure 1: Lightcurve of 866 Fatme based upon a derived period of  $20.03 \pm 0.01$  hours.

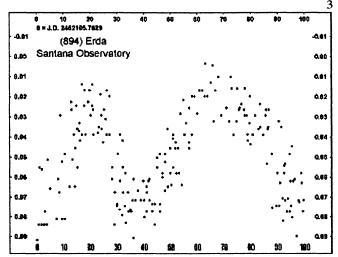


Figure 2: Lightcurve of 894 Erda based upon a derived period of  $4.69 \pm 0.01$  hours.

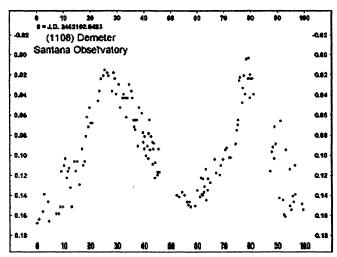


Figure 3: Lightcurve of 1108 Demeter based upon a derived period of  $9.70 \pm 0.01$  hours.

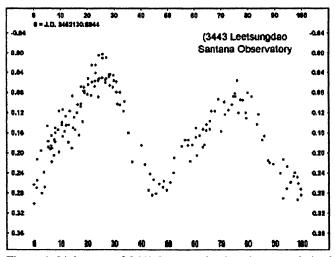


Figure 4: Lightcurve of 3443 Leetsungdao based upon a derived period of  $3.44 \pm 0.02$  hours.

### CLOSE APPROACHES OF MINOR PLANETS TO NAKED EYE STARS IN 2002

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A list is presented of approaches of minor planets brighter than magnitude 14 to naked eye stars. This may be helpful in finding some faint minor planets.

The accompanying table lists close approaches of minor planets to stars during 2002 where:

- The event takes place more than 30° from the Sun.
- The minor planet is brighter than visual magnitude 14.
- The star is brighter than magnitude 6.
- The minimum angular separation is smaller than 120".

This list can be helpful in locating some otherwise faint minor planets. By carefully drawing the star field around the predicted position down to the magnitude of the minor planet and comparing it to the situation some time later, one can detect the intruder by its apparent motion.

The information contained in the list is divided into 5 groups:

1. Date: gives the date and time in U.T. of the closest geocentric approach. All subsequent data pertain to this instant.

- 2. Closest approach: the two columns give the position of the minor planet with respect to the star:
  - the minimum geocentric distance in seconds of arc
  - the position angle in degrees, measured from north over east
- 3. Minor planet: gives information about the minor planet:
  - number and name
  - · visual magnitude
  - · apparent motion in seconds of arc per hour
  - · parallax in seconds of arc
- 4. Star: the following data of the star are given:
  - · Hipparcos star number
  - visual magnitude
  - right ascension for the equinox 2000.0
  - declination (2000.0)
- 5. Sun and Moon:
  - elongation of the Sun in degrees
  - elongation of the Moon (degrees)
  - illuminated fraction of the Moon in %

The observed minimum distance depends on the location of the observer on the Earth's surface but is always comprised between the minimum geocentric distance plus and minus the parallax. An occultation will be visible somewhere on the Earth when the parallax is greater than the geocentric separation.

The close approaches in this article were computed at the Computer Center of Agfa-Gevaert N.V., Mortsel, Belgium.

							Clo	se approa	ches of	minor	plane	ts to stars			_			
							-			REMERSE				(Dist.	< 120 ;	El. Su	n > 5	0 5
														Star	< 6.0 ; N	dn. pl	. < 1	4.0)
	Dat	e	(U.T	•)	Minim. dist.	Pos.		inor ame		n e t App. mot.	Hor.	s Designation	t a Vis. mag.	r Right ascens. (2000.0)	Decli- nation (2000.0)	Elon- gati Sun		Ill. frac Moon
			h	m		0				•/h	-			h m	0 '	0	0	•
2001	dec 2	6	9	41.5	80.54	345		Asterope	13.3	61.73	3.25	HIP 110578	5.8	22 24.11	- 4 50.2	61	69	82
	dec 2	8	5	0.4	109.29	343		Chloris	13.4	73.17	2.94	HIP 105576	5.7	21 23.01	-22 40.1	40	111	94
	dec 3	-		32.8	86.39	339		Frigga		60.20		TYC1 1577	5.1	21 53.30	-13 33.1	46	136	100
2002		1		4.2		165		Cyane		30.58		HIP 25278	5.1		+17 23.0	159	47	94
	jan	7	23	18.3	47.37	221	127	Johanna	12.3	19.86	5.13	TYC1 371	3.9	9 52.76	+26 0.4	144	79	30
	jan 1	9	7	4.6	93.49	348	66	Maja	13.4	29.68	5.22	TYC1 1081	5.8	2 58.09	+20 40.1	109	47	28
	jan 2	0	23	47.3	92.95	356	88	Thisbe	11.8	26.53	3.95	TYC1 241	2.9	6 22.96	+22 30.8	154	73	43
	jan 2	1		1.9	27.37	180	312	Pierretta	13.9	75.11	2.78	HIP 89931	2.9	18 20.99	-29 49.7	27	112	46
	jan 2			21.7	49.11	172		Isolda		28.11		HIP 13702	5.7	2 56.44	+18 1.4	101	35	86
	jan 2	:5	16	38.5	74.53	340	241	Germania	13.3	55.12	2.69	TYC1 3908	4.9	23 46.39	+ 3 29.2	52	84	86
	jan 2			53.7	14.03	149	498	Tokio		55.44		TYC1 2121	5.9	1 48.43	+ 3 41.1	78	94	99
	jan 3	10		33.4	30.80	71	140	8iwa		2.67		HIP 21604	5.9		+20 41.1	120	80	97
	feb			40.4	20.48	168		Pretoria		55.86		HIP 89439	5.3	18 15.22	-20 43.7	40	63	62
	feb			51.3	55.28	191		Ara		27.82		HIP 45751	4.9	9 19.77	-11 58.5	151	74	57
	feb	3	22	8.2	117.01	182	135	Hertha	12.6	27.64	4.54	HIP 35025	5.9	7 14.70	+24 53.1	151	110	57
	feb 1	.0	1	5.9	48.62	208	1107	Lictoria	13.5	14.30	4.53	HIP 32968	5.8	6 52.00	+23 36.1	140	166	5
	feb 1	1	2	44.5	8.21	192	476	Hedwig	13.8	24.50	3.52	HIP 17954	5.3	3 50.32	+25 34.8	98	112	1
	feb 1	.7	7	43.1	35.50	340	54	Alexandra	13.2	59.47	2.76	HIP 3632	5.4	0 46.55	+15 28.5	49	15	21
	feb 2		3	5.8	7.03	201	12	Victoria	11.4	33.71	4.63	TYC1 1223	4.1		+ 5 42.2	155	73	46
	feb 2	4	5	4.2	29.78	173	24	Themis	12.3	24.75	3.45	HIP 21881	4.2	4 42.25	+22 57.4	96	39	86
	mar	1	8	29.1	77.02	340	74	Galatea	13.6	74.95	3.19	TYC1 60	4.3	1 45.39	+ 9 9.5	47	161	94
?	mar	9		17.6	5.09	167		Boliviana	12.9	36.88	4.64	HIP 28110	6.0	5 56.47	+ 9 30.6	99	145	17
	mar 1	13	9	16.4	39.49	158		Barbara	13.8	77.32	3.22	TYC1 2173	5.3		+ 2 16.0	44	51	1
	mar 1			3.5	69.82	161		Thetis		52.13		TYC1 2220	5.6		+13 11.2	51	16	9
	mar 1	9	1	13.4	58.97	355	776	Berberici	a 13.3	55.44	2.58	HIP 103389	5.8	20 56.79	-26 17.8	49	103	21
	mar 2	23		7.8	19.52	254	71	Niobe	11.0	35.19	5.20	HIP 76618	5.4	15 38.82	-52 22.4	112	128	66
	mar 2	4		19.9	64.13	166		Europa		42.78		TYC1 1548	5.9	20 57.68		50	165	71
	mar 2			21.2	8.97	347		Echo		78.14		TYC1 1106	6.0		+17 19.6	54	88	89
	mar 2			39.3	59.41	168		Bathilde		61.83		HIP 13654	5.4	2 55.81		36	156	98
	apr	3	8	5.6	46.76	259	1036	Ganymed	13.7	25.85	4.67	TYC1 1417	4.9	15 58.19	-14 16.8	132	28	63

x

X

Date (U.T.)	Minim. Pos. dist. ang.	Minor p Name	lanet Vis. App. Hor. mag. mot. par.	S Designation	t a Vis. mag.	r Right ascens. (2000.0)	Decli- nation (2000.0)	Elon- Ill. gation frac Sun Moon Moon
apr 6 3 0.4 apr 6 4 24.7 apr 6 15 47.4 apr 7 9 41.6 apr 10 13 51.4	12.00 159 76.92 353 93.57 189 0.92 27 48.50 34	308 Polyxo 451 Patientia 423 Diotima 55 Pandora 264 Libussa	"/h " 13.6 60.49 2.66 12.3 42.19 2.92 12.8 31.43 3.18 13.2 25.94 3.36 13.8 15.95 3.42	HIP 110273 HIP 27629 TYC1 3533 HIP 37826 TYC1 3439	1.3	h m 22 20.20 5 50.97 19 13.23 7 45.31 18 8.08	- 7 49.3 +27 58.1 -25 54.4 +28 1.6 -28 27.4	0 0 <b>%</b> 42 32 36 71 144 35 90 23 31 95 155 24 108 84 5
apr 16 9 59.6 apr 23 13 5.3 apr 28 1 36.3 apr 28 19 26.3 may 1 21 55.0	94.28 197 39.14 356 107.50 207 23.41 181 113.99 338	13 Egeria 326 Tamara 107 Camilla 712 Boliviana 46 Hestia	12.0 34.74 3.45 13.4 77.99 4.76 12.5 23.39 3.35 13.5 58.51 3.42 13.2 78.65 3.21	HIP 98421 HIP 108294 TYC1 1420 HIP 34033 TYC1 3920	5.4 5.5 5.5 5.3 5.8	21 56.38 16 0.79 7 3.63	-34 41.9 -37 15.2 - 8 24.7 +10 57.1 + 0 6.6	91 130 12 76 150 82 154 13 99 68 135 96 42 75 73
may 5 18 51.7 may 6 15 3.3 may 9 22 12.4 may 14 19 35.2 may 19 22 52.5	47.79 333 7.62 334 43.69 181 14.16 354 30.53 183	639 Latona 174 Phaedra 776 Berbericia 746 Marlu 48 Doris	13.8 60.11 2.62 12.2 19.23 5.62 12.9 43.46 3.24 14.0 33.12 4.98 12.8 53.26 2.50	TYC1 902 HIP 82676 HIP 109737 HIP 74707 HIP 33929	5.7	16 54.03 22 13.74 15 16.07	+ 6 51.8 -41 48.4 -25 10.9 -41 29.5 +17 45.3	42 33 36 143 83 28 82 54 6 157 148 6 46 47 51
may 21 11 50.9 may 25 15 50.6 may 26 14 55.1 may 28 21 45.1 may 31 3 16.3	17.29 348 8.03 203 75.60 156 22.99 155 21.26 14	176 Iduna 471 Papagena 119 Althaea 75 Eurydike 76 Freia	13.5 61.70 2.54 12.3 43.48 2.73 13.5 55.72 3.42 13.2 86.90 3.66 13.8 46.11 2.64	HIP 7535 TYC1 2719 HIP 116495 HIP 5737 HIP 44001	5.6 5.4 5.8 5.3 5.8	1 37.10 9 8.00 23 36.39 1 13.73 8 57.59	+12 8.5 +29 39.3 + 2 6.1 + 7 34.5 +15 34.9	33 145 69 67 102 99 69 108 100 47 102 93 62 172 77
jun 3 23 13.0 jun 9 3 15.8 jun 15 19 41.3 jun 17 13 58.2 jun 18 8 33.3	80.81 340 86.55 5 14.00 338 90.06 192 36.86 166	240 Vanadis 451 Patientia 308 Polyxo 80 Sappho 779 Nina	14.0 73.26 3.10 12.5 56.50 2.37 13.3 36.91 3.48 13.6 51.00 2.78 13.7 22.60 3.40	TYC1 60 HIP 36393 HIP 117491 HIP 46457 TYC1 1309	5.9	1 45.39 7 29.34 23 49.46 9 28.49 11 56.02	+ 9 9.5 +28 7.1 + 1 4.6 + 8 11.3 -17 9.0	45 34 41 32 54 4 86 147 26 55 31 45 98 22 54
jun 19 5 33.7 jun 21 14 10.6 jun 22 11 20.6 jun 23 2 14.5 jun 23 6 37.4	92.84 338 68.01 10 30.69 156 13.70 146 108.25 162	186 Celuta 774 Armor 596 Scheila 679 Pax 268 Adorea	11.6 34.62 7.51 12.2 22.74 5.62 12.3 29.10 5.92 13.9 30.82 3.97 13.2 18.79 3.93	HIP 90662 HIP 96808 HIP 80079 HIP 77070 HIP 101123	4.6 2.8	18 29.93 19 40.72 16 20.64 15 44.27 20 29.90	-47 13.2 -16 17.6 -24 10.2 + 6 25.5 -18 35.0	154 84 64 155 67 87 156 9 93 134 29 96 146 54 97
jun 23 10 55.9 jul 5 15 20.2 jul 6 13 17.9 jul 11 1 26.4 jul 13 9 41.8	4.85 20 19.37 299 13.85 196 68.90 23 109.20 192	108 Hecuba 618 Elfriede 12 Victoria 148 Gallia 386 Siegena	13.6 16.90 3.28 12.9 23.31 4.24 12.8 54.45 2.66 13.8 41.28 2.45 13.6 48.13 2.25	TYC1 1335 HIP 106039 HIP 48519 TYC1 2918 HIP 49637	4.7 6.0 5.5	21 28.72		104 57 97 145 88 23 44 91 16 58 50 1 39 6 11
jul 16 1 49.2 jul 16 14 50.0 jul 17 1 57.9 jul 20 14 47.4 jul 20 21 33.0	71.67 180 39.19 348 67.67 349 78.47 173 117.28 203	774 Armor 345 Tercidina 42 Isis 415 Palatia 111 Ate	11.8 29.89 5.79 13.6 59.76 3.64 12.6 64.11 3.08 13.8 84.16 3.50 13.4 62.10 2.89	HIP 95176 HIP 13108 HIP 20542 HIP 21673 HIP 54336	4.6 6.0 4.9 5.1 5.7	19 21.73 2 48.53 4 24.10 4 39.15 11 6.90	-15 57.3 +18 17.0 +17 26.6 +15 48.0 + 1 57.3	172 101 37 68 150 43 47 136 49 47 174 85 49 89 87
jul 21 13 3.1 jul 22 11 13.0 jul 23 16 51.6 jul 27 16 59.3 jul 27 20 4.8	63.69 157 107.26 205 2.36 346 25.21 339 34.46 359	735 Marghanna 27 Euterpe 68 Leto 75 Eurydike 762 Pulcova	13.3 64.56 5.61 12.2 55.88 3.25 12.2 59.52 3.02 13.1 65.95 4.39 13.1 30.13 3.59	HIP 10305 HIP 60129 HIP 20186 TYC1 114 TYC1 773	5.8 3.9 5.3 4.4 5.1	2 12.79 12 19.91 4 19.61 3 11.63 20 40.05		88 124 91 65 92 96 53 134 99 73 69 89 176 37 88
jul 31 3 12.5 aug 3 5 55.8 aug 7 7 0.3 aug 9 1 30.5 aug 9 22 57.3	40.95 34 90.92 160 50.58 1 76.79 182 46.11 10	579 Sidonia 75 Eurydike 98 Ianthe 287 Nephthys 441 Bathilde	13.7 39.51 2.78 13.1 62.43 4.50 14.0 15.60 3.71 13.4 64.08 2.90 13.9 64.12 2.61	TYC1 3074 HIP 15737 HIP 145 HIP 29434 TYC1 269	6.0 5.3 5.1 4.9 4.0	3 22.75 0 1.82 6 12.05	- 3 1.7	72 176 62 77 8 33 135 115 3 43 48 0 32 48 2
aug 12 12 38.9 aug 13 4 4.8 aug 14 2 56.8 aug 19 0 28.2 sep 1 22 11.7	82.31 327 40.41 168 5.42 11 112.50 141 24.97 25	464 Megaira 337 Devosa 78 Diana 931 Whittemora 128 Nemesis	12.5 32.77 6.46 12.1 36.92 5.64 12.9 80.13 3.00 13.7 30.16 4.27 13.6 49.60 2.34	TYC1 1546 HIP 109332 HIP 36962 TYC1 1597 HIP 65420	4.2 5.2	22 8.98	+26 53.7 -19 36.8	165 117 19 170 127 25 30 102 35 165 56 85 42 111 32
sep     2     22     17.6       sep     4     7     11.0       sep     4     13     57.6       sep     4     18     47.6       sep     14     5     59.1	94.09 123 12.97 345 116.59 194 100.41 181 59.28 33	357 Ninina 500 Selinur 139 Juewa 103 Hera 6 Hebe	13.2 27.20 4.36 14.0 26.93 4.26 13.2 53.77 3.06 13.0 38.80 3.26 9.8 34.08 4.93	TYC1 2048 HIP 85755 HIP 72323 HIP 25499 TYC1 3454	5.6 5.4	17 31.42 14 47.38	-25 37.5 +17 57.7	150 98 22 101 141 11 65 101 9 79 48 8 102 10 55
sep 14     7 51.4       sep 14     21 0.3       sep 17 0 52.6     52.6       sep 27 13 47.3     oct 3 17 9.8	86.88 165 93.44 148 2.30 15 99.29 6 67.15 267	71 Niobe 798 Ruth 345 Tercidina 415 Palatia 312 Pierretta	11.9 58.19 3.72 13.7 31.36 4.35 12.8 26.64 5.28 13.4 67.43 4.59 13.1 13.36 4.83	HIP 80390 TYC1 1614 TYC1 1115 HIP 34440 HIP 106039	4.3 5.5 5.6		+ 6 22.7 +19 36.6 +15 55.8	79 23 56 170 80 62 109 121 81 77 37 71 127 166 11
oct 8 7 5.7 oct 12 10 55.8 oct 13 16 4.6 oct 15 17 1.1 oct 19 19 14.2	31.88 355 9.36 211 47.08 196 85.37 18 86.11 25		13.3 54.34 3.08 13.7 39.69 4.15 11.4 78.93 5.85 12.1 53.37 2.76 11.6 48.05 8.90	HIP 83176 HIP 33914 HIP 94477 HIP 49669 HIP 12828	5.9 5.9 5.1 1.3 4.3	7 2.29 19 13.71 10 8.37		61 36 5 93 169 42 90 29 55 52 170 74 163 32 98
oct 21 9 29.2 oct 22 22 27.1 oct 29 20 29.6 nov 2 0 39.6 nov 3 1 43.2	52.55 355 91.77 19 32.99 43 79.07 197 2.41 348	65 Cybele 22 Kalliope 2 Pallas 109 Felicitas 431 Mephele	12.7 30.21 2.85 12.2 54.43 2.50 10.2 28.47 2.84 10.7 31.34 9.33 13.2 27.38 4.31	TYC1 1517 HIP 56080 HIP 101847 TYC1 1088 TYC1 2288	4.9 5.9 4.5 5.7 5.9	11 29.70 20 38.34 3 9.61	-19 45.7 +15 24.8 - 1 6.3 +29 4.6 +18 11.6	87 94 100 43 120 98 95 158 43 162 124 12 158 133 5
nov 5 4 6.5 nov 11 10 8.4 nov 12 2 58.1 nov 15 10 36.3 nov 20 1 2.6	89.18 351 67.33 214 70.39 12 15.67 173 25.38 354	774 Armor 89 Julia 17 Thetis 1074 Beljawskya 161 Athor	13.9 49.55 3.29 12.2 37.70 2.96 12.9 24.50 3.60 13.7 28.03 5.30 12.2 39.46 6.17	HIP 98258 TYC1 2835 HIP 44001 HIP 20641 HIP 14893		4 25.42		75 71 0 73 158 46 97 168 53 164 66 82 169 11 100

6

U																	
	Date	⊋ (U.	т.)	Minim. dist.			ame .	Vis.	n e t App. mot.		S Designation	t a Vis. mag.	r Right ascens. (2000.0)	Decli- nation (2000.0)	Elon- gati Sun		Ill. frac Moon
													_				
			h m	•	0				•/h				h m	o ·	0	0	8
	nov 24		2.0	107.26	201		Chaldaea		63.80		TYC1 2841		10 30.29	- 0 38.2	82	48	81
	nov 24		21.8	31.14			Sorga		30.77		HIP 13654	5.4	2 55.81		164	69	80
	nov 27		57.5	109.09			Justitia		35.30		HIP 21036	5.5		+13 43.5	171	84	53
	nov 28		5.5	26.26			Germania		29.33		HIP 25695	5.5		+25 9.0	162	83	41
	nov 28	3 17	39.3	109.85	169	65	Cybele	13.0	44.97	2.42	HIP 100881	5.1	20 27.32	-18 12.7	58	134	38
	nov 29	1:	28.3	47.93	118	344	Desiderata	13.9	17.07	3.05	HIP 47080	5.5	9 35.66	+35 48.6	111	50	28
	dec 1	. 13	14.6	28.24	354	236	Honoria	13.6	67.10	2.94	HIP 96808	5.5	19 40.72	-16 17.6	45	83	11
	dec 2	2	18.8	90.21	173	206	Hersilia	12.1	34.56	5.31	HIP 20894	3.5	4 28.66	+15 52.3	173	155	5
x	dec 2	2 20	29.1	6.31		282	Clorinde	13.7	33.81	7.21	TYC1 2258	5.6	3 39.85	+ 3 3.4	156	165	3
	dec 3	3 (	12.8	36.92	343	678	Fredegundis	13.7	60.89	3.80	HIP 105761	5.8	21 25.22	- 9 44.9	70	88	2
	dec 3	3 2:	16.8	90.08	360	731	Sorga	14.0	24.86	4.49	HIP 13108	6.0	2 48.53	+18 17.0	153	158	0
	dec		35.2	91.82			Cantabia		31.13		HIP 27316	5.8		+14 29.3	165	145	14
	dec 10		5.0	60.96	333		Devosa		50.88		TYC1 3771	5.4		-11 33.9	72	5	35
	dec 10		42.3	90.44	148		Brang"ne		32.98		HIP 18434	5.5		+35 4.9	160	91	39
	dec 11		46.0	26.27	175		Hersilia		31.91		HIP 20205	3.8		+15 37.7	164	72	53
	dec 1	, ,	3 33.2	62.52	350	790	Pretoria	13.9	54.10	2.52	HIP 101847	4.5	20 38.34	- 1 6.3	50	83	82
	dec 1		5.7	17.17			Nausikaa		54.39		HIP 73184		14 57.47		35	164	83
	dec 16		11.4	42.51			Brang"ne		29.44		HIP 18081	5.8		+34 21.5	155	26	88
	dec 16		7 58.1	28.87			Ludmilla		42.75		HIP 65301	5.5			59	154	91
	dec 18		36.3	92.73	105		Marghanna		19.54		HIP 9110	5.3		+17 49.1	127	36	98
		•		22.75		. 33	····· J			,		3.3	2 27.00			30	20
	dec 18	3 2:	4.8	89.63	3	120	Lachesis	12.4	29.76	3.92	HIP 31173	5.9	6 32.45	+32 27.3	166	21	99
	dec 29	11	31.2	119.10	182	203	Pompeja	12.5	34.91	5.34	TYC1 2504	5.3	6 35.20	+28 1.3	175	125	21
2003	3 jan 2	2 8	50.2	4.13	162	110	Lydia	13.1	67.38	2.67	TYC1 3690	5.3	21 8.56	-21 11.6	32	38	0

### A COLLABORATIVE WORK ON THREE ASTEROID LIGHTCURVES: 506 MARION, 585 BILKIS, 1506 XOSA

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The lightcurve periods for three asteroids, 506 Marion, 585 Bilkis, and 1506 Xosa were determined through the collaborative work of the two authors. The period for 506 Marion was determined to be  $10.58 \pm 0.01h$  with an amplitude of  $0.35 \pm 0.03m$ . The lightcurve for 585 Bilkis was found to have a period of  $6.442 \pm 0.002h$  and an amplitude of  $0.38 \pm 0.02m$ . Asteroid 1506 Xosa was found to have a lightcurve period of  $5.9 \pm 0.01h$  and an amplitude of  $0.28 \pm 0.02m$ .

When attending the Asteroids III conference in Palermo, Sicily, in 2001 June, we met with many professionals who encouraged a more coordinated collaborative effort among amateurs and between amateurs and professionals. The primary goal of this effort is to help concentrate observations on targets that are most needing of work, e.g., radar targets, asteroids with very long periods or periods with aliases close to 24 hours, and so on. Every journey begins with a first step and ours upon returning from Italy was to work out the process whereby one or both of us would accumulate images while one of us (Warner) would reduce the data and determine the lightcurve characteristics. Eventually both of us plan to reduce data gathered from each other and by other amateurs.

For 506 Marion and 585 Bilkis, Robinson took all images from his backyard observatory in Olathe, Kansas, a short distance from downtown Kansas City. He used a 0.30m SCT and SBIG ST-9E for images of 60s duration separated by approximately three minutes. The unguided images were accumulated on a hard disk, copied to a CD-ROM, and then sent to Warner for reduction using MPO Canopus. For 1506 Xosa, both authors acquired images, Warner using a 0.5m f/8.1 Ritchey-Chretien and FLI KAF-1001E. Exposures were also 60s in duration and spaced about 3 minutes apart.

The disparity in signal-to-noise ratios between the different systems made reductions a bit more difficult than might have been, mostly because of the increased noise in the Robinson data. Too, in some cases, the data sets were separated by several weeks, meaning adjustments had to be made for changing phase angle and distances of the asteroid from the earth and sun. Despite the difficulties, reasonable results for all three asteroids were obtained and are published below.

### 506 Marion

R. Dugan discovered 506 Marion in 1903 February and it was subsequently named for his cousin. It is a type C asteroid of about 104km size. As might be guessed from its spectral classification, Marion resides in the outer reaches of the main belt, having a semimajor axis of 3.04AU and inclination of about 17°.

Robinson obtained observations on the nights of 2001 July 1, 14, 21, and 30. Data were also obtained on two other nights, 2001 June 30 and August 1. They were of short duration and suffered from considerable noise and so were not used in the period determinations. Over the total span, the phase angle of the asteroid decreased from about 14° to 7°. Using the four "good" sessions, a period of  $10.58 \pm 0.01h$  was found as well as an amplitude of  $0.35 \pm 0.03m$ . A plot of the phased curve is shown in Figure 1.

### 585 Bilkis

Bilkis is a type C asteroid of about 52 km size. It has a semi-major axis of about 2.4AU and inclination of 7.5°. It was discovered by A. Kopff in 1906 February and was given the Koran name for the Queen of Sheba. Robinson obtained data on several nights in 2001 May through August. Specifically, the data used to find a

lightcurve solution were obtained on 2001 May 14, 15, June 25, July 8, 9 and 16. During the span of used observations, the phase angle of Bilkis decreased from  $17^{\circ}$  to about  $11.5^{\circ}$ . The resulting lightcurve, shown in the phased plot of Figure 2, has a period of  $6.442 \pm 0.002h$  and amplitude of  $0.38 \pm 0.02m$ 

### 1506 Xosa

The effort to determine the period of this asteroid was truly collaborative as both authors contributed data towards finding a lightcurve. Robinson's data were acquired on 2001 July 29 and August 3 while Warner worked the asteroid 2001 July 25 and August 8. Data were also obtained on other nights but were either dominated by noise or too short of coverage to contribute effectively to a solution. Figure 3 shows a phased plot based on the period of 5.9 ±0.01h. The amplitude of the curve is 0.28 ±0.02m

C. Jackson found Xosa in 1939 May while working in Johannesburg. The name is that of a South African tribe. The asteroid's orbit has a semi-major axis of about 2.55AU and inclination of 12.5°. Its size is estimated to be about 14-15km but its spectral classification has not yet been determined.

#### More on Collaborative Work

The ultimate goal of collaboration is to obtain continuous coverage for the entire rotational period of an asteroid. In the case of long period rotators or objects with periods that are synchronous with our diurnal period, this may require observation by observers spaced around the globe. This type of collaboration will involve a variety of optics, detectors and observing conditions. The work on 1506 Xosa illustrates how it is possible to combine data collected from different sites and even at different segments of the rotational period and still obtain a reliable solution.

### Acknowledgements

The authors thank the entire professional community at the Asteroids III conference for extending a warm welcome and for their encouragement of amateur efforts around the world. In particular, we thank Richard Binzel, co-chairman of the conference, Alan Harris of Jet Propulsion Laboratory, Petr Pravec of Ondrejov Observatory, Czech Republic, and Ted Bowell of Lowell Observatory for their extraordinary support and assistance.

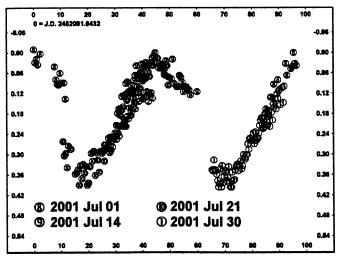


Figure 1. The phased plot for the data used on 506 Marion. The period is  $10.58 \pm 0.01h$  with an amplitude of  $0.35m \pm 0.03m$ . The data were obtained at the Sunflower observatory during 2001 July.

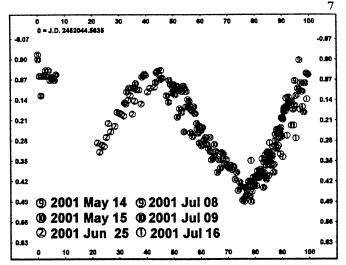


Figure 2. Data taken at Sunflower Observatory during the period 2001 May 14 - July 16 were used to create this phased plot based on a period of  $6.442 \pm 0.002h$ . The amplitude is  $0.38 \pm 0.02m$ 

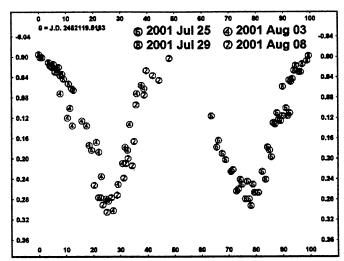


Figure 3. Observations from Sunflower and Palmer Divide Observatories taken in late 2001 July and early August produced this phased plot based on a period of  $5.9 \pm 0.01h$ . The amplitude was found to be  $0.28 \pm 0.02m$ .

8

### MINOR PLANETS AT UNUSUALLY FAVORABLE ELONGATIONS IN 2002

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A list is presented of minor planets which are much brighter than usual at their 2002 apparitions. A very close approach within 0.077 AU of 5604 1992 FE, and moderately close approaches of planets 1580 Betulia and 6455 1992 HE, are especially noteworthy.

The minor planets in the lists which follow will be much brighter at their 2002 apparitions than at their average distances at maximum elongation. Many years may pass before these planets will be again as bright as in 2002. Observers are encouraged to give special attention to those which lie near the limit of their equipment.

Planet 1580 Betulia makes 4 revolutions every 13 years, and continues in 2002 the series of close approaches of 1950, 1963, 1976, and 1989. However the period is not exactly commensurable with Earth's, and each is now farther distant than the previous. Not again in the 21st century will Betulia become as bright as magnitude 15.0.

Planet 5604 1992 FE approaches with 0.077 AU on 2002 June 22 in an event observable only at far southern declinations, being south of declination -30° the entire interval in which it is brighter than magnitude 15.

Planet 6455 1992 HE has two moderately close approaches in the year 2002, to 0.366 AU on May 4, and to 0.444 AU on September 28, brighter than magnitude 14 on both occasions.

These lists have been prepared by an examination of the maximum elongation circumstances of minor planets computed by the author for all years through 2060 with a full perturbation program written by Dr. John Reed, and to whom he expresses his thanks. Elements are from EMP 1992, except that for all planets for which new or improved elements have been published subsequently in the Minor Planet Circulars, the newer elements have been used. Planetary positions are from the JPL DE-200 ephemeris, courtesy of Dr. E. Myles Standish. Dr. Reed's ephemeris generating program, a list of minor planet elements, and the JPL planetary ephemeris are freeware which may be obtained from the author by sending a 100 Megabyte zip disk and stamped, addressed return mailer. They cannot be downloaded directly over the Internet.

Any planets whose brightest magnitudes near the time of maximum elongation vary by at least 2.0 in this interval and in 2002 will be within 0.3 of the brightest occurring, or vary by at least 3.0 and in 2002 will be within 0.5 of the brightest occurring; and which are visual magnitude 14.5 or brighter, are included. For planets brighter than visual magnitude 13.5, which are within the range of a large number of observers, these standards have been relaxed somewhat to include a larger number of planets. Magnitudes have been computed from the updated magnitude parameters published in MPC28104-28116, on 1996 Nov. 25, or more recently in the Minor Planet Circulars. In a few cases absolute magnitudes have been improved by the results of the Magnitude Alert Program.

Oppositions may be in right ascension or in celestial longitude. Here we use still a third representation, maximum elongation from the Sun, instead of opposition. Though unconventional, it has the advantage that many close approaches do not involve actual opposition to the Sun near the time of minimum distance and greatest brightness and are missed by an opposition-based program. Other data are also provided according to the following tabular listings: Minor planet number, date of maximum elongation from the Sun in format yyyy/mm/dd, maximum elongation, declination on date of maximum elongation, declination on date of maximum elongation, declination on date of minimum or brightest magnitude in format yyyy/mm/dd, minimum magnitude, date of minimum distance in format yyyy/mm/dd, and minimum distance in AU.

Users should note that when the maximum elongation is about  $177^{\circ}$  or greater, the minimum magnitude is sharply peaked due to enhanced brightening near zero phase angle. Even as near as 10 days before or after minimum magnitude the magnitude is generally about 0.4 greater. This effect takes place in greater time interval for smaller maximum elongations. There is some interest in very small minimum phase angles. For maximum elongations E near  $180^{\circ}$  at Earth distance  $\Delta$ , an approximate formula for the minimum phase angle  $\phi$  is  $\phi = (180^{\circ} - E)/(\Delta + 1)$ .

Table 1. Numerical Sequence of Favorable Elongations

Planet	Max Elon D	Hax E	RA	Dec	Min Mag D	Mag	Min Dist D	Min Dist
15					2002/09/15		2002/09/17	1.251
18			0h38m		2002/09/27		2002/09/28	0.814
20			5h 7m		2002/12/10		2002/12/12	1.103
25					2002/06/22		2002/06/26	0.929
28	2002/02/21	1//.6	10023M	+12-	2002/02/21	9.9	2002/02/20	1.406
32	2002/04/06	176.0°	12h53m	- 9*	2002/04/06	10.3	2002/04/07	1.376
77			5h 5m		2002/12/10		2002/12/08	1.344
86	2002/10/13	172.60	1h26m	+ 1*	2002/10/13	11.7	2002/10/14	1.496
109	2002/11/12	168.0°	2h59m	+29°	2002/11/12	10.5	2002/11/13	0.927
111	2002/02/10	177.6	9h34m	+11*	2002/02/10	10.7	2002/02/09	1.373
121	2002/10/20	171 60	1h49m	<b>_</b> ~•	2002/10/19		2002/10/17	1.968
137			Oh 4m		2002/10/19		2002/09/20	1.616
162					2002/04/21		2002/04/17	1.713
188					2002/06/01		2002/06/08	1.430
197					2002/08/08		2002/08/08	1.300
200			2h55m	+29°	2002/11/12	11.3	2002/11/11	1.400
229					2002/09/12		2002/09/10	1.908
239			01a 3an		2002/09/24		2002/09/27	1.342
240			4h39m		2002/12/03		2002/11/30	1.146
278	2002/04/10	170.9	13h26m	+ 0*	2002/04/10	12.4	2002/04/11	1.399
322	2002/09/06	164 9	22h 38m	+ 70	2002/09/07	11 5	2002/09/10	1.142
350			4h48m		2002/12/05		2002/12/05	1.665
352			0h20m		2002/10/02		2002/10/03	0.888
361	2002/01/28	162.9	9h 2m		2002/01/27		2002/01/24	2.291
390	2002/02/20	170.3°	10h 4m	+ 1*	2002/02/20	13.2	2002/02/19	1.340
		<b>-</b>						
428			1h13m		2002/10/13		2002/10/14	0.907
429 444			0h34m 0h35m		2002/10/06		2002/10/06 2002/09/30	1.304
459					2002/10/03		2002/03/30	1.112
464					2002/08/02		2002/08/08	1.358
488					2002/02/14		2002/02/15	1.659
503					2002/02/08	11.9	2002/02/05	1.311
549					2002/01/28		2002/01/21	1.089
593					2002/12/05		2002/12/09	1.191
596	2002/06/02	179.5	16h39m	-22°	2002/06/02	11.7	2002/06/02	1.432
638	2002/05/18	170 0	15h47m	_ 0*	2002/05/18	12 5	2002/05/18	1.288
654					2002/01/10		2002/01/13	0.808
668					2002/10/28		2002/10/19	1.418
670					2002/09/20		2002/09/21	1.269
676					2002/08/18		2002/08/18	1.654
688					2002/08/28		2002/08/25	1.342
711					2002/07/09		2002/07/10	0.798
725					2002/12/12		2002/12/05	1.151
735 753					2002/11/01		2002/10/21	0.988
/53	2002/06/02	1/3.9	TOH 2 AM	-20	2002/06/02	11.3	2002/06/05	0.807
759	2002/08/27	168.8°	22h13m	+ 0*	2002/08/26	12.7	2002/08/23	1.089
774					2002/07/12		2002/07/11	1.515
784					2002/05/05		2002/05/10	1.410
788	2002/03/15	178.6°	11h37m	+ 0*	2002/03/15		2002/03/19	1.867
789	2002/06/13	170.3	17h31m	-13°	2002/06/13	13.6	2002/06/15	1.288

Planet	Max Elon D F	ax E	RA Dec	Min Mag D )	Mag	Min Dist D I	in Dist
817	2002/10/13	160 19	1h43m -10°	2002/10/14	12 5	2002/10/15	1.210
963	2002/11/25	179.4	4h 1m +20°	2002/11/25	13.9	2002/11/24	0.953
969 984	2002/10/06 2002/10/07	175.3°	0h38m + 9° 0h28m +18°	2002/10/06 2002/10/06		2002/10/05 2002/10/05	0.966 1.276
985	2002/07/28	177.6	20h28m -16°	2002/07/28		2002/08/08	0.746
1000	2002/04/28	146.9°	13h33m -45°	2002/05/01	13.2	2002/05/04	1.507
1013	2002/12/10	163.4°	5h 7m +39°	2002/12/12	12.9	2002/12/16	1.270
1040 1052	2002/01/09	175.8°	7h18m +18° 5h29m +20°	2002/01/08 2002/12/15		2002/01/05 2002/12/10	1.576
1059	2002/05/27	166.8	16h28m - 8°	2002/05/28		2002/06/02	1.235
1060	2002/08/24	167.5°	21h53m + 0°	2002/08/23	14.1	2002/08/18	0.810
1074	2002/11/28	179.20	21h53m + 0° 4h14m +21°	2002/11/28	13.2	2002/11/27	1.641
1088 1107	2002/01/07	179.10	7h15m +21°	2002/10/03 2002/01/07		2002/10/02 2002/01/07	0.782 1.788
1116	2002/12/27	151.8	6h25m +51°	2002/12/25	12.9	2002/12/24	1.352
1125	2002/11/30	177.6	4h25m +19°	2002/11/30	14.4	2002/12/04	1.579
1130	2002/07/17	175.30	19h41m -16° 19h29m -23°	2002/07/17	13.3	2002/07/23	0.824
1131 1153	2002/08/22	174.5°	21h58m - 6°		13.4	2002/07/27 2002/08/18	0.739 0.848
1155	2002/05/31	173.3°	16h31m -28°	2002/05/31	13.5	2002/05/30	1.066
1187	2002/12/19	163.0°	5h41m +40°	2002/12/17	13.9	2002/12/12	1.207
1222 1275	2002/06/25 2002/10/24	161.0°	18h28m - 4° 2h12m + 3°	2002/06/27 2002/10/24		2002/07/01 2002/10/22	1.158
1281	2002/05/11	175.2	15h16m -13°	2002/05/11	13.9	2002/05/19	1.160
1324			20h54m -16°	2002/08/03	13.9	2002/07/29	0.861
1354	2002/05/28	173.80	16h15m -27° 21h30m -20° 23h56m - 4°	2002/05/28		2002/06/02	1.507
1397 1527	2002/08/11	174.0° 176.3°	21h30m -20° 23h56m - 4°	2002/08/10 2002/09/20	13.5	2002/08/03 2002/09/13	1.068 0.839
1530	2002/08/27	172.7	22h16m - 2° 8h11m +35°	2002/08/28	14.5	2002/09/02	0.851
1534	2002/01/16	164.7	8h11m +35°	2002/01/15	14.0	2002/01/14	1.079
1547	2002/01/15	178.4	7h48m +19°	2002/01/15	12.5	2002/01/10	1.030
1550 1580	2002/12/01 2002/05/31	178.8° 116.1°	4h28m +22° 12h54m +13°	2002/12/01 2002/05/30	13.3	2002/11/18 2002/05/29	0.944 0.238
1585	2002/08/19	176.6°	21h48m - 9°	2002/08/20	14.0	2002/08/29	1.624
1658	2002/06/23	178.8	18h 8m -22°	2002/06/23		2002/06/25	1.085
1709	2002/08/10	175.0°	21h17m -10°	2002/08/11	14.2	2002/08/14	0.884 1.149
1734 1761	2002/02/11	176.2°	1h13m + 3° 9h43m +17°	2002/10/11 2002/02/11	14.3	2002/10/08 2002/02/09	1.431
1789 1874		179.5°	19h 7m -22° 0h34m - 0°	2002/07/07	14.0	2002/07/13 2002/09/26	0.818 1.227
10/4							
1883 1885	2002/08/10	170 80	21h53m _12*	2002/11/25 2002/08/19		2002/11/22 2002/08/19	0.819 0.681
1922	2002/05/16	151.00	16h51m + 2°	2002/05/18	13.6	2002/05/21	0.742
1954 2013		179.4°	18h38m -23° 0h52m -12°	2002/07/01 2002/09/28		2002/07/12 2002/09/22	1.346
2093 2236	2002/08/03	174.7°	20h45m -12° 22h16m -26° 22h51m - 6°	2002/08/02 2002/08/20		2002/07/28 2002/08/16	0.926 0.846
2254	2002/09/04	179.4	22h51m - 6°	2002/09/04	14.1	2002/09/05	1.001
2257 2362	2002/10/27	176.4°	2h 1m +16° 23h48m - 3°	2002/10/27 2002/09/19		2002/10/20 2002/09/18	0.955 0.768
2433 2536	2002/04/04	178.1°	12h55m - 3° 21h 8m - 9° 3h32m +10°	2002/04/04 2002/08/10	14.5	2002/04/13 2002/08/17	1.245 0.884
2569 2585	2002/11/17 2002/10/29	172.0	3h32m +10° 2h28m + 1°	2002/11/16 2002/10/28	13.8	2002/11/13 2002/10/25	1.246
2755	2002/10/29	178.5°	2h41m +17°	2002/10/28	14.5	2002/10/25	0.882 1.166
2768	2002/09/11	168.19	23h37m -15°	2002/09/12	14.0	2002/09/16	0.918
2820	2002/10/03	177.1°	0h33ma + 6°	2002/10/03	14.2	2002/10/03	0.871
2829 2830	2002/04/27	166.20	18h50m -39° 14h23m - 0°	2002/07/04 2002/04/27	13.7	2002/07/06 2002/04/25	1.534 0.925
2834	2002/01/12	168.3°	7h30m + 9°	2002/01/13		2002/01/14	1.180
3034	2002/09/11	175.9°	23h24m - 8°	2002/09/11	13.6	2002/09/05	0.854
3093	2002/09/16	157.1	22h59m +18°	2002/09/14	14.2	2002/09/12	1.171
3127 3485	2002/08/30	179.5	20h41m -18°	2002/08/29 2002/07/31	14.3	2002/08/27 2002/07/29	1.078 1.023
3679	2002/10/15	171.3	22h24m - 3° 20h41m -18° 1h 6m +16°	2002/10/14	14.5	2002/10/11	0.724
	2002/07/01	170.10	18h42m =23°	2002/07/01			
3761 3768	2002/04/28	174.9°	14h31m - 9° 21h21m -11° 23h15m -12°	2002/04/29 2002/08/12		2002/05/09 2002/08/17	1.633 1.564
4155	2002/09/08	173.1°	23h15m -12°	2002/09/08	14.2	2002/09/12	0.885
4287	2002/08/12	167.5°	21h45m -26°	2002/08/12	14.5	2002/08/09	0.795
4288	2002/11/18	175.0	3h31m +24° 21h26m -23°	2002/11/18		2002/11/12 2002/08/14	1.286 1.475
4577	2002/02/05	160.9°	9h44m +33°	2002/08/11 2002/02/04	14.0	2002/02/02	0.922
4925 5095	2002/10/13	169.8	0h56m +17° 21h49m -12°	2002/10/13 2002/08/18		2002/10/12 2002/08/19	1.326 0.891
5142 5349	2002/11/03 2002/10/10	176.4°	2h39m +11° 1h 9m +33°	2002/11/03 2002/10/20			0.859 0.756
5391				2002/00/00			0.734
5438 5604	2002/01/28 2002/05/31	170.0° 143.8°	8h39m + 8° 17h44m -55°	2002/01/29 2002/06/14	13.6	2002/02/05 2002/06/22	1.155 0.077
6455 6455	2002/10/18	165.3*	13h32m - 1° 1h17m - 4°			2002/05/04 2002/09/28	
6522	2002/07/14	177.5	19h34m -24° 3h19m +15°	2002/07/14 2002/11/14	14.4	2002/07/08	0.979
6610 7564	2002/11/14 2002/03/03	178.2°	3h19m +15° 10h58m + 8°	2002/11/14 2002/03/03	14.3	2002/11/13 2002/03/02	0.909 1.235
9021	2002/06/09 2002/08/15 2002/10/25 2002/01/02	178.5°	21h39m -12°	2002/06/09 2002/08/15	14.3	2002/08/14	1.118
10142	2002/10/25	165.0	2h13m - 2°	2002/10/25	14.4	2002/10/23	1.080
15127	2002/01/02	176.6°	23h53m - 4°	2002/01/02 2002/09/20		2002/01/09 2002/09/21	
16143			22h41m + 7°			2002/08/29	
18487	2002/07/18	170.8	19h39m -12°	2002/07/19	14.2	2002/07/22	1.215

Table 2. Temporal Sequence of Favorable Elongations

	Max Elon D M		RA Dec	Min Mag E Mag	Min Dist D P	
10772 1107	2002/01/02 2002/01/07	178.0°	6h50m +20° 7h15m +21°	2002/01/02 14.0 2002/01/07 12.6	2002/01/09 2002/01/07	0.891 1.788
654	2002/01/09	171.7	7h25m +13°	2002/01/10 9.7	2002/01/13	0.808
1040 2834	2002/01/09 2002/01/12		7h18m +18° 7h30m + 9°	2002/01/08 13.6 2002/01/13 14.5	2002/01/05 2002/01/14	1.576 1.180
1547	2002/01/15	178.4*	7h48m +19°	2002/01/15 12.5	2002/01/10	1.030
1534 361	2002/01/16 2002/01/28	164.7°	8h11m +35° 9h 2m +34°	2002/01/15 14.0 2002/01/27 13.0	2002/01/14 2002/01/24	1.079 2.291
549	2002/01/28	177.8°	8h40m +16°	2002/01/28 13.0	2002/01/21	1.089
5438	2002/01/28	170.0°	8h39m + 8°	2002/01/29 13.6	2002/02/05	1.155
4577	2002/02/05	160.9°	9h44m +33°	2002/02/04 14.0	2002/02/02	0.922
503 111	2002/02/08 2002/02/10	171.7°	9h40m +22° 9h34m +11°	2002/02/08 11.9 2002/02/10 10.7	2002/02/05 2002/02/09	1.311
1761	2002/02/11	176.2°	9h43m +17°	2002/02/11 14.3	2002/02/09	1.431
488 390	2002/02/13 2002/02/20	164.0°	10h15m +27°	2002/02/14 11.5 2002/02/20 13.2	2002/02/15 2002/02/19	1.659 1.340
28	2002/02/21	177.6°	10h23m +12°	2002/02/20 13.2	2002/02/19	1.406
7564	2002/02/02	170 20	10h58m + 8°	2002/03/03 14.5	2002/03/02	1.235
788	2002/03/03	178.6°	11h37m + 0°	2002/03/03 14.5	2002/03/02	1.867
2433	2002/04/04		12155- 10	2002/04/04 14 2	2002/04/22	1 245
32	2002/04/06	176.0°	12h55m - 3° 12h53m - 9°	2002/04/04 14.2 2002/04/06 10.3	2002/04/13 2002/04/07	1.245 1.376
278	2002/04/10	170.9°	13h26m + 0° 13h32m - 1°	2002/04/10 12.4	2002/04/11	1.399
6455 162	2002/04/22	179.0°	13h58m -11°	2002/04/19 13.2 2002/04/21 12.3	2002/05/04 2002/04/17	0.366 1.713
2830	2002/04/27	166.2	14h23m - 0°	2002/04/27 14.4	2002/04/25	0.925
1000 3761	2002/04/28	174.9°	13h33m -45° 14h31m - 9°	2002/05/01 13.2 2002/04/29 14.5	2002/05/04 2002/05/09	1.507 1.633
784 1281	2002/05/04	169.5°	14h36m -26° 15h16m -13°	2002/05/05 12.1 2002/05/11 13.9	2002/05/10 2002/05/19	1.410
1922	2002/05/16	151.00	16h5lm + 2°	2002/05/18 13.6	2002/05/21	0.742
638 1059			15h47m - 9° 16h28m - 8°	2002/05/18 12.5 2002/05/28 13.4	2002/05/18 2002/06/02	1.288
1354	2002/05/28	173.80	16h15m -27*	2002/05/28 14.5	2002/06/02	1.507
1155 1580	2002/05/31	173.3	16h31m -28° 12h54m +13° 17h44m -55°	2002/05/31 13.5 2002/05/30 14.0	2002/05/30 2002/05/29	1.066 0.238
5604	2002/05/31	143.8	17h44m -55°	2002/06/14 13.2	2002/06/22	0.077
188	2002/06/01	176 69	16h39m -18°	2002/06/01 12.1	2002/06/08	1.430
596	2002/06/02	179.5	16h39m -22°	2002/06/02 11.7	2002/06/02	1.432
753 8270	2002/06/02	175.9°	16h39m -26° 17h 5m -32°	2002/06/02 11.3 2002/06/09 14.4	2002/06/05 2002/06/13	0.807 0.820
789	2002/06/13	170.3	17h31m -13°	2002/06/13 13.6	2002/06/15	1.288
25 1658	2002/06/15	142.1	18h12m +13° 18h 8m -22°	2002/06/22 10.0 2002/06/23 13.4	2002/06/26 2002/06/25	0.929 1.085
1222	2002/06/25	161.0	18h28m - 4°	2002/06/23 13.4	2002/07/01	1.158
1024	2000/07/01		101.20- 220	2002/07/01 12 0	2002/07/12	1 146
1954 3699	2002/07/01	179.10	18h38m -23° 18h42m -23°	2002/07/01 13.9 2002/07/01 14.4	2002/07/12 2002/07/08	1.346 1.030
2829	2002/07/03	163.4	18h50m -39°	2002/07/04 13.7	2002/07/06	1.534
1789 711	2002/07/07	166.9	19h 7m -22° 19h17m -35°	2002/07/07 14.0 2002/07/09 13.2	2002/07/13 2002/07/10	0.818 0.798
774	2002/07/12	174.0°	19h24m -15°	2002/07/12 11.8	2002/07/11	1.515
1131 6522	2002/07/13	178.7°	19h29m -23° 19h34m -24°	2002/07/13 13.8 2002/07/14 14.4	2002/07/27 2002/07/08	0.739 0.979
1130	2002/07/17	175.3	19h4lm -16°	2002/07/17 13.3	2002/07/23	0.824
18487 985	2002/07/18	177.6	19h39m -12° 20h28m -16°	2002/07/19 14.2 2002/07/28 13.5	2002/07/22 2002/08/08	1.215 0.746
3485	2002/07/31	179.5	20h4lm -18°	2002/07/31 14.3	2002/07/29	1.023
464	2002/08/02	171.8°	21h 0m -25°	2002/08/02 12.4	2002/08/08	1.358
1324	2002/08/03	178.7	20h54m -16°	2002/08/03 13.9	2002/07/29	0.861
2093 197	2002/08/03	174.7	20h45m -12° 21h30m -28°	2002/08/02 14.2 2002/08/08 12.0	2002/07/28 2002/08/08	0.926 1.300
2536	2002/08/09	173.10	21h 8m - 9°	2002/08/10 14.5	2002/08/17	0.884
1709 4460	2002/08/10	175.0	21h17m -10° 21h26m -23°	2002/08/11 14.2 2002/08/11 14.2	2002/08/14 2002/08/14	0.884 1.475
1397	2002/08/11	174.0	21h30m ~20°	2002/08/10 13.5	2002/08/03	1.068
3768 4287	2002/08/12	167.50	21h21m -11° 21h45m -26°	2002/08/12 14.5 2002/08/12 14.5	2002/08/17 2002/08/09	1.564 0.795
9021	2002/08/15	178.5	21h39m -12°	2002/08/15 14.3	2002/08/14	1.118
676 5095	2002/08/18	177.9°	21h39m -12° 21h46m -11° 21h49m -12°	2002/08/18 12.7 2002/08/18 14.4	2002/08/18 2002/08/19	1.654 0.891
1585	2002/08/19	176.6°	21h48m - 9°	2002/08/20 14.0	2002/08/29	1.624
1885 2236	2002/08/19	179.8	21h53m -12° 22h16m -26°	2002/08/19 14.0 2002/08/20 13.8	2002/08/19 2002/08/16	0.681 0.846
1153	2002/08/22	174.5	21h58m - 6°	2002/08/22 13.4	2002/08/18	0.848
1060 759	2002/08/24	167.5	21h53m + 0° 22h13m + 0°	2002/08/23 14.1 2002/08/26 12.7		0.810 1.089
1530	2002/08/27	172.70	22h16m - 2°	2002/08/28 14.5	2002/09/02	0.851
688	2002/08/28 2002/08/30	175.0	22h20m - 4°	2002/08/28 13.3 2002/08/29 14.3		
2254 16143	2002/09/04	179.4° 165.2°	22h5lm - 6° 22h4lm + 7°	2002/09/04 14.1 2002/09/02 14.5	2002/08/29	
322	2002/09/06	164.9°	22h38m + 7°	2002/09/07 11.5	2002/09/10	1.142
4155 5301	2002/09/08 2002/09/09	173.1°	23h15m -12° 23h11m - 4°	2002/09/08 14.2 2002/09/09 13.9		
2768	2002/09/11	168.1*	23h37m -15°	2002/09/12 14.0	2002/09/16	0.918
3034 229	2002/09/11	175.00	23h24m - 8°	2002/09/11 13.6		0.854
15	2002/09/13	162.1	22h59m +13°	2002/09/12 13.0 2002/09/15 8.0	2002/09/10	1.251
3093 2362	2002/09/16 2002/09/19	157.1	22h59m +18"	2002/09/14 14.2	2002/09/12	
670	2002/09/20	179.40	23h51m - 10	2002/09/20 12.2	2002/09/21	1.269
15127	2002/09/20	176.6	23h53m - 4°	2002/09/20 14.2 2002/09/20 13.4		1.183 0.839
239	2002/09/21 2002/09/24	179.7	0h 3m + 0°	2002/09/24 12.8		
18	2002/09/26 2002/09/28	167.3°	0h38m - 9°	2002/09/27 7.8	2002/09/28	0.814 1.616
1874	2002/09/30	175.9°	0h34m - 0°	2002/09/30 13.4		1.227
2013	2002/09/30	163.6	0h52m -12°	2002/09/28 14.1		
352	2002/10/01	173.80	0h20m + 8°	2002/10/02 11.5	2002/10/03	0.888
444 1088	2002/10/03	178.2°	0h35m + 5°	2002/10/03 10.3	2002/09/30	1.305
1088 2820	2002/10/03	177.10	0h33m + 6°	2002/10/03 14.2		
429 969	2002/10/06	172.20	0h34m +12°	0000/10/06 10 8	2002/10/06	1.304
969 984	2002/10/07	166.0°	Oh28m +18°	2002/10/06 11.8		2.2.0
5349				2002/10/20 14.2		
1734 86		172.6	1h26m + 1°	2002/10/13 11.7	2002/10/08 2002/10/14	1.149 1.496
428 817	2002/10/13	179.4	1h13m + 8°	2002/10/13 12.8	2002/10/14	0.907
4925	2002/10/13	169.8	1h43m -10° 0h56m +17°	2002/10/14 13.5 2002/10/13 14.4	2002/10/12	1.210 1.326

10					
Planet	Max Elon D Max E	RA Dec	Min Mag D Mag	Min Dist D	Min Dist
3679	2002/10/15 171.3	1h 6m +16*	2002/10/14 14.5	2002/10/11	0.724
6455	2002/10/18 165.3	1h17m - 4°	2002/10/11 13.8	2002/09/28	0.444
121	2002/10/20 171.6	1h49m + 2°	2002/10/19 11.4	2002/10/17	1.968
1275	2002/10/24 170.8	2h12m + 3°	2002/10/24 13.3	2002/10/22	1.249
10142	2002/10/25 165.0	2h13m - 2°	2002/10/25 14.4	2002/10/23	1.080
2257	2002/10/27 176.4	2h lm +16°	2002/10/27 14.5	2002/10/20	0.955
668	2002/10/28 179.8	2h10m +13°	2002/10/28 14.5	2002/10/19	1.418
2585	2002/10/29 167.5	2h28m + 1°	2002/10/28 14.1	2002/10/25	0.882
735	2002/11/01 177.3	2h29m +11°	2002/11/01 11.2	2002/10/21	0.988
5142	2002/11/03 176.4		2002/11/03 13.6	2002/11/01	0.859
2755	2002/11/06 178.5		2002/11/06 14.5	2002/11/01	1.166
109	2002/11/12 168.0		2002/11/12 10.5	2002/11/13	0.927
200	2002/11/12 168.3		2002/11/12 11.3	2002/11/11	1.400
6610	2002/11/14 177.4		2002/11/14 14.3	2002/11/13	0.909
2569	2002/11/17 172.0		2002/11/16 13.8	2002/11/13	1.246
4288	2002/11/18 175.0		2002/11/18 14.4	2002/11/12	1.286
459	2002/11/25 169.0		2002/11/25 12.7	2002/11/22	1.112
963	2002/11/25 179.4		2002/11/25 13.9	2002/11/24	0.953
1883	2002/11/26 165.7		2002/11/25 14.5	2002/11/22	0.819
1074	2002/11/28 179.2		2002/11/28 13.2	2002/11/27	1.641
1125	2002/11/30 177.6	4h25m +19°	2002/11/30 14.4	2002/12/04	1.579
1550	2002/12/01 178.8	4b28m +22°	2002/12/01 13.3	2002/11/18	0.944
240	2002/12/03 177.3		2002/12/03 11.1	2002/11/30	1.146
350	2002/12/05 167.0		2002/12/05 12.0	2002/12/05	1.665
593	2002/12/05 178.1	4h45m +20°	2002/12/05 11.5	2002/12/09	1.191
20	2002/12/10 178.8	5h 7m +21°	2002/12/10 8.4	2002/12/12	1.103
77	2002/12/10 175.9		2002/12/10 11.2	2002/12/08	1.344
1013	2002/12/10 163.4		2002/12/12 12.9	2002/12/16	1.270
725	2002/12/12 178.6		2002/12/12 13.9	2002/12/05	1.151
1052	2002/12/15 177.3		2002/12/15 13.7	2002/12/10	1.004
1187	2002/12/19 163.0		2002/12/17 13.9	2002/12/12	1.207
1116	2002/12/27 151.8	6h25m +51°	2002/12/25 12.9	2002/12/24	1.352

### CLOSE MUTUAL APPROACHES OF MINOR PLANETS IN 2002

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(Received: 25 October)

The table below lists 30 cases where one minor planet comes to within 120" of another and both are of magnitude 16 or brighter. A challenge for minor planet observers!

Here I present a list of close approaches between numbered minor planets larger than 40 km during 2002 where:

- the elongation of the Sun is more than 30°.
- both minor planets are brighter than visual magnitude 16.
- and the minimum geocentric separation is less than 120".

The table gives the following data:

- 1. Date: date and time of closest geocentric approach (in U.T.). All other information is given for this instant.
- 2. Closest approach: gives the minimum geocentric distance (in seconds of arc) and the position angle (in degrees) of the *nearest* minor planet with respect to the *farthest* one.
- 3. Minor planet 1: contains information about the *nearest* minor planet:
  - · number and name
  - visual magnitude
  - · parallax in seconds of arc
  - apparent motion in seconds of arc per hour
  - position angle of the direction of motion in degrees
- 4. Minor planet 2: information about the *farthest* minor planet. The same data as for the nearest one are given. In addition the right ascension and declination (2000.0) are printed.
- 5. Sun and Moon:
  - elongation of the Sun in degrees
  - elongation of the Moon (degrees)
  - illuminated fraction of the Moon in %

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				Close m	utual approach	es of minor planets	3			_	
								(Die	t. < 120 ; El. Sun :	0 > 30 ; magn. < 16.0	))
				_				•			•
	D a t e (U.T.)	Min. Pos. dist. ang.	Nino Name	r pla:		N 4 m 4	inor p Vis. Hor.	lanet Motion	2 Right Decli-	Flon- Ill.	
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	feb 11 21 47.5	16.52 355	253 Mathilde		1 69.44 69	1567 Alikoski		39.77 57	1 28.27 + 4 49.9	59 64 1	
?	feb 14 13 46.6	11.26 163	60 Echo	12.33 4.3	5 64.75 72	834 Burnhamia	15.93 2.33	28.46 70	2 37.30 +12 11.8	75 52 4	
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	may 20 22 17.8	107.58 250	936 Kunigunde	15.06 3.7	7 16.54 285	721 Tabora	15.32 2.83	15.26 277	13 18.95 - 6 33.7	140 37 62	
	may 21 19 17.4	116.84 328	75 Eurydike		88.39 65	167 Urda	14.99 2.56		0 58.74 + 5 46.7	44 159 71	
	jun 20 15 7.0	56.02 215	360 Carlova		1 40.96 112	893 Leopoldina		33.74 110		66 55 76	
	jun 23 10 22.6	13.61 6	230 Athamantis		2 43.92 101	539 Pamina		28.31 104	11 19.65 - 2 59.7	80 83 94	
	jun 24 14 20.0	92.40 104	1356 Byanza		6 51.78 73	260 Huberta	15.10 2.34		3 03.29 +13 07.3	45 137 97	
	jul 8 21 45.5	97.60 222	127 Johanna	13.82 2.7	7 58.28 117	758 Mancunia	14.42 2.30	45.88 113	10 37.83 +13 15.7	49 72 5	
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	sep 28 16 23.4	91.92 63	389 Industria		8 27.44 89	446 Aeternites		24.55 66	5 56.62 +28 44.0	96 24 60	
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	dec 2 5 38.9	61.60 181	660 Crescentia		8 67.20 87	190 Ismene		31.84 82	20 07.94 -16 11.5	50 81 9	
	dec 8 20 20.1	10.05 9	236 Honoria	13.55 2.8	9 68.78 83	466 Tisiphone	14.84 2.09	44.27 74	19 54.44 -15 55.3	41 18 24	
?	dec 17 8 35.7	0.35 25	116 Sirona		8 61.81 110	21 Lutetia	12.96 2.80	53.46 110	13 53.82 - 9 02.5	55 150 91	
	dec 18 20 29.9	35.66 75	977 Philippa		2 32.71 282	1669 Dagmar	15.71 3.79		4 56.28 +23 47.8	168 19 96	
	dec 27 11 35.1	68.13 129	333 Badenia		5 60.45 70	909 Ulla	15.24 2.23	47.55 80	21 25.95 -16 14.9	43 129 47	
	dec 28 2 35.6	22.01 153	308 Polyxo		7 39.04 69	34 Circe	14.07 3.09		23 56.30 - 1 50.3	82 158 40	
	dec 28 10 40.6	52.40 166	102 Miriam	13.91 3.0	8 82.62 77	190 Ismene	14.91 1.68	37.12 79	20 32.78 -15 09.7	30 103 36	

### THE MINOR PLANET OBSERVER: CAPS

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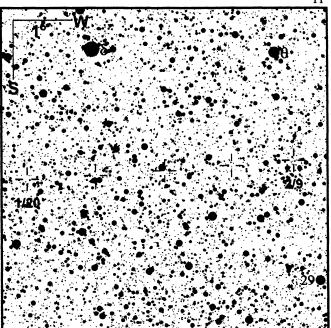
The past few issues of the Minor Planet Bulletin have shown the value of collaborating with others to get results. This is true in everyday life and is particularly valuable when working asteroid lightcurves. I've covered the reasons why a single observer cannot always get results, or at least will have a great deal of difficulty doing so: periods near 24 hours, weather, and so on. That's not to say a solitary observer can't do quite a bit of work; again these pages are testament to that. However, as more amateurs get involved with lightcurve work, there also becomes the need to help direct observing programs so that the most benefit can be derived from observations. Amateurs may have more scope time than professionals but even amateurs have their time limitations.

A couple of years ago, I started the Collaborative Asteroid Lightcurve Link (CALL), which is a web site where one can download a list of those asteroids with established lightcurves, view lists of potential targets for the upcoming months, post notifications of which asteroids are being worked, and post summaries of results obtained. The site's popularity has been growing but it lacked some resources to help direct those just getting started with lightcurve work and those who knew how to do the work but weren't exactly sure on how to proceed.

Establishing such resources and a network of amateur observers, who would work with one another and work with the professional community was one of the topics in small discussions during the Asteroids III conference. Out of those discussions came the Center for Asteroid Physical Studies (CAPS). CAPS is not a formal organization with membership dues but simply a means of trying to build that network of observers. In early September 2001, three amateurs observers, well known for their work in astrometry and photometry, and I met at my house in Colorado Springs to form the beginnings of what CAPS would be and how it would develop.

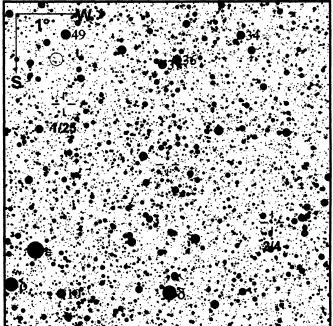
You'll find more details at the CAPS web site (http://www.MinorPlanetObserver.com/caps/default.htm – a work in progress!) but, in short, we hope to encourage amateurs to take part in directed observing programs (go for specific targets). We will also try to provide on-line resources to help the beginner to advanced observer understand the process and techniques involved in quality lightcurve work. Our first directed program is to encourage observers to determine the lightcurves for the first 1000 numbered asteroids. There are quite a few yet to be measured and as these targets tend to be brighter, they will make good starter projects for individuals or schools building a photometry program. In regards to the latter, remember that the results can be published on these pages. Hold up an MPB with a published article and tell how that lightcurve work helped the students with their math skills, and those budget purse strings may loosen just a little.

As the talks at Palermo went on, it became very clear that there was so much work that could and should be done and that amateurs could make significant contributions. Even the most prolific of observers from the MPB pages could not hope to do it alone or even with the combined efforts of all the "MPB regulars." If you're willing to give even a little time, please visit the CAPS site and join the effort to build our knowledge of the system of minor planets. Clear Skies!



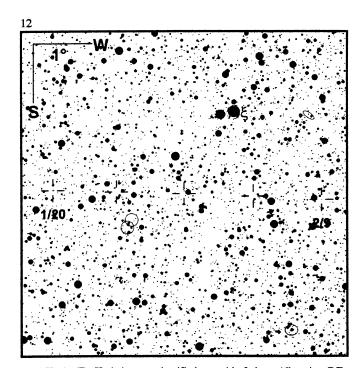
549 Jessonda (F) M. Wolf found Jessonda in 1904 November. It's named after a character in an opera by Spohr. The unclassified asteroid has a diameter of about 27km. Its path during late January take the asteroid south of Delta and Theta Cancri.

Date	RA1950	Dec1950	RA2000	Dec2000	M	PA E
01/20	8 49.21	+15 56.1	8 46.41	+16 07.2	13.2	5.1 169
01/25	8 44.38	+16 00.4	8 41.58	+16 11.3	13.0	2.3 175
01/30	8 39.50	+16 05.4	8 36.69	+16 16.0	13.0	1.4 177
02/04	8 34.76	+16 10.5	8 31.95	+16 20.9	13.2	3.9 172
02/09	8 30.37	+16 15.3	8 27.54	+16 25.4	13.4	6.7 166
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**5438 Lorre** (F) Named after Jean Lorre, a software developer for NEAT and JPL, Lorre was discovered by E.F. Helin in 1990 August. The asteroid will not be in the13s again until 2011. The field covers the Hydra-Cancer border with the three bright stars at bottom belonging to Hydra.

Date	RA1950	Dec1960	RA2000 Dec2000	M	PA E
01/20	8 51.82	+ 9 57.1	8 49.11 +10 08.4	13.8	6.7 165
01/25	8 44.91	+ 8 54.2	8 42.21 + 9 05.2	13.6	5.0 169
01/30	8 37.76	+ 7 52.7	8 35.08 + 8 03.2	13.6	4.7 170
02/04	8 30.60	+ 6 53.3	8 27.93 + 7 03.4	13.6	6.2 167
02/09	8 23.67	+ 5 56.7	8 21.01 + 6 06.4	13.7	8.5 162



1113 Katja (F) Katja is an unclassified asteroid of about 46km size. P.F. Shajn found the asteroid in 1928 August. M. Wolf found it ten days later. It is named after lab assistant and orbit calculator, Katya Iosko. Xi Cancri serves is near center. A small group of galaxies lies near the 1/25 position.

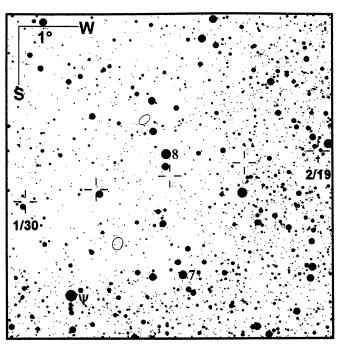
RA2000 Dec2000

RA1950 Dec1960

2/04 9 08.01 +20 31.5 9 05.16 +20 48.1 13.1 1.5 176 9 00.19 +20 43.4 13.2 3.0 171			•		
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/09 9 03.04 +20 31.5 9 00.19 +20 43.4 13.2 3.0 171	.10			<b>4</b> 13	
	/09 9 03.04				

4577 Chikako (F) Chikako is named in honor of Japanese educator Chikako Mihashi, who contributed considerable effort towards developing the "Astro Village" in Yatsugatake. The field covers a corner of Leo and includes part of Leo Minor, including 20 Leo Minoris.

Date	RA1950	Dec1950	RA2000	Dec2000	M	PA E
01/20	9 55.75	+31 44.7	9 52.83	+31 58.9	14.2	13.7 153
01/25	9 52.97	+32 29.3	9 50.03	+32 43.4	14.1	11.9 157
01/30	9 49.43	+33 09.6	9 46.47	+33 23.6	14.1	10.6 160
02/04	9 45.34	+33 43.8	9 42.36	+33 57.6	14.0	9.9 161
02/09	9 40.92	+34 10.3	9 37.92	+34 24.0	14.1	10.1 160



1152 Pawona (F) Here's a perfect testament to collaboration. Pawona is a combination of Palisa and Wolf, two famous asteroid discovers who worked together. The 21km asteroid wends its way through Leo in late January, pass close to several galaxy fields.

RA2000 Dec2000

9 44.27 +15 50.8 9 39.29 +16 04.1

5.8 166 3.2 172

RA1950 Dec1950

1/30	3 41.00			113 30.0	14.0	0.0 100
2/04	9 42.03	+15 50.4		+16 04.1	14.2	
2/09	9 36.80	+16 03.9	9 34.05	+16 17.4	14.0	
2/14	9 31.49	+16 16.6		+16 29.9	14.1	2.2 175
2/19	9 26.28	+16 27.9	9 23.51	+16 40.9	14.3	4.8 169
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1988 CA (F) There's not much to tell about this asteroid save that rarely does it rise above 15th magnitude and that this is the best (brightest) the asteroid gets from now through 2050. R. Rajamohan found the asteroid in 1988 February. The field is in Leo.

Date	RA1950	Dec1950	RA2000 Dec2000	M	PA E
02/20	11 06.39	+ 6 21.6	11 03.80 + 6 37.8	14.9	6.2 166
02/25	11 03.22	+ 7 17.3	11 00.63 + 7 33.5	14.7	3.6 172
03/02	10 59.84	+ 8 14.3	10 57.24 + 8 30.4	14.6	1.1 178
03/07	10 56.39	+ 9 10.9	10 53.77 + 9 26.9	14.7	2.2 175
02/12	10 E2 D1	110 OF 6	10 50 20 410 21 5	1/ 0	A 9 160

# ASTEROID PHOTOMETRY AT SUNFLOWER OBSERVATORY: RESULTS FOR 507 LAODICA AND 1147 STAVROPOLIS

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(Received: 16 September Revised: 23 October)

Lightcurves for two asteroids were measured in August and September 2001 at the Sunflower Observatory located in Olathe, KS. The period of 507 Laodica was found to be  $4.705 \pm 0.02h$  with an amplitude of  $0.22 \pm 0.02m$ . 1147 Stavropolis was determined to have a period of  $5.66 \pm 0.01h$  and the amplitude measured was  $0.45 \pm 0.02m$ .

Sunflower Observatory is a participant in the Center for Asteroid Physical Studies (CAPS), a goal of which is to determine accurate lightcurves on the 1000 – 1500 brightest asteroids. Target selection is made from the Collaborative Asteroid Lightcurve Link (CALL) web site maintained by Brian Warner (http://www.MinorPlanetObserver.com/astlc/default. htm). All planned targets are posted to avoid duplication of effort unless needed for collaboration and results are posted when obtained.

The current instrumentation at Sunflower Observatory is a 0.30-m Meade LX-200 SCT and Santa Barbara Instruments Group ST9E CCD camera. The effective focal ratio of the system is f7.2 yielding a pixel size of 2.3 arc sec. square. During August and September the camera is cooled to -5°C and the usual exposure is 120 sec. using the tracking feature of the SBIG camera. The objective is to reach a S/N ratio ≥50 for the target. Exposures are generally started at 300 second intervals while the target is ≥30 degrees elevation. All images are unfiltered. All images are dark subtracted and flat fielded before measurement. Usually four comparison stars are chosen in the same field and selected on the basis of similar magnitude and spectra to the target.

Canopus software is used to measure the images and determine best fit of a lightcurve to the data. More information on this software may be obtained at <a href="http://www.MinorPlanet">http://www.MinorPlanet</a> Observer.com.

### 507 Laodica

Discovered 1903 February 19 by R. S. Dugan at Heidelberg. Named for a daughter of Priam and Hecuba. She fell in love with Acamas, son of Theseus and Phaedra, when he and Diomedes came from the Greeks to demand the restoration of Hellen. The minor planet was not named by the discoverer. It had previously been observed by A. Charlois in Nice, but later lost. The name appeared when an ephemeris was published.

A total of 152 observations were used from data acquired on August 12, 17, 18, 19, 20, and 29 covering the entire lightcurve with no gaps. The period of the lightcurve was determined to be  $4.705 \pm 0.005h$ . The curve is somewhat asymmetrical with a maximum amplitude of  $0.22 \pm 0.02m$ .

### 1147 Stavropolis

1929 LF was discovered 1929 June 11 by G. N. Neujmin at Simeis. Numbered 1147 and named by the discoverer, it honors the city of Stavropol' in northern Caucasus region. Between 1936 and 1946 the city was named Woroschilowsk.

A total of 150 observations were used from data acquired on September 2, 3, and 5 covering the entire lightcurve with no gaps. The period of the lightcurve was determined to be  $5.66 \pm 0.01h$ . The curve is asymmetrical with a maximum amplitude of  $0.45 \pm 0.02m$ .

#### References

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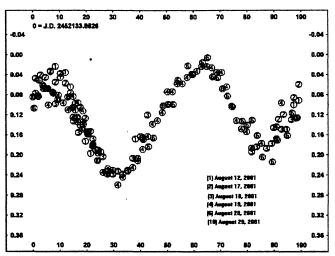


Figure 1. The lightcurve for 507 Laodica. The period is  $4.705 \pm 0.005h$  with an amplitude of  $0.22 \pm 0.02m$ .

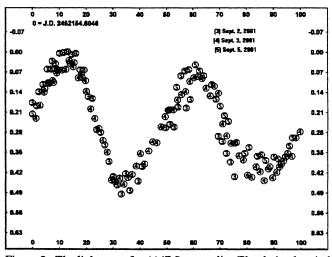


Figure 2. The lightcurve for 1147 Stavropolis. The derived period is  $5.66 \pm 0.01h$ ; the amplitude is  $0.45 \pm 0.02m$ .

# ASTEROID PHOTOMETRY AT THE PALMER DIVIDE OBSERVATORY: RESULTS FOR 573 RECHA, 1329 ELIANE, AND 8041 MASUMOTO

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(Received: 26 August)

Lightcurves for three asteroids were measured in 2001 January, April, and August at the Palmer Divide Observatory located near Colorado Springs, CO. The period of 573 Recha was found to be  $6.53 \pm 0.05h$  with an amplitude of  $0.25 \pm 0.05m$ . 1329 Eliane was determined to have a period of  $8.0 \pm 0.1h$  with an amplitude of  $0.08 \pm 0.02m$ . For 8041 Masumoto, the period was found to be  $14.10 \pm 0.02h$  and the amplitude  $0.30 \pm 0.02m$ .

The goal of the initial Palmer Divide Observatory asteroid photometry program is to measure the lightcurves for as many asteroids as possible. A successful measurement is considered when the period has been determined to within 0.05h or better and 0.02m or better for the amplitude. The current instrumentation is a 0.5m Ritchey-Chretien working at f/8.1 The camera is a Finger Lakes Instrumentation camera using the Kodak KAF-1001E non-antiblooming CCD chip working at -30°C and 2x2 binning (0.48um pixels). This yields a scale of about 2.4 arcseconds per pixel. Given that the average seeing at the Palmer Divide Observatory is on the order of 4-5 arcseconds, this is usually a satisfactory match. All observations for this paper were unfiltered with unguided exposures of either 60s or 70s. All images were dark subtracted and flat-fielded before measuring.

Initial targets are determined by referring to the list of lightcurves maintained by Dr. Alan Harris (Harris 2001) as well as those posted on the Collaborative Asteroid Lightcurve Link (CALL) web site maintained by the author (http://www.MinorPlanetObserver.com/astlc/default.htm). At least two nights are dedicated to the initial run for every target. Depending on the preliminary analysis of the lightcurve data from those two nights, additional runs are allocated as necessary to assure full coverage of the lightcurve with no significant gaps, if possible. For asteroids with periods approaching and exceeding 24 hours or with aliases close to 24 hours, this becomes difficult.

Custom software written by the author, MPO Canopus, is used to measure the images since it allows automatic storage of the measured magnitudes of the comparison stars and targets. It uses aperture photometry with magnitudes determined by calibrating images against field or, preferably, standard stars. The package includes a Fourier analysis routine, the original FORTRAN code for which was supplied by Alan Harris (Harris et al, 1989) and converted to Delphi Pascal. If the data from a single night appears to cover at least half a period or more, then an estimate based on a plot of the raw data is used to help narrow the possibilities when using data from two or more nights.

For two of the measured asteroids, 573 Recha and 1329 Eliane, the old addage "If at first you don't succeed, try, try again" comes to mind. The data for these targets was gathered in early 2001, initially measured, and no satisfactory results determined. During a cleanup of the computer's hard drive to archive images and data

reduction files to CD-ROM, I was determined to give these two and a couple of other unresolved asteroids another try. Eventually, all the images for all the asteroids were remeasured, taking care to avoid measuring images where field stars may have affected the data. This effort paid off in the case of Recha and Eliane as the new data allowed reasonable solutions to be determined. Part of the success can also be attributed to refinements in the code for determining the centroid and total flux made after discussing the topic with a number of professional and amateur photometrists.

### 573 Recha

This 39km unclassified asteroid was discovered by Max Wolf in 1905 September. It is named after a character in Lessing's Nathan the Wise. The semi-major axis of 3.017AU places the asteroid towards the outer main belt. The inclination is about 9° and the eccentricity 0.109. Observations were obtained on the nights of 2001 January 1 (my way of celebrating the 200th anniversary of Ceres' discovery) and January 4. Data were also obtained on January 5 but suffered from excessive noise due to less than ideal conditions and so the data were not used in the final solution, which included about 135 observations from the nights of January 1 and 4. In the original measurements, the data from all three nights showed considerable noise and no solution could be found. The remeasuring "quieted" the data for two of the nights and so lead to a least a preliminary and apparently reasonable solution.

The observations are plotted in Figure 1, which is phased against the period of  $6.53 \pm 0.05h$ . The amplitude of the curve is  $0.25 \pm 0.05$ . While the extrema of the plots matched well, there was a decided difference in the amplitude of the curve between the two nights, with the data from January 4 showing a smaller amplitude by about 0.06-0.08m.

### 1329 Eliane

E. Delporte found Eliane in 1933 March. It was subsequently named after the daughter of astronomer Paul Bourgeois. The asteroid is of type SU and about 26km in size. The semi-major axis of the asteroid's orbit is 2.615AU, the inclination 14.5°, and the eccentricity is 0.175.

Data from the nights of 2001 April 27 and 28 were used to find a period of 8.0  $\pm$ 0.1h. The amplitude of the curve is 0.08  $\pm$ 0.02m. Figure 2 shows the data phased to this period. There is a considerable gap in the data. However, prior to remeasuring the original images, no solution could be found at all. In addition, based on a typical double maximum/minimum curve, the data from April 27 agree fairly well with the eight-hour period. Assuming this is the correct value, the with the period having an integral alias close to twenty-four hours, any single observer would have to spend a considerable amount of time and effort to obtain sufficient data to cover the entire curve. This asteroid presents an excellent opportunity for collaborative efforts at future apparitions. The next is in 2002 September when the asteroid will be at -16° when at its brightest of about 14.1. Unfortunately for Northern Hemisphere observers, Eliane is generally at its brightest when at or south of the celestial equator

### 8041 Masumoto

There was no lack of data for this asteroid, with observations being made on the nights of 2001 August 12, 17, 18, and 19. Figure 3 shows a phased plot on the derived period of  $14.10 \pm 0.02h$ . The amplitude of the curve is  $0.30 \pm 0.03m$ . The curve shows a decided asymmetry between the first and second minima, with the latter having a markedly smaller amplitude and different shape.

8041 Masumoto was discovered by F. Uto at Kashihara in 1993 November. It is named in honor of the man who built Uto's observatory, Takeji Masumoto. Masumoto has a semi-major axis of 2.625AU, an orbital inclination of 11.4°, and eccentricity of 0.186.

### Acknowledgements

Thanks go to Dr. Alan Harris of the Jet Propulsion Laboratory for making available the source code to his Fourier Analysis program and his continuing support and advice.

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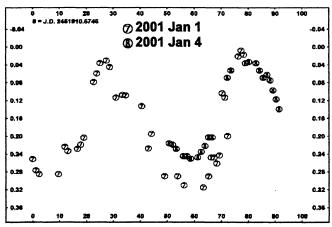


Fig. 1 – The lightcurve plot for 573 Recha. The period is  $6.53 \pm 0.05$ h with an amplitude of  $0.25 \pm 0.05$ m.

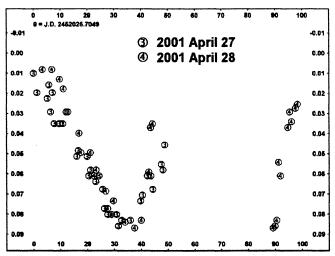


Fig. 2 – The lightcurve plot for 1329 Eliane. The derived period is  $8.0 \pm 0.1h$ ; the amplitude is  $0.08 \pm 0.02m$ .

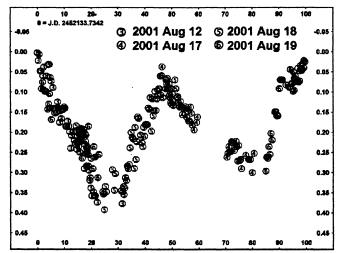


Fig. 3 – The lightcurve plot for 8041 Masumoto. The period is  $14.10\pm0.02h$  with an amplitude of  $0.3\pm0.03m$ .

### **ASTEROID-DEEPSKY APPULSES IN 2002**

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A list of favorable appulses between asteroids and brighter deepsky objects during the year 2002 is presented.

The following list is a subset of the results of a search for asteroid-deepsky appulses meeting the following criteria:

- 1. The asteroid was brighter than 14.0.
- 2. The separation between the two was less than 360 arcseconds.
- The phase of the moon was between mid-waning gibbous to mid-waxing gibbous
- 4. The event was at least 45° from the Sun.

The list below is not comprehensive by any means. However, it's a good first check. For a more complete check, the Minor Planet Center's web site at

http://cfaps8.harvard.edu/~cgi/CheckSN.COM

allows you to enter the location of a suspected asteroid or supernova and check if there are any known targets in the area.

The complete set from which the table below is a small subset can be found at the MPO web site:

http://www.MinorPlanetObserver.com/htms/dso.htm

The table gives the following data:

Date/Time	Universal Date and Time of closest approach
#/Asteroid	The number and name of the asteroid
RA/Dec	The J2000 position of the asteroid
Mag	The approximate visual magnitude of the asteroid
Sep/PA	The separation in arcseconds and the position angle from the DSO to the asteroid
DSO	The DSO name or catalog designation
Mag	The approximate total magnitude of the DSO
Туре	The type of DSO: OC = Open Cluster; GC = Globular Cluster; G = Galaxy
SE/ME	The elongation in degrees from the sun and moon respectively
MP	The phase of the moon: $0 = \text{New}$ , $1.0 = \text{Full}$ .

Positive = waxing; Negative = waning

Date	UT		Asteroid		RA	Dec	M	Sep	PA	DBO	Type	Mag	82	ME	ЖР
01/05	00:04	498	Tokio	1	22.26	- 0 50.4	13.4	198	321	MGC 497	G	13.0	94	160	-0.630
01/06	19:36	665	Sabine	12	06.19	-14 29.3	13.7	301	59	MGC 4094	G	11.8	99	25	-0.425
01/10	00:16	71	Wiobe	14	07.97	-33 18.5	11.9	36	225	MGC 5488	G	11.9	69	33	-0.126
01/11	23:17		Palisana	6	37.74	+ 4 40.9	13.1	187	169	Cr 107	oc	5.1	158	160	-0.025
01/11	19:45	80	Sappho	8	57.57	+ 2 59.6	11.7	326	185	MGC 2716	G	11.8	152	134	-0.030
01/12	02:01		Sappho	8	57.33	+ 2 59.9	11.7	297	5	MGC 2713	G	11.8	152	137	-0.021
01/14	01:43	411	<b>Eanthe</b>	9	36.99	+23 32.0	13.8	251	226	MGC 2927	G	12.9	153	160	0.003
01/14	05:53	5438	Lorre	8	59.19	+11 10.2	13.9	12	326	MGC 2720	G	12.8	159	165	0.005
01/17	22:07	291	Alice	6	10.86	+20 32.0	13.9	240	189	MGC 21758	oc	6.8	155	106	0.169
02/07	13:31	241	Germania	0	04.54	+ 5 13.3	13.3	145	339	MGC 7820	G	12.9	45	98	-0.202
02/08	19:52	247	Eukrate	1	38.33	+35 17.6	12.4	336	160	MGC 634	G	13.0	77	114	-0.111
02/08	00:20	145	Adeona	18	02.27	-23 06.9	14.0	293	186	ж20	CMB	6.3	49	1	-0.167
02/10	04:48	423	Diotima	18	03.06	-24 28.3	13.1	322	186	ж8	CMB	5.0	51	27	-0.041
02/16	14:48	74	Galatea	1	21.40	+ 7 01.0	13.5	2	144	MGC 485	G	13.0	54	8	0.159
02/18	05:26	247	Eukrate	1	57.80	+36 24.6	12.5	283	341	MGC 759	G	12.7	73	28	0.284
02/18	23:20	774	Armor	18	07.18	-23 15.1	14.0	297	356	MGC 6546	OC	8.0	59	131	0.351
02/19	04:25	85	Io	12	47.08	-11 33.1	12.6	285	35	MGC 4680	G	13.0	135	149	0.370
02/19	15:36	100	Hekate	9	05.49	+18 22.2	13.0	207	23	MGC 2749	G	11.8	163	83	0.415
02/21	05:05	326	Tamara	18	59.59	-36 38.5	13.7	32	190	MGC 6723	GC	7.3	52	145	0.572
02/21	12:04	28	Bellona	10	23.74	+12 33.8	9.9	62	40	MGC 3230	G	12.8	178	79	0.602
03/04	05:23	186	Celuta	17	32.22	-32 39.3	13.7	206	201	FGC 6374	oc	9.0	80	36	-0.693
03/04	00:45	61	Danae	9	38.60	+17 04.9	12.8	171	359	MGC 2943	G	12.4	158	86	-0.713
03/05	13:40	704	Interamnia	2	03.65	+23 59.9	11.7	249	170	UGC 1551	G	12.5	53	146	-0.550
03/08	16:49	186	Celuta	17	39.74	-33 09.8	13.6	331	21	MGC 6404	oc	10.6	82	24	-0.243
03/09	06:12	667	Denise	6	30.45	+ 5 01.5	13.6	162	123	MGC 2237	CMB	5.5	109	153	-0.197
03/12	12:29	584	Semiramis	13	07.09	-23 43.2	13.2	131	183	MGC 4968	G	12.8	144	127	-0.023
03/12	12:22	192	Mausikaa	11	20.24	+ 2 55.0	11.0	190	196	MGC 3630	G	11.9	178	164	-0.023
03/12	09:07	123	Brunhild	3	44.86	+24 11.4	13.9	267	353	<b>∀dB</b> 20	BH	11.6	68	87	-0.027
03/12	04:33	454	Mathesis	10	49.35	+16 09.6	12.4	207	191	MGC 3399	G	12.8	164	173	-0.033
03/13	22:44	1011	Laodamia	11	45.88	+10 51.3	13.6	102	42	MGC 3869	G	12.8	171	176	0.002
03/20	19:59		Celuta		59.11	-34 32.4	13.4	231	24	MGC 6453	<b>e</b> c	9.9	90	161	0.369
04/05	03:40		Lumen		07.15	-23 38.2	13.0	173	14	NGC 4968		12.8	162	86	-0.447
04/08	05:31		Ursula		47.59	-30 53.3	12.3	162	357	NGC 5292	G	11.9	154	109	-0.178
05/02	08:40		Themis		05.00	+24 02.1	12.7	125	1	IC 2157	oc	8.4	49	162	-0.688
05/06	16:11	174	Phaedra	16	54.02	-41 48.4	12.2	28	147	MGC 6231	OC	2.6	144	84	-0.279
05/07	19:35	31	Euphrosyne	12	26.79	+ 7 56.2	11.8	15	324	MGC 4416	G	12.4	135	171	-0.188
05/08	22:22	639	Latona	0	03.79	+ 7 27.4	13.8	26	333	MGC 7816	G	12.8	45	12	-0.113
05/10	00:03	6455	1992 HE	11	10.04	-23 47.3	13.8	228	147	IC 2627	G	12.0	125	142	-0.056
05/13	07:02	88	Thisbe	7	09.38	+20 36.8	13.2	130	188	MGC 2342	G	12.6	54	44	0.007
05/14	01:09	387	Aquitania		53.77	+19 23.7	13.8	346	355	UGC 3587	G	12.8	50	31	0.026
05/14	02:47		Oceana		40.48	+14 54.4	13.8	42	208	MGC 2954	Ğ	12.4	89	70	0.028
05/16	11:35	1018	Arnolda	15	39.88	-30 37.9	14.0	236	184	MGC 5968	Ğ	12.2	168	136	0.164
05/18	19:42	326	Tamara	22	57.21	-36 26.6	13.2	94	176	IC 1459	Ğ	10.0	87	153	0.388
			_								-		••		

Date	UT		Asteroid		RA		Dec	AM	Sep	PA	DSO	Type	Mag	SE	ME	ЖР
06/01	15:36	477	Italia	17	40.17		2 17.4	13.0	270	167	ж6	OC .	4.2	163	60	-0.627
06/01	18:21		Tina		45.01		9 25.6	13.3	190	241	N26	OC.	8.0	149	72	-0.426
06/03	07:13		Portlandia		49.84		9 21.0	13.8	240	8	MGC 6000	G	12.2	166	103	-0.470
06/06	22:45	477	Italia	17	34.75	-3	2 28.6	12.8	325	351	NGC 6383	oc	5.5	168	123	-0.157
06/07	18:18	8	Flora	10	44.41	+1	4 05.4	11.3	25	26	MGC 3357	G	12.6	81	118	-0.103
06/09	14:14	1349	Bechuana	17	37.99	-3	5 05.5	13.6	343	197	MGC 6396	OC.	8.5	166	153	-0.020
06/09	05:15		Italia		32.27	_	2 32.2	12.8	228	353	MGC 6374	OC	9.0	169	151	-0.032
06/10	10:04		Klytia		02.57		6 20.3	12.6	262	185	Ж19	GC	7.2	176	174	-0.002
06/11	11:44		Italia		29.73		2 35.0	12.7	302	175	Antalova 2	oc	8.8	170	171	0.003
06/13	04:28		Aegle		32.30	-	3 14.7	12.4	300	97	MGC 5193	G	11.6	128	105	0.056
06/15	20:49		Astraea		58.77		8 16.6	12.6	214	163	MGC 766	G G	12.7 12.9	54 105	116 31	0.263 0.359
06/16	18:44		Genua		44.48		0 27.2 1 15.7	13.4	107 52	191 233	UGC 7911 MGC 4103	oc	7.4	113	70	0.359
06/17 06/17	13:44 22:28		1992 FE Althaea		04.57	-	5 11.9	13.3	52 57	336	MGC 7820	a	12.9	83	168	0.490
06/17	05:53		Aurora		22.96		8 58.4	13.5	311	153	MGC 502	G	12.8	65	170	0.640
00/19			waiora					23.5								
07/01	07:59	640	Brambilla	18	39.75	_	8 32.6	13.1	217	191	Tr 34	oc	8.6	165	76	-0.629
07/05	05:33	128	Nemesis	12	27.06	+	2 31.3	13.3	92	33	MGC 4420	G	12.1	82	145	-0.267
07/08	22:18	358	Apollonia	18	18.40	-1	8 24.3	13.7	41	355	MGC 6603	oc	11.1	167	171	-0.026
07/15	22:06	96	Aegle	13	47.59	-3	0 55.2	12.9	48	346	MGC 5292	G	11.9	102	38	0.355
07/17	14:09	1032	Pafuri	18	55.16	-3	0 31.4	13.8	150	160	<b>X54</b>	GC	7.7	165	73	0.545
07/30	02:33	192	Yausikaa	11	48.96	-	2 00.0	13.2	130	25	UGC 6780	G	13.0	52	165	-0.713
08/03	06:42	415	Palatia	5	10.70	 +1	6 31.6	13.7	21	174	MGC 1807	oc	7.0	53	17	-0.322
08/07	22:39		Thyra		13.35		9 27.7	13.2	206	14	MGC 5022	G	12.9	70	82	-0.010
08/07	16:31		Photographica				7 11.1	13.7	223	145	M18	OC	6.9	140	153	-0.017
08/08	11:13	63	Ausonia	12	49.46	-1	0 04.7	12.2	147	23	AN 3	G	11.6	60	65	-0.002
08/10	13:20	201	Penelope	1	16.01	+	5 12.2	12.1	75	7	MGC 455	G	12.6	118	142	0.046
08/10	17:16	63	Ausonia	12	53.18	-1	0 27.3	12.3	177	23	MGC 4760	G	11.4	59	33	0.054
08/30	18:52	631	Philippina	0	21.06	+2	2 26.1	13.5	330	338	MGC 80	G	12.1	139	52	-0.531
08/30	05:47	631	Philippina	0	21.33	+2	2 27.5	13.5	163	338	INGC 83	G	12.5	139	47	-0.584
09/14	15:13	106	Dione	16	17.06		2 54.5	13.7	278	12	ж80	ec	7.2	75	26	0.595
09/15	22:56	753	Tiflis	17	36.20	-3	3 34.5	13.9	329	180	Tr 27	oc	6.7	92	26	0.724
09/29	22:44	480	Hansa		56.66	+	7 28.1	13.6	63	211	MGC 2485	G	12.2	68	26	-0.475
10/01	02:45		Julia	9	40.49		4 54.7	12.3	22	206	MGC 2954	G	12.4	45	28	-0.353
10/03	23:38		Engenia		39.17		6 35.1	12.0	217	327	MGC 1026	, G	12.6	150	116	-0.092
10/06	07 : 57		Tamara		57.80		6 01.3	12.7	81	64	IC 5269	G	12.2	132	136	-0.001
10/07	00:24		Venetia		52.86		2 24.2	13.6	194	190	MGC 6469	oc	8.2	75	67	0.006
10/08	15:34	209			34.02		0 57.1		173	175	MGC 976	G	12.4	152	172	0.071
10/11	11:02		Kacivelia		28.16		1 52.5		254	146	MGC 124	G	13.0	167	99	0.315
10/12	12:30		Ceres		56.53		9 58.8	7.6	108 157	166	MGC 309	e oc	11.9	162 81	88 33	0.427 0.694
10/15 10/15	04:42 12:08	5	Hebe Hispania		53.53 17.29		0 20.4	10.1 13.1	157 42	11 353	Cr 394 Bockum 13	OC	6.3 7.2	60 81	33 57	0.722
10/15	18:34		Fides		06.72		4 19.6		215	200	MGC 3121	a a	12.6	69	3	-0.332
11/06	10:48		Maria		20.34		1 19.2		110	162	MGC 1282	ø	12.9	154	157	0.037
11/07	09:00		Athene		42.42		4 47.5		107	321	MGC 1050	G	12.6	161	146	0.089
11/08	16:44		Panopaea		24.04		9 34.1		121	206	Berk 19	oc	11.4	144	164	0.191
11/25	10:31		Bathilde		26.33		7 53.0		233	221	MGC 2882	6	12.6	101	20	-0.732
11/26	13:23 	404	Arsinoe		54.73	+1	7 42.8	13.8	351 	4	MGC 3457		12.6	<b>8</b> 6	19	-0.619
12/04	13:49		Chryseis		23.60		9 18.7	13.8	21	350	MGC 6903	6	11.9	51	48	0.001
12/07	05:08		Thomis		55.56		1 10.2		316	203	UGC 6903	æ	12.3	76	114	0.106
12/09	06:59		Tergeste		10.76		4 45.9		277	231	UA 205	G	11.5	100	154	0.275
12/14	01:10	79	Enlances	13	39.43	-1	1 31.9	13.5	186	200	MGC 5254	G	13.0	55	169	0.719

### ASTEROID PHOTOMETRY OPPORTUNITIES JANUARY-MARCH 2002

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Winter season with its long nights and the opposition point high in the sky for northern hemisphere observers provides more opportunities for collaborative work among observers at far different longitudes (across oceans), that is particularly suitable for working longer period asteroids. We encourage observers from different continents to take this opportunity and establish a collaboration on such objects. That will help to suppress a bias against longer periods in the available sample.

In the Table below, we present a list of suitable photometric targets for the January-March 2002 period. Most of the objects have been selected from a more extensive list prepared by Brian Warner. We selected objects with the predicted V<14 in opposition and unknown or not reliably established periods. Some of the targets have listed previous less reliable period estimates that are generally relatively long; that is a selection effect – long periods are usually more difficult to be established reliably. We have comments on some of them: The period of 654 Zelinda was derived by Schober (1975) and it is worth checking, it may be wrong. The amplitude of 0.03 mag for 151 Abundantia has been observed by Shevchenko et al. (1992); the asteroid may display different amplitude at the present apparition. The period of 111 Ate may be a half of the listed value (Harris and Young 1983). The lower limits on the

period and amplitude of 488 Kreusa was derived by Harris and Young (1989); the period may be actually very long - it will be interesting to know if it is not the largest (D~150 km) member of the unexplained population of very slow rotators with periods longer than a few days. The period estimates for 259 Aletheia and 823 Sisigambis have reliability codes of 1 (Weidenschilling et al. 1990, Barucci et al. 1994), Aletheia being probably the largest (D~180 km) of the targets listed below. The lower limit on the period of 514 Armida was derived by Lagerkvist (1978) from photographic observations; it may be an easy target for today's CCD observers. 58 Concordia is the lowest number asteroid with reliability code 1 (its lower limit period estimate was derived by Gil Hutton 1993). Observers interested in asteroids fainter than V=14 are encouraged to check the full list on the Brian Warner's website (http://www.MinorPlanetObserver.com/astlc/ default.htm). Note: Full references to the period and amplitude estimates cited in this article can be found in the list prepared by Harris which is available at http://cfawww.harvard.edu/iau/lists/LightcurveDat.html .

		Opp':	n	Opp'r	<u> </u>	
Aste	eroid	Date		V	Per	Ampl
		2002			[h]	
1176	Lucidor	Jan	07	13.7		
1107	Lictoria	Jan	07	12.6		
1040	Klumpkea	Jan	80	13.6		
654	Zelinda	Jan	10	9.7	31.9	0.3
1547	Nele	Jan	15	12.5		
549	Jessonda	Jan .	28	12.9		
5438	Lorre	Jan :	28	13.6		
1113	Katja	Feb	03	13.1		
412	Elisabetha	Feb	80	12.7	15	0.1
151	Abundantia	Feb	09	12.6		0.03
111	Ate	Feb	10	10.7	22.2	0.1
384	Burdigala	Feb	13	12.7		
488	Kreusa	Feb	19	11.5	>28	>0.2
259	Aletheia	Mar	11	12.2	15	0.19
754	Malabar	Mar	11	13.1		
823	Sisigambis	Mar	15	13.4	>12	>0.2
1011	Laodamia	Mar	16	13.6		
514	Armida	Mar	17	13.4	>20	>0.3
58	Concordia	Mar	22	12.1	>16	>0.07
976	Benjamina	Mar	24	13.3		
399	Persephone	Mar	30	12.9		

20

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The deadline for the next issue (29-2) is January 15, 2002. The deadline for issue 29-3 is April 15, 2002.